BODDINGTON

# Petrography, mineralogy and trace element chemistry of $\mathbf{C u}$-Au-Mo mineralisation from Central Diorite, Boddington, W.A. 

Ryan J. Guerin

Supervisor<br>Cristiana L. Ciobanu<br>Co-supervisor<br>Nigel J. Cook

Centre for Tectonics, Resources and Exploration
School of Earth and Environmental Sciences
The University of Adelaide, South Australia
ryan.guerin@student.adelaide.edu.au


#### Abstract

The world-class Boddington $\mathrm{Cu}-\mathrm{Au}-\mathrm{Mo}$ deposit has a complex genetic history. The relative importance of different ore-forming processes during the period $3.0-2.6 \mathrm{Ga}$ is debated, particularly with respect to the role played by the $2611 \pm 3 \mathrm{Ma}$ Wourahming granite. LA-ICPMS analysis of trace element concentration in molybdenite represents a valuable new metallogenetic tool to track mineralising events in deposits with protracted geologic histories. The Re content and trace-element signatures in molybdenite from diorite and granite show three distinct populations, attributed to porphyry-style (hundreds of ppm) orogenic- and granite-related systems ( $<1$ to a few ppm, respectively). Rhenium concentrations in molybdenite are highly variable on the deposit-scale. Economic concentrations of Re occur only at shallower levels in both Central Diorite and ABreccia. The Au content correlates with high-concentrations of chalcophile elements (CE). This is seen in the association of Auminerals and $\mathrm{Bi}-(\mathrm{Pb})$-tellurides present as inclusions in the molybdenite from diorite and is inferred from an LA-ICP-MS element map for molybdenite in granite.

The FIB-SEM and TEM study show that visible telluride inclusions extend down to the nanoscale as coherent intergrowths with host molybdenite. Nanoporosity is accompanied by a whole range of structural defects and twinning. The telluride species identified include unnamed $\mathrm{Bi}_{4} \mathrm{~Pb}_{7} \mathrm{Te}_{4} \mathrm{~S}_{9}$. Analysis of stacking sequences show co-precipitation of Bi-tellurides and molybdenite under equilibrium conditions. In corroboration with EPMA data, this is the first confirmation that minerals from the aleksite series are characteristic components of the ore at Boddington. Molybdenite with high-concentration of chalcophile elements is present as the 2 H polytype only, contrary to previous hypothesis that incorporation of trace elements is assisted by 3R structural modification. Instead, a new mechanism is presented in which coherent lattice-scale intergrowths between molybdenite and tellurides are reasons for the measured high CE concentrations. Knowing that $\mathrm{Bi}-(\mathrm{Pb})$-tellurides are Au -carriers, this may also explain the observed, unusual Au-enrichment in molybdenite from Boddington. Nucleation of Au fine particles is inferred from element map correlations but further work is necessary to prove if Au nanoparticles are also present.

Petrographic, mineralogical and geochemical evidence support a three-stage model for Boddington. An early porphyry event can account for the bulk of the Cu mineralisation, as well as some of the Au and Mo . A subsequent orogenic-Au event led to shearing and remobilisation of ore components. New constraints on metamorphic conditions are offered by


chlorite and stannite-sphalerite geothermometry $\left(200-420{ }^{\circ} \mathrm{C}\right)$ and the occurrence of two coexisting pyrrhotite species. The granite introduced some Au, Mo and other 'granitic' elements, notably Bi leading to substantial upgrading of Au grades by Bi -melt scavenging. The study concludes however that hydrothermal activity associated with granite was not the most important concentrator of ore minerals.

## INTRODUCTION

The Boddington $\mathrm{Cu}-\mathrm{Au}$ mine is located 100 km SSE of Perth, W.A., and is situated within the Archean Saddleback Greenstone Belt, a segment of the Yilgarn Craton. The deposit was originally discovered in 1980. Newmont Asia-Pacific developed the existing mine on a much larger scale than the original mine; starting production in 2009 (Hayes 2009).

A genetic model for the Boddington deposit which explains all observations has been difficult to derive since the deposit shows a number of features which are atypical of many other deposits in the Yilgarn Craton. As a consequence, a number of different models have been put forward (e.g. Roth \& Anderson 1993, Allibone et al. 1998, McCuaig et al. 2001, Stein et al. 2001, Hagemann et al. 2007).

A previous University of Adelaide honours project (Kalleske 2010) made two important new findings: firstly that Re concentrations, and to some extent also Au concentrations within molybdenite are commonly extremely high; and secondly that Bi-tellurides are widespread throughout the deposit, generally associated with elevated Au grades. The geochemistry of molybdenite showed promise as a new method to constrain ore genesis, in particular the role of the late Wourahming granite as a source of ore metals or heat source driving remobilisation.

Using a variety of microanalytical methods, the present project aims to significantly enlarge the geochemical dataset for molybdenite, focussing in particular on high-grade areas of the Central Diorite orebody and the Wourahming monzogranite. The primary goal is to identify whether the geochemical trends reported by Kalleske (2010) are consistent within the larger dataset and, if so, to understand the underlying controls on molybdenite geochemistry and its apparent inhomogeneity in the context of multistage ore genesis.

Secondary goals include an attempt to better constrain the types of alteration present at Boddington, an assessment of metamorphic conditions (using chlorite and sphalerite-stannite geothermometry) and an evaluation of the petrogenetic information within a number of trace
minerals in the deposit. The large number and varied compositional character of different trace minerals in the Boddington deposit carries inferences for any genetic model and any attempt to constrain conditions of ore formation or superimposed modification.

## GEOLOGICAL SETTING OF THE BODDINGTON DEPOSIT

A large portion of Western Australia is composed of Archean metavolcanic and metasedimentary rocks initially formed at 3050-2600 Ma (Myers 1993, Duuring et al. 2007). These rocks were originally part of volcanic arcs, back arc basins and microcontinents that amalgamated during a time of tectonic movement and formed the Yilgarn Craton (Myers 1993). Crustal fragments were aligned and became joined during a period of increased tectonic activity dated at 2780-2630 Ma (Myers 1993). A significant section of the Yilgarn Craton is comprised of granitoids and associated greenstone belts that underwent metamorphism. All greenstone belts in W.A. contain packages of igneous and sedimentary rocks that display similar lithologies and broad-scale structural controls. Greenstone belts are widely associated with metallic mineral deposits, since the latter form in volcanic arc and back arc basin environments where faulting can occur (Barley et al. 1990).

Boddington is a $\mathrm{Cu}-\mathrm{Au}$ deposit with associated $\mathrm{Mo}, \mathrm{W}$ and Bi . The deposit is located in the northern part of the 50 km -long, NNW-striking Saddleback Greenstone Belt (SGB). This belt forms a part of the southwestern Yilgarn Craton (Allibone et al. 1998). The trace element chemistry of diorite intrusions ( $2715-2690 \mathrm{Ma}$ ) from the SGB implies formation in an island arc setting (Allibone et al. 1998, McCuaig et al. 2001). Initial basaltic and dacitic volcanism occurred at 2714-2696 Ma. Magmatic activity was initiated again at 2675 Ma as a series of dikes. The greenstone belt is enclosed by the Western Gneiss Terrane which hosts the Wourahming monzogranite dated at $2611 \pm 3 \mathrm{Ma}$ (Allibone et al. 1998).

The Boddington deposit strikes NW-SE over a length of 4 km , with a width of 1 km (Le May 2010). Mineralisation is contained within felsic to intermediate volcanic and sedimentary rocks. Mineralisation occurs within shear zones, brittle ductile faults, veins and reactivated veins, veinlets, lenses and disseminated ores. The deposit is currently mined via two open pits (North and South). The North Pit is further divided into ABreccia, Blob, Son of Blob and North Diorite and the South Pit is divided into Blackbutt, Central Diorite, Pipeline and Southern Diorite Deep (Figure 1). The Southern Diorite Deep zone has not yet been mined as it is below the depths of the pit. The Central Diorite is a high-grade zone that is
currently being mined (at time of fieldwork) at the U108 level. At the end of 2010, in addition to $728,000 \mathrm{Oz} \mathrm{Au}$, the mine had produced $\sim 27,000 \mathrm{t} \mathrm{Cu}$. Copper reserves are estimated at 1.2 Mt .

The Boddington greenstone sequence is divided into three formations. The formations hosting the Boddington deposit are the 2714-2696 Ma Wells and Marradong Formations; the Hotham Fm. is structurally below these two formations (Wilde et al. 1986). The Wells Fm. has a thickness of 2-5.5 km and comprises felsic volcanic rocks alternating with lavas and breccias that recrystallized during metamorphism at $\sim 2640$ Ma. The Marradong Fm. (on top of the Wells Fm.) consists of metamorphosed actinolite- and albite-epidote-dominant basalts (Wilde et al. 1986). Deposit mineralogy shows that a dominant lower amphibolite facies metamorphism was followed by retrograde metamorphism (greenschist-amphibolite facies boundary conditions), especially around lithological contacts (Wilde et al. 1986). Brittleductile faults were widely developed during this metamorphic event (Allibone et al. 1998).

The series of deformational events at Boddington is hotly debated. Allibone et al. (1998) propose seven distinct post-mineralisation deformational events. In this scheme, molybdenite and chalcopyrite veins formed between 2714-2696 Ma before the first deformational event. Widespread metamorphism led to a quartz $\pm$ oligoclase $\pm$ albite-clinozoisite-muscovitebiotite $\pm$ chlorite mineralogy and well-defined rock fabric. Sericite-quartz alteration then took place within shear zones after peak metamorphism. $\mathrm{D}_{3}$ shears dated at post-2675 Ma hold quartz-albite-epidote alteration and cut through the previous two sets. Dikes were intruded along $\mathrm{D}_{4}$ faults introducing actinolite-bearing veins and biotite-clinozoisite $\pm$ actinolite alteration halos. Further movement on the $\mathrm{D}_{4}$ faults introduced superimposed (remobilised?) mineralisation $(\mathrm{Au}+\mathrm{Cu}+\mathrm{Mo}+\mathrm{W}$ quartz vein mineralisation and clinozoisite-biotite-pyrrhotitechalcopyrite veins).

The giant size and unusual geochemical features of the Boddington deposit would suggest at least two mineralisation events, even if researchers disagree about what these events represent (see 'Discussion’ section).

## SAMPLING, APPROACH AND METHODOLOGY

## Fieldwork

The Boddington deposit is confined between two main shear zones and the Central Diorite is placed within the southern part of the Open Pit; it is truncated to the South by an E-W fault (Figure A1a). A dense swarm of mafic (and locally ultramafic) dykes, both longitudinal and oblique, occur within the deposit (Figure A1b). Even though some of these dykes are Archean, the majority are Proterozoic, post-dating mineralisation. They are monitored because they represent mine waste.

Sampling was carried out in April 2011 in the Open Pit at RL 108 (11 samples) and in 4 drillcores ( 42 samples) (Figure 1) where molybdenite is present. One drillcore intersects the Wourahming granite ( $\sim 1 \mathrm{~km} \mathrm{SE}$ from the other drillcores). The granite plunges westwards and intersects the South-East Shear on the eastern side (Figure A2a, b). High-grade Mo and Au intervals were sampled from three drillcores in the Central Diorite (Figure A3a, b). Sampling intervals were located within the high-grade core defined as 'Southern Diorite Deep'. Sulphide mineralisation was also found in one of the drillcores intersecting the deepest margin of the Central Diorite. Samples from the Open Pit were collected from high Au-grade blocks (Figure A4). Drillcore intervals were chosen on the basis of assays (Figure A5), targeting the overlap between high-grade Au, Mo and Bi. Samples were selected on the basis of visible sulphides, including molybdenite. Details of the sampling intervals in all four drillcores, highlighting lithology and mineralisation, are presented in Figure A6-9.

## Analytical Methods

A Nikon optical microscope equipped with a digital camera was used in reflected light mode (magnifying lenses up to 50x) to identify mineral associations and relevant textures.

Philips XL-30 and XL-40 scanning electron microscopes (SEM) at Adelaide Microscopy were used to examine selected sections. These instruments possess back-scatter electron (BSE) detectors to enable observation of subtle compositional differences. Both instruments were operated at an accelerating voltage of 20 eV , with a spot size of 4 (XL-30) or 4-6 depending on the mineralogy (XL-40). Special attention was given to micron-scale inclusions and fine-scale compositional zoning within minerals. Both SEMs are equipped with EDAX software permitting generation of semi-quantitative compositional information on 1-2 $\mu \mathrm{m}$ -
sized spots. Imaging and EDAX analysis were also used to select specific target areas for electron probe microanalysis (EPMA) and laser-ablation inductively-coupled plasma mass spectroscopy (LA-ICP-MS).

A CAMECA SX-51 EPMA instrument at Adelaide Microscopy was used to provide quantitative compositional data, mineral stoichiometry and chemical characterisation of zoning patterns in sulphides and gangue minerals. Standards and spectral lines for rockforming silicates were: Na (albite, $\mathrm{K} \alpha$ ), Mg (almandine, $\mathrm{K} \alpha$ ), Al (almandine, $\mathrm{K} \alpha$ ), Si (almandine, $\mathrm{K} \alpha$ ), P (apatite, $\mathrm{K} \alpha$ ), Cl (tugtupite, $\mathrm{K} \alpha$ ), K (sanidine, $\mathrm{K} \alpha$ ), Ca (apatite, $\mathrm{K} \alpha$ ), Ti (rutile, $\mathrm{K} \alpha$ ), Cr (pyrope, $\mathrm{K} \alpha$ ), Mn (rhodonite, $\mathrm{K} \alpha$ ), Fe (almandine, $\mathrm{K} \alpha$ ) and F (fluorite, $\mathrm{K} \alpha$ ). Standards and spectral lines for ore minerals were: $\mathrm{Au}(\mathrm{Au}$ metal, $\mathrm{M} \alpha), \mathrm{Ag}\left(\mathrm{Ag}_{2} \mathrm{Se}, \mathrm{L} \alpha\right), \mathrm{Cu}$ $\left(\mathrm{Cu}_{2} \mathrm{~S}, \mathrm{~K} \alpha\right), \mathrm{Zn}(\mathrm{ZnS}, \mathrm{K} \alpha), \mathrm{Fe}\left(\mathrm{FeS}_{2}, \mathrm{~K} \alpha\right), \mathrm{S}\left(\mathrm{FeS}_{2}, \mathrm{~K} \alpha\right)$, $\mathrm{Se}\left(\mathrm{Ag}_{2} \mathrm{Se}, \mathrm{L} \alpha\right)$, Co (Co metal, $\left.\mathrm{K} \alpha\right)$, $\mathrm{Ni}(\mathrm{Ni}$ metal, $\mathrm{K} \alpha)$, As (AsGa, $\mathrm{L} \alpha$ ), $\mathrm{Cd}(\mathrm{CdS}, \mathrm{L} \alpha)$, $\mathrm{Mn}(\mathrm{MnS}, \mathrm{K} \alpha), \mathrm{Te}\left(\mathrm{Bi}_{2} \mathrm{Se}_{3}, \mathrm{~L} \alpha\right), \mathrm{Pb}(\mathrm{PbS}$, $\mathrm{M} \alpha)$ and $\mathrm{Bi}\left(\mathrm{Bi}_{2} \mathrm{Se}_{3}, \mathrm{M} \alpha\right)$.

A LA-ICP-MS system located at CODES, University of Tasmania was used to acquire the trace element data for molybdenite. This machine is a 7500 series ICP-MS instrument coupled with a UP-213 laser-ablation system. The following isotopes were measured: ${ }^{57} \mathrm{Fe}$, ${ }^{59} \mathrm{Co},{ }^{60} \mathrm{Ni},{ }^{65} \mathrm{Cu},{ }^{66} \mathrm{Zn},{ }^{75} \mathrm{As},{ }^{77} \mathrm{Se},{ }^{107} \mathrm{Ag},{ }^{118} \mathrm{Sn},{ }^{121} \mathrm{Sb},{ }^{125} \mathrm{Te},{ }^{182} \mathrm{~W},{ }^{185} \mathrm{Re},{ }^{197} \mathrm{Au},{ }^{204} \mathrm{~Pb},{ }^{207} \mathrm{Bi}$, ${ }^{238} \mathrm{U}$, i.e. the same element set as used by Kalleske (2010). In-house standard STDGL2b2 was used for calibration (Danyushevskiy et al. 2011). Pre-ablation at 1 Hz was undertaken to clean the surface of contaminants. Sample measurement took place at a laser frequency of 5 Hz and $35 \mu \mathrm{~m}$ spot size. Analysed isotopes were selected to avoid interference from other isotopes with the same weight or a combination of isotopes with the same weight.

In addition to spot analysis, 4 element maps were created using the same instrument (see Large et al. 2009 for experimental details). Each mapped a different grain of molybdenite in four samples (NK25, RG10, RG27 and RG45). The maps highlight grain-scale element distributions. Map NK25 was compiled using $15 \mu \mathrm{~m}$ resolution, map RG10 used a $10 \mu \mathrm{~m}$ resolution and maps RG27 and RG45 used an $8 \mu \mathrm{~m}$ resolution with each map taking 3-6 hours to complete. Mapped elements were $\mathrm{Al}, \mathrm{Si}, \mathrm{S}, \mathrm{Ca}, \mathrm{Fe}, \mathrm{Co}, \mathrm{Ni}, \mathrm{Cu}, \mathrm{As}, \mathrm{Se}, \mathrm{Mo}, \mathrm{Ag}, \mathrm{Sb}$, $\mathrm{Te}, \mathrm{W}, \mathrm{Re}, \mathrm{Au}, \mathrm{Hg}, \mathrm{Pb}$ and Bi in all cases.

Focussed Ion Beam - SEM (FIB-SEM) methods were used for high-resolution crosssection imaging of $\mu \mathrm{m}$ and sub- $\mu \mathrm{m}$-sized inclusions in molybdenite and for extraction and thinning of in-situ foils for Transmission Electron Microscopy (TEM). A FEI-Helios
nanoLab DualBeam system at Adelaide Microscopy was used, following methodology described by Ciobanu et al. (2011). Standard operating conditions are 30 kV for ion milling and 20 kV and 5 kV for normal and immersion mode imaging, respectively, using the electron beam.

TEM study (electron diffraction and high-resolution) imaging of the FIB-prepared foils was performed on a Philips 200CM TEM instrument (Adelaide Microscopy) operated at 200 kV . The instrument is equipped with a double-tilt holder and Gatan digital camera. Measurements on the diffractions were performed using DigitalMicrograph ${ }^{\mathrm{TM}}$ 3.11.1.

## RESULTS

## Lithologies and alteration assemblages

## PETROGRAPHY AND MINERALOGY

At Boddington the mineralisation occurs as disseminations and veinlets throughout several types of magmatic rocks, including diorite, andesite, rhyodacite, dolerite and granite, all with various degrees of alteration. All these rocks, except the granite, are deformed and metamorphosed at lower amphibolite facies (biotite-plagioclase present). The present study was focused on sulphide-rich samples from diorite and granite in the southern part of the deposit. For the sake of comparison, a sample of dolerite with scattered sulphides (RG19) from the same part of the deposit was also analysed. Mineral associations (ore minerals and gangue) in the 53 polished blocks are listed in Table 1 and abbreviations for mineral names throughout the tables and text are given in Table 2.

Diorite is a term attributed to foliated rocks consisting of variable amounts of plagioclase, quartz, biotite and subordinate muscovite. This rock can vary from coarse- to very finegrained (Figure A10a, b) and in some cases shows extremely sheared mylonitic texture; in the latter the proportion of muscovite increases. Albitisation of diorite is seen throughout large intervals in the lower half $(400-900 \mathrm{~m})$ of drillcore WBD 105900002. This diorite is characterised by a prominent sheared porphyroblastic texture, in which the porphyroblasts consists of symplectites of quartz and albite.

In contrast, the granite and dolerite are both coarse-grained. The granite (Figure A10c) consists of perthitic feldspar (potassium feldspar with exsolutions of albite), plagioclase,
quartz and mica (both biotite and muscovite). In comparison with the diorite, the granite is less altered and the alteration is mostly confined to, or in and around the sulphide patches. A micro-grained aplitic variety occurs at depth (Table 1). The studied dolerite is a quartzbearing variety in which the pyroxene is pseudomorphed by hornblende and Fe -Ti-oxides (Figure A10d, e) and the plagioclase is minor and largely replaced by epidote.

This study focused on alteration assemblages that are tied to (hydrothermal) sulphide deposition, rather than those resulting from metamorphism. Such alteration assemblages, seen in both diorite and granite, include actinolite, intermediate members of the clinozoisiteepidote (Clz/Ep) group, biotite, chlorite, albite, K-feldspar (adularia), quartz, stilbite and clay minerals.

Minor components of the alteration assemblages seen in both diorite and granite samples are Fe -Ti-oxides (ilmenite, rutile), titanite, apatite, scheelite, fluorite, thorite and a variety of REE-bearing minerals as well as niobates. Apatite and scheelite are widespread and in some samples very abundant (e.g. RG52, Figure A10f). Among the niobates, fergusonite, (Y,REE) $\mathrm{NbO}_{4}$, is quite common in the granite (Figure A 10 g ) whereas Ca-niobates are present in both granite and diorite. Hochelagaite, $\left[(\mathrm{Ca}, \mathrm{Na}, \mathrm{Sr}) \mathrm{Nb}_{4}\right]_{11} \cdot 8 \mathrm{H}_{2} \mathrm{O}$, was identified In the latter, together with scheelite as dusty inclusions within ilmenite surrounding a core of rutile (Figure A10h, i). Fluorite is more abundant in the granite where it can occur as veinlets together with sulphides; in the diorite fluorite is seen as small patches tied to the sulphides.

It is very difficult to establish specific mineral sequences since all these minerals are ubiquitously present, except for actinolite and stilbite which are much rarer. In most cases, the abovementioned minerals are associated, in variable proportions, with one another. Illustrating the complexity of the alteration assemblages, for example, in a diorite sample (RG52), actinolite was seen filling the inner part of a quartz veinlet with sulphides where the veinlet selvage includes both epidote and biotite (Figure A10j). The most common and abundant alteration mineral, always seen embedding the sulphide patches, is $\mathrm{Clz} / \mathrm{Ep}$ (Figure A10k, 1). This often displays oscillatory zonation patterns seen on the back scatter electron (BSE) images as various shades of grey (Figure A10m). Biotite (Figure A10n) is the second most common alteration mineral and is tied to the sulphide occurrence even though it is difficult to discriminate between different generations of biotite. In diorite, some of the biotite is formed by replacement of mafic minerals (pyroxene/amphibole). In this case it occurs as coarse grains containing abundant inclusions of titanite. In contrast, the hydrothermal biotite (associated with sulphide deposition) forms short lamella packages.

Chlorite is less abundant but nonetheless widespread. Chamosite, the Fe-rich variety, is seen interlayered with muscovite (Figure A100) - in particular in the granite samples.

A zeolite facies alteration is also present based on the fact that a Ca-rich variety of stilbite has been identified in three diorite samples (RG67, 27, 60). Stilbite occurs either intergrown with molybdenite (Figure A11a), replacing feldspars (Figure A11b, c), or forming veinlets. In the sample from the open pit (RG67) stilbite is found in an intensely-altered diorite featuring high porosity and silica + clay minerals replacement of feldspars (Figure A11d). Hydrothermal biotite interlayered with chlorite (Figure A11e) and overgrown by Clz/Ep (Figure A11f) is also observed in the same rock. These minerals coexist with allanite that contains dusty inclusions of REE-minerals (Figure A11g, h) and rims of thorite (Figure A11i),

## Silicate Mineral Chemistry

Electron microprobe analysis was carried out on the rock-forming silicates and alteration assemblages hosting the mineralisation. Results for feldspar, biotite, muscovite, epidote and chlorite are summarised (as means for each sample) in Tables A1-5; means for chlorite are given in Table 3.

## FELDSPAR

Plagioclase feldspars from the granite are Na-rich with a narrow compositional variation within the albite ( $\mathrm{Ab}_{99.1-99.7}$ ) or oligoclase $\left(\mathrm{Ab}_{82} \mathrm{An}_{16} ; \mathrm{RG} 50\right)$ fields. Similarly, the potassium feldspar is close to end-member Kfs showing a narrow variation in all samples except RG50 ( $\mathrm{Kfs}_{97.6-96.8}$ and $\mathrm{Kfs}_{86.6} \mathrm{Ab}_{12.8}$, respectively). The compositional spread for the plagioclase feldspars in the diorite is also in the albite-oligoclase range but, unlike in the granite, this variation is observed in each of the analysed samples; a single analysis plots in the andesine field. End-member potassium feldspar has been found in two samples (RG36). The Ab-AnKfs plots (Figure 2a, b) show there is little difference in terms of compositional variation for plagioclase but the granite has more variation in terms of K-feldspar composition.

## BIOTITE

The data for biotite in both granite and diorite shows little variation around an intermediate composition between the Fe -rich (annite) and Mg -rich (phlogopite) end-members ( $\mathrm{Ann}_{43}$ ${ }_{50} \mathrm{Phl}_{47-53}$ ). A slightly Fe-richer variety ( $\mathrm{Ann}_{57.6}$ ) is present in a granite sample (RG50) and a second population of biotite, a Mg -rich variety $\left(\mathrm{Phl}_{60.6}\right)$, is observed in a diorite sample (RG29). The biotite has variable but consistent F content in both granite and diorite ranging from 0.5-2.71 $\mathrm{wt} \%$, giving fluorphlogopite components in the range $10.8-32.5 \mathrm{~mol} \%$.

## MUSCOVITE

Muscovite also has little variation throughout the dataset for granite and diorite with a phengite component ranging from $8.5-18.2 \mathrm{~mol} \%$. The fluorine content in the muscovite is appreciably lower than in the biotite, ranging from $0.12-0.69 \mathrm{wt} \%$, corresponding to a $\mathrm{mol} \%$ F-end-member component in the range 1.2-7.4.

## EPIDOTE GROUP MINERALS

Data for minerals from the epidote group show the presence of dominantly Al-rich compositions with $\mathrm{X}_{\mathrm{Al}}$ between $76.4-82.9 \mathrm{~mol} \%$ and $72.2-86.3 \mathrm{~mol} \%$ in the granite and diorite, respectively. This indicates that clinozoisite is the mineral in both diorite and granite. The Mn component is minor throughout the dataset ( $<1 \mathrm{wt} \%$ ). The compositional plots (Figure 2c, d) show somewhat greater spread among the analyses from granite than those from the diorite. Compositional variation across profiles in individual grains that show oscillatory zonation (Figure A10m) indicate a narrower range of variation than the entire compositional interval with only 6-7 mol $\% \mathrm{X}_{\mathrm{Al}}$ difference across the profile. This means that in a given sample the variation in composition is higher from grain to grain than across the zonation.

## CHLORITE

In contrast to the other silicates chlorite shows a wider variation between different samples and from granite to the diorite. One of the granite samples (RG42) is characterised by varieties richest in $\mathrm{Fe}(94.9 \mathrm{~mol} \%$ chamosite, the Fe -end-member). Otherwise, the chamosite component in other granite samples ranges from 70.9 to $81.7 \mathrm{~mol} \%$. In the diorite, Mg -richer
compositions are noted; $\mathrm{Fe} /(\mathrm{Fe}+\mathrm{Mg}+\mathrm{Mn})$ ranges from 0.38 to 0.78 , inferring that both chamosite and clinochlore (Mg end-member) are present. Mn contents are consistently <0.7 $\mathrm{wt} \%$. $\mathrm{Al}^{\mathrm{vi}}$ (octahedral Al ) varies from 1.2 to 1.6 , with no statistical difference between granite and diorite. In the diorite sample RG27, there are two distinct compositional types of chlorite: one is appreciable richer in Fe ( 78.5 against $63.5 \mathrm{~mol} \%$ chamosite) but poorer in $\mathrm{Al}^{\text {vi }}$ ( 1.26 against 1.36 a.p.f.u.). This suggests two distinct generations (see below). The ternary $\mathrm{Fe}-\mathrm{Mg}$ $\mathrm{Al}^{\mathrm{vi}}$ plot (Figure 2e) discriminates chamosite in the granite from the intermediate compositions in other samples.

Application of the chlorite geothermometer (Cathelineau 1988, Jowett 1991, De Caritat et al. 1993) uses the $\mathrm{Al}^{\mathrm{iv}}$ (tetrahedral) component in chlorite to derive a peak metamorphic temperature. The calibrations of Cathelineau (1988) and Jowett (1991) are used: for chlorites with $\mathrm{Fe} /(\mathrm{Fe}+\mathrm{Mg}+\mathrm{Mn})<0.6$, whereas the calibration of Cathelineau (1988) is preferred for the Fe-rich varieties. Realistic temperatures are obtained (Table 3), ranging from 194 to $366{ }^{\circ} \mathrm{C}$ in the granite and 219 to $421^{\circ} \mathrm{C}$ in the diorite. The two populations in RG 27 give temperature estimates at the minimum of the range $\left(219{ }^{\circ} \mathrm{C}\right)$ and towards the upper limit $\left(367^{\circ} \mathrm{C}\right)$, further suggesting these two populations were formed under different conditions. Temperature estimates from individual analyses are also summarised in Figure 2f-h.

## Mineralisation

The major components of the $\mathrm{Cu}-\mathrm{Mo}$ mineralisation at Boddington are chalcopyrite, pyrrhotite and molybdenite. Other sulphides identified here are cubanite, pyrite, sphalerite, stannite, pentlandite, argentopentlandite, a $\mathrm{Fe}-\mathrm{Ni}$-thiospinel possibly violarite, $\mathrm{Co}-$ mackinawite, bornite, chalcocite and covellite; native copper was also found in one of the samples with zeolite alteration (RG67). The Au mineralisation at Boddington is characterised by an $\mathrm{Au}-\mathrm{Ag}-\mathrm{Bi}-(\mathrm{Pb})-\mathrm{Te}-(\mathrm{Se})$ mineral association which includes native gold, electrum, maldonite $\left(\mathrm{Au}_{2} \mathrm{Bi}\right)$, native bismuth, hessite $\left(\mathrm{Ag}_{2} \mathrm{Te}\right)$, altaite $(\mathrm{PbTe})$, intermediate members of the galena-clausthalite series $(\mathrm{PbS}-\mathrm{PbSe})$, naummanite $\left(\mathrm{Ag}_{2} \mathrm{Se}\right)$, abundant Bi-tellurides from the tetradymite group (Cook et al. 2007a) and subordinate $\mathrm{Bi}-\mathrm{Pb}$-sulphotellurides from the aleksite series (Cook et al. 2007b).

## PETROGRAPHY AND MINERALOGY OF SULPHIDES

Most commonly, the sulphides occur as narrow veinlets (mm-cm size) or as patches and disseminations (Figure A12). Chalcopyrite is often associated with pyrrhotite and in some cases with molybdenite; the latter however is often found in monomineralic patches of cm to mm-size. All these minerals are found in both granite and diorite samples. Also common in the two types of rocks is the presence of sphalerite where the latter is found as small grains at the boundaries between chalcopyrite and pyrrhotite.

In the diorite the sulphide assemblages are more varied. In many cases pyrrhotite is present as two compositionally distinct species (see below), with lamellar exsolution textures recognisable on the BSE images by the different shades of grey (Figure A12a). Either of the two species can be the host for lamellar exsolutions. Cubanite, commonly seen in minor amounts as lamellar exsolution in chalcopyrite, is also the main sulphide in a couple of samples (RG24 and RG59). In sample RG24 cubanite occurs as a fine dissemination along the foliation in the diorite. In many grains, skeletal exsolution of pyrrhotite (Figure A12b), as well as small inclusions of sphalerite and pentlandite are present. In the same sample patches of chalcopyrite/pyrrhotite host grains of pentlandite that is compositionally inhomogeneous; argentopentlandite (Figure A12c) was identified within such grains. In sample RG59 the cubanite contains fields of tiny exsolutions of stannite + sphalerite (Figure A12d). Coarser inclusions of sphalerite and stannite (up to several microns; Figure A12e) also occur either throughout the cubanite or at the margin between cubanite and pyrrhotite. The association stannite+sphalerite is also found coexisting with pyrrhotite in a sample where chalcopyrite is the main copper sulphide (RG27; Figure A12f).

Pyrite is rare and is found either coexisting with (RG59) or without pyrrhotite (RG17). In the latter case, however, lamellar textures are observed indicating replacement of pre-existing pyrrhotite. In the dolerite, pyrite is the main sulphide and contains rounded inclusions of chalcopyrite. Cobalt-nickel minerals are represented by Co-mackinawite (RG11; Figure A12g) and $\mathbf{F e}$-Ni-thiospinel (RG36, RG59). These were identified from two samples in associations containing pyrrhotite; in RG59, it also coexists with pyrite. In sample RG59, the thiospinel occurs as unusual intergrowths with (REE-bearing) (bastnäsite) (Figure A12h-j).

Apart from chalcopyrite, very minor amounts of other $\mathrm{Cu}-(\mathrm{Fe})$-sulphides (bornite, covellite, chalcocite) are found in two samples. Bornite and covellite (Figure A12k) are seen along thin veinlets or grain boundaries of chalcopyrite where some of the grains contain
lamellar cubanite (RG25). Several grains of chalcocite (Figure A12l) occur in association with native copper (Figure A12m) in a sample from the open pit containing zeolite alteration (RG67). Native copper is quite abundant in this sample where is present as visible disseminations and thin veinlets or staining the rock surfaces. Some of the grains occur as inclusions in epidote or stilbite; in former case they have a narrow rim with clay minerals.

## MOLYBDENITE

Molybdenite is present in 24 samples, a third of which are from the granite. In the majority of cases, it is associated with chalcopyrite and pyrrhotite. It displays great morphological variety from bundles and knots to lamellar aggregates along the foliation or single lamella scattered in the surroundings of the larger patches (Figure 3). It is commonly coarse-grained with contorted and kink deformation lamellae (Figure 3a-c). In some samples, however such coarse aggregates coexist with meshes formed by shorter lamellae (RG64, Figure 3d) and also aggregates of idiomorphic short lamellae.

In the present sample suite a conspicuous feature of the molybdenite is the presence of inclusions (up to several microns; Figure 3e) of different minerals (see also next section). In molybdenite from both granite and diorite Bi-tellurides from the tetradymite group are abundant; galena and hessite are also noted. Phases from the aleksite series are present in the granite whereas chalcopyrite, native gold, electrum, native bismuth and minor altaite are seen in the diorite.

The molybdenite in the granite displays regular, dense fields of telluride inclusions. In the diorite, however, chalcopyrite inclusions are regular and dense (Figure 3f). In the latter case all other inclusions, including tellurides, have an irregular distribution and range in abundance from scattered to very dense in certain areas. In both cases the inclusions are seen preferentially along the molybdenite lamellae with a tendency to coarsening across intervals of kink deformation or in the axial plane of tight microfolds (Figure 3g, h). Ductile remobilisation of inclusions is also observed along microshears oblique to the molybdenite lamellae (Figure 3i). The thicker chalcopyrite inclusions display pointing edges indicating they are formed under stress conditions (Figure 3f).

## Sulphide Mineral Chemistry

Compositional variation in the sulphides was measured using EPMA. Results are presented in terms of means and standard deviation for each sample or distinct populations in a given sample in Tables 4-7 and Table A6 in Appendix.

Pyrrhotite, $\mathrm{Fe}_{1-\mathrm{x}} \mathrm{S}$, where $0<\mathrm{x}<0.125$, is an interesting mineral group both in geological and crystallographic terms. The pyrrhotite structures have hexagonal close packed (hcp) layers of S with Fe atoms occupying octahedral interstices between the S layers. All octahedral positions are occupied in troilite ( FeS ; hexagonal), whereas the non-stoichiometric pyrrhotites have vacancies on the Fe positions (Posfai \& Buseck 1997). The most common forms in nature are monoclinic 4C pyrrhotite (ideally $\mathrm{Fe}_{7} \mathrm{~S}_{8}$ ) and hexagonal 5C pyrrhotite (ideally $\mathrm{Fe}_{9} \mathrm{~S}_{10}$ ) (Becker et al. 2010).

The EPMA data shows distinct compositions for Po in the granite relative to the diorite and also distinct compositional fields for the 2-phase Po with lamellar exsolution textures identified on BSE images (Figure A12a) in diorite (Table 4). Compositional variation is expressed as $\mathrm{M} / \mathrm{S}$, where $\mathrm{M}=\mathrm{Fe}+\mathrm{Co}+\mathrm{Ni}+\mathrm{Cu}$, ranging from 0.98 to 0.87 in terms of mean values. Nickel is present in highest amounts (means of 0.19 and $0.2 \mathrm{wt} \%$ ) in the dark lamellae from RG52 and RG24 and is detectable (means of 0.05 and $0.09 \mathrm{wt} \%$ ) in 2 other samples, one of which is the single phase Po (RG27). Cobalt is present in detectable amounts only in 2 samples, both from the single phase Po, one from the granite (RG48) and one from the diorite (RG66). Copper is detectable in a majority of the samples with mean values between 0.25 and $0.07 \mathrm{wt} \%$.

The plot of M versus S shows that the single phase Po corresponds to the $\mathrm{Fe}_{7} \mathrm{~S}_{8}$ field in the granite and spreads into the field of $\mathrm{Fe}_{9} \mathrm{~S}_{10}$ for single phase Po in the diorite (Figure 5a, b). About a third of the Po in the granite analyses plot in a field with still lower M, which is difficult to explain unless we consider that the Po is replaced by Py and the latter is not observable on the BSE images (submicroscopic grain replacement). There is a clear compositional gap between the 'dark' (showing a wide spread across $\mathrm{Fe}_{7} \mathrm{~S}_{8}$ to $\mathrm{Fe}_{10} \mathrm{~S}_{11}$ ) and 'bright' lamellae (towards the upper limit of stoichiometric FeS) from the 2-phase Po in the diorite (Figure 5c).

Sphalerite (Table 5) from both diorite and granite samples is rich in Fe , with $\mathrm{mol} \% \mathrm{FeS}$ in the range between 14.23 (RG48) and 20.67 (RG24). Detectable amounts of Cd were found in all samples (means from $0.22 \mathrm{wt} \%$ in RG27 to $1.16 \mathrm{wt} \%$ in RG48). Copper is also present in
the majority of samples with mean values between $0.27 \mathrm{wt} \%$ (RG66) and $1.72 \mathrm{wt} \%$ (R59). Stannite, present together with sphalerite in 2 diorite samples shows variable $\mathrm{Fe}: \mathrm{Zn}$ ration from 1.624:0.376 to 1.905:0.103 in samples RG59 and RG27, respectively.

Co-existing Sph-Stn pairs were used to calculate formation temperatures based on equilibrium $\log X_{\mathrm{FeS}} / \mathrm{X}_{\mathrm{ZnS}}$ ratios in the two minerals (Table 5) following the calibration of Shimizu \& Shikazono (1985). Even though the Stn grains are relatively small to ensure accurate measurements, the geothermometer gave realistic temperatures of $\sim 350{ }^{\circ} \mathrm{C}$ a cluster of points (Figure 5d).

Pyrite was measured in one sample (RG18, Table A6) and shows detectable amounts of Co and Ni (means 0.08 and $0.04 \mathrm{wt} \%$, respectively). Although As was not measured, the good totals indicate that, if present, it is only present at very minor levels, concordant with pyrite data given in Kalleske (2010). Pentlandite shows stoichiometric composition whereas argentopentlandite is slightly metal deficient, probably due to analytical error brought in by the small grain size (Table A6). Copper is present in both minerals (means 1.19 and 2.17 $\mathrm{wt} \%$, respectively); mean Co contents are 4.68 and $0.19 \mathrm{wt} \%$ in Pn and $\mathrm{Ag}-\mathrm{Pn}$, respectively. Two different Co-Ni-bearing varieties of stoichiometric ( $\mathrm{M}: \mathrm{S}=1$ ) mackinawite are present (Table A6). The Fe-Ni thiospinel phase (Wagner \& Cook 1999) gives compositions along the $\mathrm{Fe}_{3} \mathrm{~S}_{4}$ (greigite) - $\mathrm{Ni}_{3} \mathrm{~S}_{4}$ (polydymite) join, but conspicuously closer to $\mathrm{Fe}_{2} \mathrm{NiS}_{4}$ than the named mineral $\mathrm{FeNi}_{2} \mathrm{~S}_{4}$ (violarite).

## AU-AG-BI-PB-TE-SE ASSOCIATION

Gold and silver minerals (native gold, electrum, maldonite and hessite; Table 6) as well as tellurides (Figure 4) were found in the diorite samples, both as inclusions within molybdenite and as clusters of minute grains (up to 10 microns) in the alteration assemblages surrounding the sulphides; the latter type of occurrence is also present in the granite. Even though native gold and electrum are present in larger amounts as monomineralic filaments or small droplets/patches (RG60, RG62) (Figure 4a, b), these are also observed together with Bi- and Pb -tellurides (RG25; Figure 4c) within molybdenite from the diorite.

The genetic link between gold and bismuth minerals is clearly seen in the presence of crystallised droplets containing the assemblage maldonite + hedleyite + bismuth (RG52, Figure 4d), an association representing the equivalent of the eutectic at $235^{\circ} \mathrm{C}$ in the $\mathrm{Au}-\mathrm{Bi}-$

Te system (Prince et al. 1990) Such an association is part of the larger clusters including native gold, native bismuth, hessite and several Bi-tellurides from the Bi-rich end of the tetradymite group (Figure 4e, f). These clusters are tied to local brecciation and chloritisation superimposed onto an earlier alteration assemblage, e.g. actinolite-quartz-epidote in RG52. Hessite is always associated with native bismuth $\pm$ bismuth tellurides and is relatively widespread as inclusions in either molybdenite, other sulphides or in the gangue (Figure 4 g i).

The bismuth tellurides (Table 7) found in molybdenite are either tsumoite or species with ratios of $\mathrm{Bi} / \mathrm{Te}>1$ (including several unnamed species) in the granite (Figure 4 j ) and diorite, respectively. Several tellurosulphides from the aleksite series (saddlebackite, aleksite and phase ' C '; see below) were found in the molybdenite from the granite. Variation in the sulphur content is observed in the Bi-tellurides occurring within molybdenite but not in those from the outside clusters.

Other tellurides and selenides identified by EDAX analysis include altaite (Figure 4k), clausthalite, ikunolite (Figure 41), laitakarite and naummanite.

## Trace Element Chemistry of Molybdenite (LA-ICP-MS data)

Molybdenite from 19 samples was analysed using LA-ICP-MS to determine trace element concentrations. The sample selection comprises 6 granite and 13 diorite samples, of which one was a sample from the ABreccia orebody (NK25) from the Kalleske (2010) sample suite. 250 individual spot analyses were carried out. Results are summarised as means, standard deviations, minima and maxima for each sample in Table 8; the full dataset is listed in Table A7, Appendix.

## ELEMENT SPECTRA AND CONCENTRATIONS

In molybdenite from both diorite and granite, some elements, e.g. Re, W and Se , show flat signals on the time-resolved spectra. Most chalcophile elements ( $\mathrm{Bi}, \mathrm{Te}, \mathrm{Ag}, \mathrm{Pb}$ ) and Au typically display irregular signals during spectra acquisition. Representative spectra for both diorite and granite are shown in Figure 6.

Rhenium concentrations in molybdenite from diorite (expressed as mean within an individual sample) vary across 4 orders of magnitude, from $<1$ ( 0.16 ; RG29) to 374 ppm
(NK25). Mean values of several hundred ppm Re were all obtained from samples from the open pit (NK25, RG8, RG10) and one sample from the drillcores gives a mean of tens of ppm $\operatorname{Re}(36 \mathrm{ppm}$; RG64). All other samples had mean Re concentrations of $\ll 10 \mathrm{ppm}$. The range of Re concentrations in individual spot analyses across the diorite sample suite is 0.03 ppm (RG29, RG30) to 905 ppm (NK25). In terms of variation within a single sample, the highest variation is seen in one sample (RG64), which shows a range over 3 orders of magnitude. All samples from the open pit show the least intra-sample variation. Results obtained on samples from the open pit in the Central Diorite are comparable in terms of Re concentration with some of the highest concentrations obtained in the ABreccia orebody sample at a difference in depths of $\sim 42 \mathrm{~m}$. Maximum values of hundreds of ppm Re are also recorded from some of the drillcore samples that intersect mineralisation $\sim 200 \mathrm{~m}$ below, e.g. 347 and 218 ppm for RG27 and RG64, respectively. In contrast with the diorite, molybdenite from granite has means Re concentrations that are consistently low ( 0.4 to 18 ppm ).

Mean concentrations of W in molybdenite from diorite samples show a range from 79 to 306 ppm . The range of W values within each diorite sample is low, although one sample from the open pit (RG10) displays a variation up to 2 orders of magnitude. Mean W concentrations in the granite samples show a spread from 69 to 206 ppm . Tungsten variation in the granite is low with all samples showing a range of no more than one order of magnitude.

Mean Au concentrations in molybdenite from the diorite range from 9.4 (RG27) to 116 ppm (RG30), with a maximum value of 1,110 and a minimum value of 0.01 ppm . The highest variation within a single sample is 4 orders of magnitude (RG30 and RG65). Variation is, in general, always greater than one order of magnitude. Although comparably higher mean Au concentrations and maximum values are obtained for the two samples from Central Diorite in the open pit than in drillcores at depth, there is a suggestion that the shallower samples are poorer in Au than the deeper ones. In addition, shallower samples from the Central Diorite have a higher spread in Au values than that of the ABreccia. The mean Au concentrations in molybdenite from granite varied between 0.01 and 11 ppm , with 4 of the 6 samples having all values <mdl and an individual maximum value of only 53 ppm (RG40).

Mean Ag concentrations in molybdenite from the diorite range from 2.2 to 96 ppm . The range in variation in Ag is high (up to 4 orders of magnitude; RG65). In the granite, Ag means range between 1.3 and 32 ppm .

Means for Te in molybdenite from the diorite vary between 46 and 865 ppm with individual values of $20-3,912 \mathrm{ppm}$. The variability in Te concentration within each sample is low (up to 2 orders of magnitude). The range of variation for Te means in the granite is between 36 and 5,093 ppm (maxima and minima are 35,599 and 8.0 ppm , respectively). Two of the 6 samples (RG49, RG51) have much lower Te concentration (tens of ppm) in comparison with the other 4 which show hundreds or thousands of ppm Te . These granite samples are slightly finer-grained than the other 4.

Mean Bi concentrations in molybdenite from the diorite show variation between 113 and $2,929 \mathrm{ppm}$ (RG25, RG27) with maxima of $12,108 \mathrm{ppm}$ and minima of 1.3 ppm (RG25, RG29, respectively). Bi values in the diorite samples display greater variation than Te (up to 3 orders of magnitude). In granite molybdenite, Bi shows a still higher range of variation; mean concentrations are from 32 to 7,885 , with the maximum $(52,682)$ and minimum $(0.02$ $\mathrm{ppm})$ concentrations in molybdenite from the same samples (RG45, RG51). The highest variation in a single sample is up to 5 orders of magnitude (RG40). The lowest Bi concentrations are recorded in the fine-grained granite samples.

Mean Pb concentrations in molybdenite from the diorite ranges from 57 (RG62) to 1,532 ppm (RG25), with a maximum individual value of 3,440 and a minimum of 0.28 ppm . The range of Pb concentration in each sample shows a high variation (up to 3 orders of magnitude; RG29 and RG64). Mean Pb concentrations for molybdenite in the granite show variation between 146 (RG51) and 4,774 pm (RG45). Variation within individual samples is high (up to 3 orders of magnitude).

Selenium concentrations in molybdenite from the diorite show a range from 311 to 1,149 ppm, with a maximum value of $1,543 \mathrm{ppm}$ and minimum of 268 ppm . Molybdenite in the granite has mean Se concentrations that range from 205 to 444 ppm . Unlike most other elements, Se has a low variance within each sample in both diorite and granite.

Detectable concentrations of Co and Ni are recorded in molybdenite from the diorite. Mean concentrations are $0.85-14 \mathrm{ppm}$ and $2.2-53 \mathrm{ppm}$ for Co and Ni , respectively. In molybdenite from the granite, Co and Ni concentrations are extremely low ( $0.44-3.0 \mathrm{ppm}$ and $0.06-0.27 \mathrm{ppm}$, respectively).

Mean Fe concentrations in molybdenite from the diorite lie in the range 1,390-26,853 ppm (samples RG60 and RG25, respectively). Samples from the granite show mean
concentrations of 420-5,549 for Fe in molybdenite. Variance within individual samples is high in both rock types (up to 3 orders of magnitude).

Copper is present at detectable concentrations (means $15-1,261 \mathrm{ppm}$ in the diorite and 2.1400 ppm in the granite). Overall the diorite samples have higher Cu concentrations in molybdenite than the granite. Detectable amounts of several other elements (Zn, As, $\mathrm{Sn}, \mathrm{Sb}$, Tl and U ) are also recorded in the analysed molybdenites and in some samples show somewhat higher values.

## HISTOGRAMS

The distribution of trace elements in molybdenite from both diorite and granite is expressed in histograms (Figure 7). Elements which are commonly incorporated within the molybdenite structure (i.e. Re and W) show contrary distribution patterns in molybdenite from the two rock types. The rhenium distribution is bimodal and W shows a normal distribution in diorite, but in granite Re is normally distributed and W is bimodal (Figure 7a, b). The population from molybdenite in granite overlaps with the low-concentration population in the diorite for Re. Histograms for Au and Ag in molybdenite from both granite and diorite show normal distributions (Figure 7c, d). The distribution of Au is however positively skewed for diorite with granite in the lower concentration range. Silver shows the best similarity between the population in the diorite and granite. The 3 chalcophile elements $(\mathrm{Bi}, \mathrm{Te}, \mathrm{Pb})$ associated with Au and Ag (see section above) show broadly normal distributions in the entire dataset (Figure $7 \mathrm{e}-\mathrm{g}$ ) although there is a slight positive skew towards higher concentrations in the Bi and Pb populations. In the granite population there is a bimodal tendency for all 3 elements, best expressed for Te , largely due to the fact that molybdenite in the fine-grained granite is characterised by lower concentrations of these elements (see above).

## VARIATION TRENDS

Plots of element concentration in molybdenite against depth (Figure 8) are illustrative of the spread in variation across the depth interval sampled in the Central Diorite ( $\sim 250 \mathrm{~m}$ ) and in relation to the sample from ABreccia ( 42 m above). Rhenium shows a slight decreasing trend and a wide variation with depth (Figure 8a). The wide variation is seen in those diorite samples that overlap with the granite depth. In contrast, W shows no apparent trend or range of variation with depth (Figure 8b). Gold shows a wider variation at depth but indication of a
decreasing trend with depth if the 2 fine-grained granite samples that record the lowermost values ( $<0.1 \mathrm{ppm}$ ) are included (Figure 8c). Silver shows no trend but a wider range of variation with increasing depth (Figure 8d); in this case the fine- and coarse-grained granite samples have comparable ranges of variation in Ag concentration. A clear increasing trend and wider spread of variation with depth is seen for Te , whereas both Bi and Pb show comparable wide ranges of variation but no trends with depth (Figure 8e-g). For all three elements the highest values and widest ranges are in molybdenite from the granite but the finer-grained granite variety records the lowermost concentrations in each case. For Pb there is, however, an overlap in the range of values with those of the coarse-grained granite.

Binary plots of Au against chalcophile elements (Figure 9a-c) display much higher correlation coefficients for molybdenite from diorite ( $r^{\prime}=0.79$ for $\mathrm{Au}-\mathrm{Bi} ; \mathrm{r}^{\prime}=0.65$ for $\mathrm{Au}-\mathrm{Te}$; $r^{\prime}=0.69$ for $\mathrm{Au}-\mathrm{Pb}$ ) than the granite ( $\mathrm{r}^{\prime} \sim 0.50$ for $\mathrm{Au}-\mathrm{Bi}, \mathrm{Au}-\mathrm{Te}, \mathrm{Au}-\mathrm{Pb}$ ). The $\mathrm{Au}-\mathrm{Ag}$ plot (Figure 9d) is, however, an exception in that the trends are parallel ( $\mathrm{r}^{\prime}=0.64$ ) for the 2 sample groups. A strong distinction is seen in the correlation plot between Ni and Co (Figure 9e), where the correlation coefficient for the diorite population $\left(r^{\prime}=0.91\right)$ is much higher than that for the granite population ( $r^{\prime}=0.38$ ). The plot of Re versus W shows a funnel-shaped distribution where the highest Re values (in diorite) record the highest spread in W and this spread narrows down to intermediate $\operatorname{Re}$ values (in granite) and further to the lowest $\operatorname{Re}$ values (in other diorite samples) (Figure 9f). The plot for Au versus Re (Figure 9g) shows that the diorite has the highest values in Au over a wide range of Re and that the analyses from granite cluster within a distinct field with the lowermost values for both $\operatorname{Re}$ and Au .

## GRAIN SCALE INHOMOGENEITY

In molybdenite from the diorite the spot analyses demonstrate a high degree of grain-scale inhomogeneity (Figure 10). This depends, to some extent, upon textures such as bending, kink deformation, grain boundaries and replacement by silicates or other sulphides such as chalcopyrite. For example, spot analyses taken along a trajectory following apexes of bends in molybdenite lamellae from sample RG10 show relatively constant, high Re concentrations (hundreds of ppm ) and constant Au concentrations (tens of ppm) (Figure 10a). In another case typifying the same type of texture (sample RG27), an inverse trend of variation is seen for the two elements, when comparing values from one spot to another i.e. consistently hundreds of ppm Re but an order of magnitude variation in Au (Figure 10b). In the same sample, in areas where there is replacement along molybdenite lamellae by chalcopyrite,
spots located on the apex of a kink show values of Re in the lower range for this sample (a few ppm) but Au values are in the upper range (tens of ppm) (Figure 10c).

An interesting case is seen in sample RG64, where coarse molybdenite coexists with packages of idiomorphic short molybdenite lamellae. In the former type, Re concentrations are lowest (a few ppm) but Au is highest (tens of ppm) (Figure 10d). In the latter type, however, the opposite is seen. Rhenium records the highest values (hundreds of ppm) but Au is at most a few ppm (Figure 10e). In the same sample, inhomogeneity throughout packages of short lamellae that display slight deformation show variation in Au content over 2 orders of magnitude, from few to a hundred ppm (Figure 10f). The highest concentration of Au in the diorite samples (Figure 10 g ) was recorded along grain boundaries where particle nucleation takes place, explaining the highly irregular Au signal across the time-resolved downhole spectra. In some cases, increases in Au concentration (up to tens of ppm) are also observed along the axes of kink folds (Figure 10h). In another sample (RG62), Au in the same textural context is only few ppm (RG56, Figure 10i), but increases to tens of ppm in the apexes of tight microfolds (Figure 10j).

Variability of Re and Au concentration depending upon comparable microscale textures is seen in molybdenite from the granite samples, even though the overall values are much lower than in the diorite. In the granite, trace element inhomogeneity in molybdenite is much more strongly expressed in terms of $\mathrm{Te}, \mathrm{Bi}, \mathrm{Pb}$ variation, and to some extent also of W , rather than Re and Au (Figure 10k).

## LA-ICP-MS MAPPING OF MOLYBDENITE

LA-ICP-MS element mapping was carried out on 4 samples, 3 from the diorite (RG27, RG10, NK25; Figure 11, A13, A14, respectively) and 1 from the granite (RG45; Figure 12), to better understand the observed correlation between textures, inclusion distribution and grain scale inhomogeneity obtained from spot analysis in the molybdenite.

Sample RG27 (Figure 11) represents the case of molybdenite in diorite where replacement of Cp and Clz takes place along the lamellar boundaries. An area with small kink deformation was also included. Native Bi is also observed as tiny inclusions. This is one of the diorite samples where wider variation was recorded for Re with upper values of several hundred ppm (Table A7b). The elemental maps show broadly similar distribution for Re and W with enrichment overlapping the Cp inclusions. Strong enrichment of Re is observed in the
kink area. Bismuth and Pb are higher than Te and seems to be concentrated within the lamellar rather than towards the Cp inclusions.

Sample RG10 (Figure A7) typifies molybdenite with consistently high Re concentration in the South Open Pit. Inclusions of Bi and Bi -tellurides are present. The map area is over a coarse lamella with a tight micro-fold on one side and silicate replacement on the other side. Rhenium shows high, pervasively concentration throughout the entire molybdenite. The micro-fold is marked by highest concentration whereas depletion in Re is observed through the middle of the lamella on the side with silicate replacement. Tungsten shows an inverse trend to Re with clear depletion in the coarser molybdenite lamella with highest concentration on the lamella boundaries as well as in the fold axis. Tellurium, Bi and Pb express very similar patterns with spots of highest concentration along the lamella in the micro-fold and also in the oblique alignment; such spots indicate sites for sub-micron particle nucleation. Gold and silver show very similar trends with the 3 chalcophile elements above even though at much lower concentration.

Sample NK25 (Figure A14) is representative for coarse-grained, highly deformed molybdenite with consistently high Re concentration in the diorite from the North Open Pit (ABreccia). Careful re-examination of the sample has shown the presence of minute inclusions of Au in the molybdenite. Rhenium and W show inverse patterns with trends of Re depletion and W concentration with the deformation. Bismuth and Pb are higher than Te and they broadly show a similar pattern with W , re-emphasising that deformation controls trace element distribution. Similar trends observed for Au and Ag even though in this case the concentration is much lower. It is interesting to note that Se shows comparable trends with Re.

Sample RG45 (Figure 12) is representative of the coarse molybdenite in the granite where the abundance of $\mathrm{Bi}-(\mathrm{Pb})$-tellurides is greatest. The selected area includes 2 sets of kinkbends on each side; replacement of silicates strongly marks the kink-bends. Slight deformation oblique to lamella orientation is also seen on the optical microscopic image (Figure 10k). Rhenium shows zonation with strong enrichment in the areas with the kinkbends relative to the middle part of the mapping area. In contrast W only shows strong enrichment along the kink bends and also along the oblique deformation. Lead, Bi and Te show high concentration with similar pattern to the W , strongly suggesting the relationship between textures and trace element concentration. Notably, Au show spot concentrations
overlapping the chalcophile elements enrichment despite the fact that the measured Au values in the sample are at best a few ppm.

## CRYSTAL-STRUCTURE AND NANOSCALE CHARACTERISATION OF MOLYBDENITE AND BI-(PB)-TELLURIDES (FIB-TEM DATA)

Molybdenite occurs as 2 different polytypes: hexagonal (2H) and rhombohedral (3R) of which the first type is the most common in nature. Newberry (1979a, b) stipulated that there should be a correlation between the crystal structure and the amount of minor/trace elements incorporated since 3 R molybdenite has a wider $c$ parameter than 2 H molybdenite (18 and 12 $\AA$, respectively). This assumption was based on the fact that many natural molybdenites appear to be rich in a range of trace elements including $\mathrm{Te}, \mathrm{Bi}, \mathrm{Pb}, \mathrm{Re}, \mathrm{W}$. For one such molybdenite, from Whipstick (NSW, Australia), Plimer (1974) concluded that the Mo-Au-Bi breccias in this deposit were zoned in terms of molybdenite polytypes, with the 2 H polytype in inner parts and the 3 R polytype dominating in the halo.

The petrographical, mineralogical and trace element distribution in molybdenite indicates that Boddington represents an ideal case to test Newberry's hypothesis. The high degree of inhomogeneity in terms of trace element distribution observed from both LA-ICP-MS spot analyses and element mapping suggests the widespread presence of sub-microscopic inclusions, including $\mathrm{Bi}-(\mathrm{Pb})$-tellurides, in all cases. High Au and Ag contents of both overlapping $\mathrm{Bi}-\mathrm{Pb}-\mathrm{Te}$ zones on the element maps and the strong correlation between the precious metals and these chalcophile elements suggest that underlying controlling factors operate at the sub-microscopic scale. Therefore, in order to understand the reasons for such correlation at the sub-microscopic scale, and whether this is controlled by crystal-structural modification in the host molybdenite, a TEM study that addresses these issues at the grainscale was considered necessary. This is now possible by using FIB-SEM methods for TEM sample preparation (Figure A15). The technique offers the advantage that TEM foils can be obtained in-situ, from specific sites where the sample shows compositional variation and the presence of inclusions (Ciobanu et al. 2011).

The main target was molybdenite from one granite sample (RG45) that contained abundant inclusions of phases from both the tetradymite group and aleksite series. A diorite sample (RG62) containing Bi-telluride inclusions within molybdenite was also studied. Four TEM foils were obtained from the molybdenite in the granite sample and one from the diorite
(Figure A16). High-resolution FIB imaging during slice cutting shows that sub-microscopic inclusions of tellurides are always present adjacent to the larger telluride inclusions; abundant kink-bending, high porosity and sub-micron inclusions were observed in the molybdenite (Figure 13a-c; Figure A16b-e, i, j, m, n).

Bright-field (BF) HR-TEM imaging shows that the telluride inclusions and porosity are present down to the nm-scale (Figure 13d-j). Lattice fringes with distortions closing in towards the pores are also a common feature (Figure 13k, l). Lattice fringes and electron diffraction patterns (EDPs) of the molybdenite from all 5 TEM foils indicate only the presence of the 2 H polytype (Figure 13 m and inset). EDPs obtained from boundaries between molybdenite and telluride inclusions show coherent intergrowths between the 2 minerals with preferential orientation along the $c^{*}$ axis (Figure 13d inset). Tiny inclusions of Gn were also identified from EDP (Figure 13g inset).

Bismuth-tellurides (-selenides, -tellurosulphides and -telluroselenides) comprising the homologous tetradymite group (Cook et al. 2007a, Ciobanu et al. 2009a) and the 'aleksite' series (Cook et al. 2007b) are mixed-layer compounds with rhombohedral or trigonal symmetry. They are derived from the same 5 -layer module ( $\mathrm{X}-\mathrm{Bi}-\mathrm{X}-\mathrm{Bi}-\mathrm{X} ; \mathrm{X}=\mathrm{Te}, \mathrm{Se}, \mathrm{S}$ ), known as the 'tetradymite archetype' ( $a=4.25 \AA ; c=29.6 \AA$ ), by incremental addition of Bi-Bi and $\mathrm{M}-\mathrm{X}(\mathrm{M}=\mathrm{Pb}, \mathrm{Bi})$, respectively.

Ciobanu et al. (2009a) show that all phases are N -fold superstructures ( $\mathrm{N}=$ total number of layers in the unit cell) of a rhombohedral subcell and that homology is underpinned by the structural formula $S^{\prime}\left(\operatorname{Bi}_{2 k} X_{3}\right) \cdot L^{\prime}\left(\operatorname{Bi}_{2(k+1)} X_{3}\right)$, where $X=$ chalcogen and $S^{\prime}$ and $L^{\prime}$ are the number of short and long modules, respectively. Phases from the aleksite series are isoconfigurational with those of the tetradymite group and have the general formula: $\mathrm{M}_{\mathrm{p}} \mathrm{X}_{\mathrm{p}+1}$; $\mathrm{p} \geq 2$; homology is given by the structural formula: $\mathrm{S}\left(\mathrm{M}_{1} \mathrm{X}_{m}\right) \cdot \mathrm{L}\left(\mathrm{M}_{m} X_{m}\right) ; \mathrm{S}, \mathrm{L}=$ numbers of short and long modules. The only named minerals in the aleksite series are: aleksite $\left(\mathrm{Bi}_{2} \mathrm{PbTe}_{2} \mathrm{~S}_{2}\right)$ and saddlebackite $\left(\mathrm{Bi}_{2} \mathrm{~Pb}_{2} \mathrm{Te}_{2} \mathrm{~S}_{3}\right.$; Clarke 1997), with $\mathrm{p}=3$ and 4, respectively. Crystal-structure in each of these minerals is built by a single type of layer, i.e. 7 -atom layer: $\mathrm{S}-\mathrm{Bi}-\mathrm{Te}-\mathrm{Pb}-\mathrm{Te}-$ Bi-S and 9-atom layer: S-Bi-Te-Pb-Te-Pb-S-Bi-S, respectively. Unnamed 'phase C' ( $\mathrm{Bi}_{4} \mathrm{PbTe}_{5} \mathrm{~S}_{2}$; Cook et al. 2007b) identified here by EPMA has a structure built by a stacking sequence where 5 - and 7 -atom layers alternate.

Homology implies that the mineral-chemistry and structure are related closely to one another so that the stacking sequences can be calculated from either the chemical composition or from electron diffraction patterns (Ciobanu et al. 2009a). The small grain size
and inherent fine-scale intergrown nature of minerals from homologous series hamper accurate EPMA measurements. Thus, assesment of stacking sequences using EDPs is necessary given the importance of the ' $k$ ' or ' p ' factors in establishing equilibrium in multiphase assemblages. At equilibrium k or p factors in the formulae of the component phases must be consecutive, e.g. joséite- $\mathrm{B}\left(\mathrm{Bi}_{4} \mathrm{Te}_{2} \mathrm{~S}\right)$ and hedleyite $\left(\mathrm{Bi}_{7} \mathrm{Te}_{3}\right)$, where $\mathrm{k}=2$ and 3, respectively, can coexist at equilibrium, whereas tetradymite $\left(\mathrm{Bi}_{2} \mathrm{Te}_{2} \mathrm{~S} ; \mathrm{k}=1\right)$ and hedleyite ( $\mathrm{k}=3$ ) cannot.

Based on empirical observations on a large number of occurrences, assemblages from the tetradymite group reflect the reduced/oxidised character of the mineral association (Ciobanu et al. 2005, Cook et al. 2009). Phases with $\mathrm{Bi}: \mathrm{X}>1$ indicate reduced conditions (pyrrhotiteor magnetite-stable) and coexist with native bismuth, maldonite. In contrast, phases with $\mathrm{Bi}: \mathrm{X}<1$ are found in oxidised associations (pyrite- or hematite-stable), coexisting with native tellurium and Au-tellurides. Thermodynamic modelling for the stability of native elements, Au - and main Fe -minerals in $f \mathrm{~S}_{2}-f \mathrm{O}_{2}$ space at $300^{\circ} \mathrm{C}$ also support the above (Cook et al. 2009).

Stacking sequences corresponding to Bi-tellurides present as inclusions in molybdenite were interpreted using EDPs following the method of Ciobanu et al. (2009a) (Figure 14). The Bi-telluride identified from the diorite is $\mathrm{Bi}_{4} \mathrm{X}_{3}$ (Figure 14a, h). The EDPs confirm the EPMA data for the dominant telluride in the granite as tsumoite; slight stacking disorder is indicated from both EDP's and BF HR-TEM images (Figure 14b, c). The EDPs obtained from foil 4 on an area with inclusions from the aleksite series indicate a higher-order phase. Stacking sequences interpreted from the EDPs indicate alternating 11- and 13-atom layers corresponding to composition $\mathrm{Bi}_{2} \mathrm{~Pb}_{7} \mathrm{Te}_{4} \mathrm{~S}_{9}$, an unnamed phase in the aleksite series (Figure $14 \mathrm{~d}-\mathrm{g}, \mathrm{j}$ ). The observed lattice fringes support this interpretation (Figure 14f). Disorder is observed on both EDPs and lattice fringes obtained for the unnamed phase (Figure 14e, g, j), as well as for tsumoite (Figure 14c). The domains of disorder observed in both cases comprise combination of layers with consecutive p or k , respectively, indicating crystallisation under equilibrium. This infers that the studied telluride inclusions embedded within molybdenite are not affected by an overprint inducing disequilibrium.

## DISCUSSION

## Previous work

The Boddington deposit was first considered as an Archean porphyry-style Cu-Mo deposit based on fluid inclusion evidence (high-salinity ore fluids; Roth \& Anderson 1993). This hypothesis was refuted by Cassidy et al. (1998) based on the lack of similarity between whole rock geochemistry of the diorite relative to diorites with associated porphyry-style mineralisation worldwide. An orogenic style of Au deposit is interpreted by Allibone et al. (1998) based on the strong structural control and the presence of high-grade Au in shear zones (e.g. ‘Jarrah’ bonanza ore along the South-East shear zone).

Stein et al. (2001) provide Re-Os dating of Boddington molybdenite and discuss a twostage model of Au mineralisation: (i) porphyry $\mathrm{Cu}-\mathrm{Au}-\mathrm{Mo}(2707 \pm 17 \mathrm{Ma}$ ), (ii) orogenic Au $(2623 \pm 9 \mathrm{Ma})$. Even though description of the samples giving the later date correspond to the granite as described here, i.e. "coarse-grained intervals of cloudy Qz, perthitic feldspar, Mu and Bt ", the age is attributed to an orogenic event based on both the low Re content and by similarity with orogenic Au mineralisation in the Yilgarn Craton. The error on the second age overlaps with SHRIMP U-Pb zircon dating of the granite ( $2611 \pm 3 \mathrm{Ma}$; Allibone et al. 1998) . Attribution of the first date to the porphyry event is favoured by the high Re-concentration, comparable with porphyry-style mineralisation elsewhere. McCuaig et al. (2001) chose to interpret the second Re-Os age in terms of an intrusion-related Au-Mo model rather than orogenic Au.

This intrusion-related Au model is reinforced by Duuring et al. (2007), based on the typical Bi-signature present at Boddington - a key factor in this class of deposits (e.g. Lang et al. 2000, Hart 2007). Hagemann et al. (2007) also support an intrusion-related Au model based on the high Bi and W measured in melt inclusions from granite. These authors consider the granite (with affinities to A- or fractionated I-type, unusual for Archean granitoids in the Yilgarn Craton) to be an important $\mathrm{Au}-\mathrm{Cu}-\mathrm{Mo}$ metal source and a good exploration model for the Boddington area.

Kalleske (2010) acknowledges the multi-stage character of mineralisation in which a preexisting $\mathrm{Cu}-\mathrm{Mo}-\mathrm{Au}$ mineralisation is attributed to mafic rocks (diorite) and an inferred porphyry-style mineralisation. The $\mathrm{Au}-\mathrm{Bi}-\mathrm{Te}$ association is recognised not only in mineralogical terms but also discussed in terms of overprinting during metamorphism and
granite emplacement The novelty of this work was that it raised the question of why molybdenite from the diorite and the granite have different trace element signatures (the granite is much poorer in $\operatorname{Re}$ and Au ) and whether these data could help validate a genetic model for Boddington. The Au-Bi-Te signature and reduced character of the ore assemblages led Kalleske (2010) to suggest that an intrusion-related model could not be ignored, despite the fact that only modest values of chalcophile elements were measured in molybdenite from granite (fine-grained aplitic variety).

The present study brings in a consistent package of new data concerning ore stages, formation temperatures, remobilisation and Au deposition in the Central Diorite together with new mineral chemistry data for molybdenite-dominant mineralisation within the granite.

## Stages of mineralisation

The main ore assemblages and alteration in diorite and granite share similarities but also differ in several aspects. Among the similarities are: (i) the ore is defined by the assemblage $\mathrm{Cp}+\mathrm{Po} \pm \mathrm{Mo} \pm \mathrm{Sph}$; (ii) $\mathrm{Bi}-(\mathrm{Pb})$-tellurides present in both molybdenite and altered country rock; (iii) a reduced character for the sulphide assemblage (Po stable, high Fe-content in Sph ) and telluride association ( $\mathrm{Bi} \geq \mathrm{Te}$ ); (iv) clinozoisite is the main alteration mineral associated with the ore. Notable differences are: (i) a generation of sulphides that clearly pre-date the granite (the difference in Po composition, Cub present in the diorite but not the granite; the bimodal distribution of Re and Au concentrations in molybdenite in diorite and the difference in the $\mathrm{Ni}-\mathrm{Co}$ signature of molybdenite from diorite and the granite); (ii) the absence of any visible Au in the granite; (iii) the low degree of alteration of the granite relative to that of the diorite; and (iv) that chamosite (Fe-rich chlorite) is present only in the granite.

This summary highlights that at Boddington there are at least 2 main ore stages (A) relating to host diorite and (B) formed during emplacement of granite. The similarity in the main sulphide and alteration assemblages, as well as the reducing character indicate that the metals and the fluids involved in both mineralising stages have either undergone a similar mechanism of ore deposition or that a strong superposition or obliteration occurred during granite emplacement.

## Metamorphism and hydrothermal alteration

Actinolite was recognised as an important component of alteration assemblages by Kalleske (2010). In the present study, however, it is scarce, a fact that can be at least partially attributed to replacement by Clz. In the Central Diorite and the granite there is a strong association between ore and Clz which could be indicative of redox reactions controlling metal precipitation during fluid-rock interaction. Such reactions would also favour crystallisation of molybdenite which is considered to be facilitated by reduction of $\mathrm{Mo}^{6+}$ (the species considered for hydrothermal transport) to $\mathrm{Mo}^{4+}$ (e.g. Stein et al. 2003).

Whereas Clz is part of the $\mathrm{Ab} / \mathrm{Olig}+\mathrm{Qz}+\mathrm{B} \pm \mathrm{Mu} / \mathrm{Chl}$ alteration compatible with low amphibolite facies metamorphism, there is evidence of superimposed greenschist facies $(\mathrm{Chl}+\mathrm{Mu}+\mathrm{Qz}+\mathrm{Ab})$ alteration in both diorite and the granite (e.g. Cham $+\mathrm{Mu} \rightarrow \mathrm{Bt}$ ). This is more pronounced towards depth and probably relates to the granite contact aureole. The lowest temperature type of alteration (zeolite facies, Stb+Qz+clay minerals) is attributable to localised action of faults. The presence of native copper instead of sulphides in samples from the open pit or the formation of $\mathrm{Bn}+\mathrm{Cov}$ in samples from the drillcores indicates various degrees of reducing conditions during the fault (re)activation.

The compositional data in this study allowed two independent geothermometers to be applied. The wide range of formation temperatures indicated by chlorite geothermometry and corresponding ( $6-\mathrm{Al}^{\mathrm{vi}}$ ) values can be clearly seen on Figure 2 h , reinforcing a model of regional metamorphism with pronounced retrograde overprint (e.g. in sample RG27). The temperatures obtained from sphalerite-stannite geohermometry (Figure 5d) overlap with upper temperature ranges obtained for chlorite from both diorite and granite. Considering the depth of the analysed samples with similar depth to the granite it is most likely that this temperature represents re-equilibration of $\mathrm{Sph}-\mathrm{Stn}$ in the thermal contact aureoles of the granite.

## Syn-metamorphic and late-granite overprinting

The ore textures, associations and trace element distribution in molybdenite are all highly suggestive of mineral re-equilibration and remobilisation during metamorphism and subsequent granite emplacement.

## LOCAL SULPHIDE REMOBILISATION

The relationships between Cu -Fe-sulphides ( Cp and Cub ) and Po and the inclusions they contain are indicative of localised variability in the composition of sulphide assemblages followed by different degrees of re-equilibration. For example, variable $\mathrm{Cu}: \mathrm{Fe}$ ratios within the Cu -Fe-sulphides would result in exsolution of Cub exsolving from Cp or Po from Cub. Minor contents of Zn and Sn in the Cu -Fe-sulphides would result in either exsolution of Sph Stn pairs from Cub or their accumulation at the boundaries between Cp and Po. The low Ni and Co contents determined in Po reflect release of these elements during metamorphism and subsequent formation of discrete Co-Ni-bearing minerals ( $\mathrm{Pn}, \mathrm{Mkw}$ and Thio). The diversity of Co-Ni-mineral species within the same sample, e.g. coexisting $\mathrm{Ag}-\mathrm{Pn}$ and Pn and from one sample to another e.g. differences in the thiospinel composition (see also Kalleske 2010) further emphasises the role of syn-metamorphic remobilisation in generation of the highly variable trace mineral signature seen throughout the deposit. The replacement of Po by Py, the abundance of REE-minerals observed in such samples (e.g. RG59), as well as the abundance of scheelite and W-REE-bearing minerals within $\mathrm{Fe}-\mathrm{Ti}$ oxides (e.g. RG52) are all indicative of a late-stage overprint due to interaction with fluids attributable to hydrothermal activity associated with granite emplacement.

The wide compositional range for Po (Figure 5a-c) and the textural relationships observed between the two pyrrhotite species in the diorite suggest the presence of different structural types, one of which is close to troilite, and the others, defined by different stoichiometries ( $\mathrm{Fe}_{7} \mathrm{~S}_{8}, \mathrm{Fe}_{9} \mathrm{~S}_{10}, \mathrm{Fe}_{10} \mathrm{~S}_{11}$ ), represent superstructures (Posfai \& Buseck 1997). Natural Po superstructures are commonly ferromagnetic. Troilite, however, is antiferromagnetic at temperatures below $140^{\circ} \mathrm{C}$ and 1 bar, but undergoes two magnetic phase transitions of which the $\beta$-phase transition is at the Curie temperature ( $315{ }^{\circ} \mathrm{C}$; e.g. Fleet 2006). The textural expression of lamellar exsolution of the two Po species (with pointed edges) suggests formation during syn-metamorphic deformation. Although the crystal structure of the $\sim \mathrm{FeS}$ phase would need to be determined, an analogy can be made with comparable lamellar exsolution between coexisting Po species at Challenger, S.A. (Haese 2010), one of which is troilite. As in the Challenger case, the coexisting Po types can be used to infer Po transition at the Curie point. Troilite is very rare in natural samples and is restricted to meteorites and liquid-magmatic Ni-Cu deposits (Becker et al. 2010). The possible occurrence at Boddington thus represents a second occurrence where troilite has formed in response to metamorphism.

## RHENIUM AND W BEHAVIOUR IN MOLYBDENITE

Previous LA-ICP-MS studies of molybdenite were focused on identifying sub-grain-scale decoupling of Re and Os to assess the validity of this method for dating (e.g. Stein et al. 2003, Kösler et al. 2003, Selby \& Creaser 2004). All these studies concluded that Re-Os decoupling readily takes place within molybdenite grains that undergo deformation. Rhenium mobility is discussed in terms of molybdenite crystallographic orientation but in general it is considered that Os is more mobile than Re , in particular under deformation conditions, and this explains the grain-scale decoupling observed. Other sulphides such as Cp accommodate much lower Re than molybdenite.

The trace element distributions and grain-scale inhomogeneity in molybdenite shown here are strong evidence for remobilisation of a whole range of elements, including Re, and has particular implication for Au (re)concentration. The distinctive flatness in the signal of both Re and W throughout the majority of the spectra clearly (Figure 6) indicates that these two elements are incorporated in the molybdenite lattice. This is commonly discussed in terms of the similarity between the ionic radii of $\mathrm{Mo}^{4+}(0.65 \AA), \mathrm{Re}^{4+}(0.63 \AA)$ and $\mathrm{W}^{4+}(0.66 \AA)$ (e.g. Stein et al. 2003). In the present dataset, both Re and W have wide variation from one sample to another and even within the same sample in the diorite and differ fundamentally from the very constant values measured in molybdenite from granite. This is a clear indication that both elements have undergone grain-scale remobilisation during deformation (kinks, folds and dislocation) and/or fluids introduced by the granite. There is, however, a marked difference in the patterns observed for Re and W on the element maps, where only 1 of the 4 maps show comparable patterns (RG27) for both elements. Tungsten concentration correlates with textures attributable to either deformation or replacement. In contrast, Re remains highly concentrated within molybdenite lamellae (RG10, 25 NK ) that, although strongly deformed, do not show too extensive replacement. Mobilisation of Re outside of the grain is seen in the case where there is evidence for fluid percolation, e.g. dense inclusions in the granite, porosity and replacement by either Cp or silicates.

McCandless et al. (1993) suggested that the presence of Re in molybdenite directly induces point defects and screw-dislocation growth due to the small difference in $\mathrm{Re}^{4+}$ and $\mathrm{Mo}^{4+}$ ionic radii. Such defects could assist Re loss during post-crystallisation alteration/deformation. The present study has shown that such defects, in particular kink bending extends to the lattice scale and these are also associated with nanoporosity (Figure
13). The element maps show that W is more mobile relative to Re . The latter can display stunning zonation patterns with discrete remobilisation along pathways of fluid percolation (Figure 12).

## MOLYBDENITE ROLE IN UPGRADING AU CONCENTRATION

The present study shows the consistent presence of Au and chalcophile elements $(\mathrm{Ag}, \mathrm{Te}, \mathrm{Bi}$, $\mathrm{Pb})$ throughout the molybdenite. All these elements are typically characterised by irregular signals on the LA-ICPMS spectra (Figure 6). This is concordant with the presence of abundant inclusions of tellurides in molybdenite from both granite and diorite, and also gold minerals in the diorite. The irregularity in the signal of these elements is much higher in the present dataset than in that of Kalleske (2010) implying orefield heterogeneity in terms of processes that control the concentration of these elements.

High correlation coefficients between Au and all chalcophile elements (Figure 9a-d) in the diorite indicate a common behaviour in response to processes that remobilise them. Similarly a strong association between chalcophile elements and Au is observed on the element maps, with clear indication of gold particle nucleation, even in the case of molybdenite from the granite. The presence of $\mathrm{Bi}-(\mathrm{Pb})$-tellurides as sub-microscopic lamellae adjacent to micronscale inclusions (Figures 13, A16) indicates coarsening of telluride lamellae from coherent lattice-scale intergrowths with molybdenite to visible inclusions. This hypothesis is reinforced by the presence of nanoporosity in the same areas. On the other hand, Au can be incorporated into the lattice of $\mathrm{Bi}-(\mathrm{Pb})$-tellurides (Ciobanu et al. 2009b) and this will be released as nanoparticles during overprinting. The presence of high Au anomalies overlapping the areas with $\mathrm{Bi}, \mathrm{Te}, \mathrm{Pb}$ can be attributed to such Au concentrations in the tellurides in molybdenite from the granite. Release of this Au to form discrete gold inclusions in the same areas with visible tellurides, or further away in the same sample, is characteristic for molybdenite in the diorite.

The present study shows no 2 H to 3 R transformation in molybdenite. This was the crystalstructural mechanism for high concentration of trace elements suggested by Newberry (1979a, b). It gives, however an alternative explanation- formation of lattice-scale intergrowths between molybdenite and phases with crystal-structural compatibility such as $\mathrm{Bi}-(\mathrm{Pb})$-tellurides. Considering the capacity of Bi-tellurides as Au carriers, molybdenite can
act as a true trap for Au and chalcophile elements if these elements are present in the same fluids from which molybdenite precipitated.

## LATE-STAGE GRANITE AND AU SCAVENGING BY BISMUTH MELTS

Gold is seen outside molybdenite in the same $\mathrm{Au}-\mathrm{Ag}-\mathrm{Bi}-\mathrm{Te}$ association in diorite from the contact aureole of the late-stage granite. Such associations are not locked within sulphides but rather tied to areas of chloritisation in the country rock. Temperatures of $235{ }^{\circ} \mathrm{C}$, inferred from the equilibrium assemblages in Mld-Hed-Bi droplets, overlap with the lower temperature range estimated from chlorite geothermometry. The appearance of such droplets is indicative of crystallisation from melts precipitated out of a fluid (Ciobanu et al. 2005). The most likely scenario is that redox reactions between the oxidised, granite-derived fluid and reduced, altered diorite country rock neutralise $\mathrm{Bi}^{3+}$ which in turn scavenges Au and other elements ( Te ) from the fluid. The $\mathrm{Bi}^{0}$ will precipitate as a composite $\mathrm{Au}-\mathrm{Bi}-\mathrm{Te}$ melt (instead of mineral crystallisation upon saturation) if the temperature of the redox reaction is above the melting temperature of the precipitated assemblage. This is feasible given the lowtemperature assemblages mentioned above.

Bismuth-melts can scavenge Au even from a fluid undersaturated in this element (e.g. Tooth et al. 2008, 2011) and will greatly improve the grade in a deposit with a protracted geological history. This interpretation is concordant with the fact that the granite-derived fluid is Bi-rich but does not seem to carry high Au concentrations (gold is absent in the granite-hosted mineralisation and granite-hosted molybdenite has the lowest Au content).

## Genetic model

The present study supports a multi-stage genetic model in which (i) an early porphyry-style mineralisation is overprinted by (ii) regional metamorphism and (iii) granite emplacement (Figure 16). Each of these events are characterised by metal input, the majority of Cu introduced during (i), Au in (ii) and $\mathrm{Bi}-\mathrm{Te}$ in (iii). Although primary textures from the earliest event are obliterated by metamorphism, the dominant Cu character and style of mineralisation (disseminations, patches, small veinlets) in the deposit are concordant with a porphyry $\mathrm{Cu}-\mathrm{Mo}-(\mathrm{Au})$ system.

Rhenium content in molybdenite is a good indicator of the mineralisation type (see above) and the present LA-ICPMS dataset is large enough to statistically characterise distinct populations of molybdenite. Despite grain-scale variability, the Re values in a given sample and also across sampling intervals cluster within a narrow range, a fact also supported by the element maps, thus proving the 'robustness' of Re relative to overprint. The three distinct populations of molybdenite can be attributed to: (i) porphyry-Moly (highest Re, mean 254 ppm ) (ii) orogenic-Moly (lowest Re, mean 1.76 ppm ) and (iii) granite-Moly (low Re, mean 3.94 ppm ). The first two are hosted in diorite and the last in granite. Although, low-Re Moly in the diorite could be considered input from the granite in the contact aureole, given the small difference in the Re concentration, the distinct Ni-Co signature of all Moly hosted by diorite (Figure 9e) shows that the three Moly generations are proven as discussed above.

The spatial distribution shows that the porphyry-Moly is dominant at shallower levels in both Central Diorite and ABreccia; orogenic Moly is at an intermediate level between the shallow levels and the granite (Figure 8a). The spatial overlap of these two types within the same sample or sampling interval (2.5m interval for RG27, 29, 30; WBD105900002 and RG64; WRD10600004) is observed in diorite located within the granite contact aureole; in the latter example re-crystallisation of porphyry Moly is also observed.

Higher Au values in molybdenite, both porphyry and orogenic, throughout the Central Diorite are the result of increase in Au concentration during both the metamorphic and granite emplacement events as discussed above. The distal location of ABreccia relative to the granite can thus explain the lower Au values in the molybdenite from this orebody where only metamorphism would have overprinted porphyry Moly. High-grade, shear-hosted Au mineralisation within the Eastern Shear Zone ('Jarrah'), as well as the majority of Au locked as inclusions within the sulphides, is attributable to the orogenic event.

The granite introduced chalcophile elements ( Bi and Te ) that greatly assisted concentration of Au in diorite-hosted Moly. The granite is clearly multi-phase where the finer-grained varieties are part of later aplitic suites and are depleted in such elements. Rhenium values are, however, comparable indicating the common magmatic reservoir for both coarse- and aplitic-granite. The granitic fluids are probably the lowest in Au but they have contributed to the overall Au balance in the deposit given the high Bi-Te content and melt-scavenging by Bi . The granite was probably emplaced during or shortly after peak metamorphism. The ages obtained by Allibone et al. (1998) and Stein et al. (2001) both carry
a sizeable analytical error. Based on the present observation the low-Re Moly in the diorite could be used for a more accurate dating of the orogenic event.

The strong $\mathrm{Bi}-\mathrm{Te}$ signature in Au mineralisation throughout the deposit, as well as molybdenite from the diorite, suggests that this is a characteristic that may have been inherited from the earlier porphyry- or orogenic-style systems rather than introduced only by the granite. It is interesting to note that even the bonanza ore has a marked Bi-Te signature (e.g. saddlebackite).

Although the intrusion-related model for the mineralisation has merit (high Bi-Te, reduced ore types, localised alteration), it cannot be proven in totality (low-Au fluids, lack of the Stype granite character). However, the overprinting effects introduced by granite emplacement (heat flow, acidic Bi-bearing fluids) greatly increased Au grades in the contact aureole. The intrusion-related exploration model is thus valid in the SGB. The recent discovery of the high-grade Katanning Au prospects in a granite aureole (Ausgold 2010), 200 km SE of Boddington supports this idea. The unusual characteristics of Boddington are however the result of the protracted geological history where gold- and sulphide-remobilisation greatly contributed to the giant size and economic Au grades.

The data presented here, complementing Kalleske (2010), strongly support the hypothesis that molybdenite trace element chemistry represents a valuable new metallogenetic and vectoring tool.

## CONCLUSIONS

- Molybdenite can incorporate not only Re but a whole range of other trace elements via structural defects and coherent lattice scale-intergrowths
- Concentrations of Re in molybdenite are highly variable on the deposit-scale. Economically significant concentrations of Re occur only at shallower levels in both Central Diorite and ABreccia; and not at deeper levels in the contact aureole of the granite.
- Grain-scale inhomogeneity in molybdenite is documented and can be related to the relative mobility of the different trace elements during overprinting.
- There is a marked association between gold and molybdenite, particularly when accompanied by intergrowths of Bi-tellurides, suggesting trapping of Au within Bi-
tellurides in both reworked early, high-Au molybdenite and later, Au-poor molybdenite.
- Available petrographic, mineralogical and geochemical evidence points to a threestage (porphyry-style, 'orogenic' and intrusion-related) genetic model for Boddington.
- Whereas the orogenic event introduces most of the Au , the granite emplacement upgrades the Au grades in the deposit by Bi -melt scavenging.
- Molybdenite trace element chemistry represents a valuable new metallogenetic and vectoring tool in comparable terranes.


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## FIGURE CAPTIONS

FIGURE 1. Geological sketch showing geology and main orebodies in the Open Pit at Boodington at level 150RL (modified from Kalleske 2010). Location of samples from drillcore and Open Pit area is indicated in red. Inset shows the location of the Boddington deposit in the Yilgarn Craton.

FIGURE 2. (a-e) Ternary plots showing the composition of silicates in the granite and diorite. (a, b) Ab-An-Kfs plots showing the compositional variation in feldspars. (c, d) XFe-XAl-XMn compositional plots for epidote-group minerals; note that granite (d) shows a larger compositional field but the range remains within Clz. (e) Ternary $\mathrm{Fe}-\mathrm{Mg}-\mathrm{Al}^{\text {vi }}$ plot showing the compositional range for chlorite in diorite (diamonds) and granite (circles). Note that Fe-richer varieties (end-member chamosite) are present in granite, whereas the Mg-richer chlorites are from diorite. (f-h) Binary plots of T vs. $\mathrm{Al}^{\mathrm{iv}}(\mathbf{f}, \mathbf{g})$ and vs. $6-\mathrm{Al}^{\mathrm{vi}}(\mathbf{h})$ in chlorite, showing the temperature ranges for distinct populations in diorite and granite.

FIGURE 3. Reflected light photomicrographs showing textures in molybdenite. (a-c) Coarse-grained molybdenite with typical deformation (a), kink-bending in marginal lamellae (b) and around silicates (c). (d) Atypical short-lamellar molybdenite forming an aggregate mesh. (e) Typical occurrence of coarser, micron-scale Bi-telluride inclusions adjacent to kink-bending. (f) Abundant, coarse chalcopyrite inclusions occurring in molybdenite from diorite; note the pointed edge termination of chalcopyrite lamellae. (g) Field of Bi-telluride inclusions coarsening across kinked lamellae in molybdenite from granite. (h) Bismuthtelluride segregation through axial planes of microfolds. (i) Remobilisation of Bi-telluride and chalcopyrite inclusions along microshears.

FIGURE 4. Back-scattered electron (BSE) images showing aspects of the $\mathrm{Au}-\mathrm{Ag}-\mathrm{Bi}-(\mathrm{Pb})-$ $\mathrm{Te}-(\mathrm{Se})$ association. ( $\mathbf{a}, \mathbf{b}$ ) Filaments of gold and electrum in molybdenite. (c) Composite gold and $\mathrm{Pb}-\mathrm{Bi}$-telluride inclusion in molybdenite. (d) Droplet of maldonite+hedleyite+bismuth in gangue. This association represents the equivalent of the eutectic at $235{ }^{\circ} \mathrm{C}$ in the system $\mathrm{Au}-\mathrm{Bi}-\mathrm{Te}$. (e) Cluster of Au-minerals and Bi-tellurides in areas of chloritisation along an $\mathrm{Act}+\mathrm{Clz}+\mathrm{Qz}+\mathrm{Bt}$ veinlet [the droplet in (d) is part of this area]. (f) Gold-bismuth assemblage in similar context as shown in (e). (g-i) Different contexts for hessite-Bi-tellurides $\pm$ bismuth inclusions: at the boundary between adularia and chalcopyrite in diorite (g); in gangue adjacent to chalcopyrite in granite (h) and within molybdenite from granite (i). (j) Typical fields of telluride inclusions along lamellae and kink-bends in
molybdenite. (k) Altaite associated with Bi-tellurides in molybdenite from diorite (l) Ikunolite as part of a complex telluride-bismuth assemblage within molybdenite from diorite.

FIGURE 5. (a-c) Binary plots of total metal (M) versus $S$ for pyrrhotite in granite and diorite. (a) Single-phase pyrrhotite corresponding to $\mathrm{Fe}_{7} \mathrm{~S}_{8}$ from granite. (b) Single-phase pyrrhotite corresponding with $\mathrm{Fe}_{7} \mathrm{~S}_{8}-\mathrm{Fe}_{9} \mathrm{~S}_{10}$ from diorite. (c) Distinct compositional fields for two-phase pyrrhotite occurring as lamellar exsolutions in diorite. The phase that is 'bright' on BSE images forms a field close to FeS (troilite) whereas composition of the 'dark' phase straddles across $\mathrm{Fe}_{10} \mathrm{~S}_{11}-\mathrm{Fe}_{7} \mathrm{~S}_{8}$. (d) Sphalerite-stannite geothermometry plot showing formation temperatures from $\log \mathrm{X}_{\mathrm{FeS}} / \mathrm{X}_{\mathrm{ZnS}}$ ratios in the two minerals. The cluster of points along the $350^{\circ} \mathrm{C}$ calibration line (Shimizu \& Shikazono 1985) is the most reliable.

FIGURE 6. (a-f) LA-ICPMS spectra for molybdenite from diorite (a-f) and granite (g, h). Note the difference between the flat signals for Re and W highly irregular signals for $\mathrm{Au}, \mathrm{Bi}$, Te and Pb across the time-resolved spectra.

FIGURE 7. Histograms showing the distribution of trace elements in molybdenite from granite (green) and diorite (red, purple and blue colours). (a) Rhenium in molybdenite from granite has a normal distribution whereas in molybdenite from diorite shows bimodal distribution. (b) In contrast, W in molybdenite from granite has bimodal distribution and normal distribution in molybdenite from diorite. (c) Gold shows normal distribution in molybdenite from granite (lower concentration) and bimodal distribution in molybdenite from diorite; note that majority of data plots at higher concentrations in the latter case. (d) Silver displays normal distribution for molybdenite from both granite and diorite. (e-g) Bismuth, Te and Pb show slight bimodal and normal distributions for molybdenite in granite and diorite respectively. The high-range corresponds to molybdenite from the coarse-grained granite, whereas the lower range is for the molybdenite in the aplitic variety.

FIGURE 8: LA-ICP-MS plots showing element concentration vs. depth in all analysed molybdenite. (a) Rhenium shows a decreasing trend and wider variation at depth. (b) Tungsten shows no variation or trend with depth. (c) Gold shows a wider variation and slight decreasing trend with depth. (d) Silver shows no trend but slightly wider variation at depth. (e) Tellurium shows marked increasing trend and wider variation at depth. (f, g) Bismuth and lead both display wider variation at depth where the highest and lowest concentrations are seen in molybdenite from coarse- and fine-grained granite, respectively.

FIGURE 9. LA-ICP-MS correlation plots between Au and chalcophile elements (a-d), NiCo (e) Re-W (f) and Au-Re (g) in molybdenite. Au and chalcophile elements show much higher correlation coefficients for molybdenite in diorite compared to molybdenite in granite (a-c). Gold and Ag have comparable correlation coefficients for molybdenite from diorite and granite (d). Very high correlation coefficient for Co-Ni in molybdenite from diorite compared to molybdenite from granite (e). (f) Rhenium and W show a funnel-shaped distribution. (g) Gold versus Re shows no correlation but it discriminates the four Re populations in molybdenite. The higher Au is clearly characteristic for molybdenite from diorite.

FIGURE 10. LA-ICP-MS spot analysis of trace element concentration in molybdenite from diorite ( $\mathbf{a}-\mathbf{j}$ ) and granite ( $\mathbf{k}$ ) showing grain scale inhomogeneity relative to textures. (a) Apex of bends in molybdenite lamellae from diorite showing both Re (hundreds of ppm) and Au (tens of ppm) are relatively constant. (b) Apex of bends in molybdenite lamellae from diorite showing constant Re values but variable Au (c) Replacement of molybdenite lamellae by chalcopyrite showing low Re and high Au values. (d) Coarse, deformed molybdenite from diorite showing low Re but higher Au values. (e) Idiomorphic, short molybdenite lamellae from the same diorite sample as in (d) showing high Re (hundreds of ppm ) but very low Au values. (f) Slightly deformed molybdenite lamellae from the same diorite sample as in (d) and (e) showing Au variation up to two orders of magnitude but constant high Re. The examples in d-f show coexistence of molybdenite from two generations: recrystallised porphyry in (e, f) and orogenic in (d). (g) Molybdenite from diorite showing the highest Au value measured suggesting particle nucleation along grain boundaries. (h) Axis of a kink folds in molybdenite from the diorite higher Au but low Re. (i) Similar texture as in (h) in molybdenite from diorite showing lower Au. (j) Apex of a tight microfold in molybdenite from diorite showing higher Au. Aspects of molybdenite in ( $\mathrm{g}-\mathrm{j}$ ) are characteristic for the orogenic event with low Re but highly variable Au content. (k) Large, coarse molybdenite aggregate in a granite sample showing microscale textures with large inhomogeneity in $\mathrm{Bi}, \mathrm{Pb}, \mathrm{Te}$ but relatively constant, low Re, W and Au. Rectangle in red shows the area for the LA-ICP-MS element map in Figure 12.

FIGURE 11. LA-ICP-MS element map for molybdenite in a diorite sample (RG27). The area is representative for trace element variation relative to replacement by Cp and Clz and also in kink bend. Both Re and W are concentrated towards the lamellae boundaries and within the kink. Bismuth and Pb show strong correlation and have similar patterns to Au and Ag.

FIGURE 12. LA-ICP-MS element map for molybdenite in a granite sample (RG45). The area shows kink bends and silicate replacement and minute $\mathrm{Bi}-(\mathrm{Pb})$-telluride inclusions. The chalcophile elements ( $\mathrm{Bi}, \mathrm{Pb}, \mathrm{Te}$ ) show strong correlation with each other; Au shows highconcentration spots overlapping enrichment in chalcophile elements. Re displays grain zonation where the core is relative free of replacement.

FIGURE 13. High-resolution FIB-SEM (secondary electron) images (a-c) and bright-field TEM images (d-m) of molybdenite from granite and diorite. (a-c) Sub-micron inclusions and fine particles (arrowed on c) of Bi-tellurides in molybdenite; note also the presence of galena in (a). (d-f) Nanometer-scale lamellar intergrowths between $\mathrm{Bi}-(\mathrm{Pb})$-tellurides and molybdenite. Note high-porosity and 'hair-pin' folds in molybdenite in (e). Selected area of electron diffraction (SAED) as insets in (d) and (f) show parallel orientation of molybdenite and $\mathrm{Bi}-\mathrm{Pb}$-tellurides along $\mathbf{c}^{*}$ and [1-10] zone in galena, respectively. ( $\mathbf{g}-\mathbf{j}$ ) Nanoporosity in molybdenite (arrowed) adjusted by bending (g), curvature (h), 'hair-pin’ folds (i) and twinning ( $\mathbf{j}$ ) in molybdenite. SAED (inset $\mathbf{j}$ ) shows twinning at $\sim 90^{\circ}$. ( $\left.\mathbf{k}-\mathbf{l}\right)$ Distortion of lattice fringes (arrowed). (m) Lattice fringes and SAED (inset) showing 2H molybdenite on [1100] zone; 3+1 integer ( $h \mathrm{klm}$ ) indexation is with respect to the hexagonal cell setting.

FIGURE 14. SAEDs, bright field TEM images and interpretation of stacking sequences in Bi- $(\mathrm{Pb})$-tellurides from the diorite and granite samples. (a, b) SAEDs showing [$110]_{\mathrm{R}}=[110 .]_{\mathrm{H}}$ zone axis in pilsenite (a) and tsumoite (b). Three-integer indexation, corresponding to the subscript R , is with respect to rhombohedral subcell and $(3+1)$ integer ( hklm ) indexation, whereas subscript H is with respect to the hexagonal cell setting. (c) Lattice fringes in tsumoite showing the typical 5- and 7-atom stacking sequence in which minor disorder is observed (' 55 ' layer repeats). (d-e) SAEDs showing the $[-110]_{\mathrm{R}}=[110 .]_{\mathrm{H}}$ zone axis in unnamed $\mathrm{Bi}_{4} \mathrm{~Pb}_{7} \mathrm{Te}_{4} \mathrm{~S}_{9}$. Stacking disorder is indicated by streaks and the pronounced ellipsoidal shape of reflections along $\mathbf{c}^{*}$ and parallel rows in (e). (f, g) Lattice fringes in unnamed $\mathrm{Bi}_{4} \mathrm{~Pb}_{7} \mathrm{Te}_{4} \mathrm{~S}_{9}$ showing ordered sequences (alternative 11- and 13-atom layers, f ) and disordered sequences (intervals of 13-atom layers, g ). ( $\mathbf{h - j}$ ) interpretation of the stacking sequences using the reflection distribution and their relative intensities in the characteristic ' d ' interval (Ciobanu et al. 2009a) corresponding to the rhombohedral sub-cell (marked on SAEDs in $a, b, d$, e). The number of divisions in the ' $d$ ' interval corresponds to the number of layers in the unit cell. The number of divisions within the central part of ' d ' (between the two second brightest reflections) corresponds to the number of layer-units defining the stacking sequence, e.g. 2 in both tsumoite and the unnamed phase (i, $\mathbf{j}$ ) but single
layer in pilsenite (h). In (j) the upper and lower ' d ' intervals show the difference between ordered and disordered state in terms of reflection morphology.

FIGURE 15. Schematic cartoon illustrating the three-stage mineralisation model for Boddington. (a) Porphyry $\mathrm{Cu}-\mathrm{Mo}-(\mathrm{Au})$ mineralisation formed during diorite intrusion within volcanic arc environment. High Re concentration in molybdenite is concordant with a mantle source. (b) Orogenic Au-style mineralisation emplaced during shearing and syn-metamorphic deformation at ca. 2675 Ma synchronous with the first (Archean) dyke emplacement (not shown on figure) and sulphide remobilisation. Molybdenite is characterised by low Re concentrations ( $<1 \mathrm{ppm}$ ) but a $\mathrm{Co}-\mathrm{Ni}$ signature, common also to the porphyry-style molybdenite. (c) Intrusion-related Mo-Bi-Te-(Au) mineralisation associated with granite emplacement. Molybdenite is characterised by low Re concentration (a few ppm). Aplitic granite present at depth (yellow crosses) has the same Re concentration but much lower Bi and Te . Molybdenite lacks a Co-Ni signature. Further remobilisation of metals and sulphide re-equilibration in the contact aureole and Au scavenging by Bi -melts is interpreted as the result of interaction between high-chalcophile element but low-Au granitic fluids. (d) Present-day exposure level of the Boddington mineralisation. Abundant Proterozoic dyke swarms are not shown for simplicity.

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## Figure 2



Figure 3


Figure 4


Figure 5


Po granite single
phase (42RG, 46, 48)
Diorite

- Po single phase
(25RG, 27, 59, 66)
Po two-phases
(11RG, 24, 36, 52)
$\square$ 'bright'
- 'dark'





## Figure 6



Figure 7


$\mathrm{N}=250$ mean $=2.187697$ std dev $=.8603354$


Figure 8


Figure 9



Re 368
Au 11

b


Re 3.1, W 32, Au 0.01, Bi 4911 Te 3037

Re 0.09, W 30, Au 0.04, Bi 1540, Te 962

Re 2.0, W 27, Au 0.05, Bi 1865 , Te 1188

Re 1.8, W 32, Au 0.07, Bi 1210, Te 785


Re 3.1, W 39, Au 0.09, Bi 73, Te 64

Re 2.5, W 37,
Au 0.18, Bi 22156, Te 13832

Re 4.9, W 25,
Au 0.4, Bi 23342, Te 4902

Re 2.0, W 27,
Au 0.5, Bi 1865, Te 1188



RG45

Figure 12


Re185_Cps





W182_cps

Au197_cps



Figure 13


Figure 14


7 divisions;
stacking sequence 7 Bi4X3; X=Te, S, Se; if $\mathrm{X}=\mathrm{Te}$-->pilsenite


12 divisions;
stacking sequence 5.7 $\mathrm{Bi} 2 \mathrm{Te} 3+\mathrm{Bi} 4 \mathrm{Te} 3=\mathrm{Bi} 6 T \mathrm{TE} 6$, tsumoite

24 divisions;
stacking sequence 11.13 Bi2Pb3Te2S4+Bi2Pb4Te2S5 =Bi4Pb7Te4S9 (unnamed)


Figure 15


Table 1. Sample description

| Sample description | meter | Sample <br> ID | Ore minerals |  | Gangue/Alteration |  | SEM | EPMA | LAICPMS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Major\&Minor sulphides/ore min. | Au-Ag-Bi-Pb-Te-Se association | Major | Others |  |  |  |
|  |  | OPEN PIT |  |  |  |  |  |  |  |
| Veinlet Moly in porph. diorite | RL108 |  | Moly, Cp, Po |  | Qz, Ab/Plag, Ep, Bt, Mu | Tit, REE-min | x |  | x |
| Patch Moly in bleached diorite | RL108 | RG10 | Moly | Bi, Bi-tell | Qz, Ep/Clz, Ab, Plag | Cal, Clay min | x | X | x |
| Sulphide dissem. in diorite | RL108 | RG11 | Cp, Po, Sph | $\mathrm{Hs}, \mathrm{Bi}, \mathrm{Ag}$ ? | Qz, Ep, Ab, KFs, Bt, Chl | Tit, Calc, Sch | x | X |  |
| Sulphide vein in diorite | RL108 | RG12 | Cp , Po |  | Plag, Qz, Bt, Ep | Tit, Ap | x |  |  |
| Large sulph. veinlet in diorite | RL108 | RG13 | Cp , Po | Hs | Qz, Bt, Ep, Plag |  | x |  |  |
| X -sulph. veinlets in bleached diorite | RL108 | RG14 | $\mathrm{Cp}, \mathrm{Po}$ |  | Plag, Qz, Ep , Bt, Mu | Ap, Tit | x |  |  |
| Sulphide veinlet in diorite | RL108 | RG16 | $\mathrm{Cp}, \mathrm{Po}$ |  | Qz, Plag, Bt, Mu, Ep |  |  |  |  |
| Sulphide veinlet in diorite | RL108 | RG17 | Cp, Py | $\mathrm{Bi}, \mathrm{Cls}$ | Qz, Bt, Plag, Ep | Bast, Zr, Tit | x |  |  |
| Diorite wt large sulphide patches | RL108 | RG18 | $\mathrm{Py}, \mathrm{Cp}$ |  | Qz, Plag, Bt, Ep |  |  |  |  |
| Dolerite | RL108 | RG19 | Py, Cp |  | Hbn/Act, Ep, Qz,Plag, <KFs | Ilm, Ru, Tit, Ap | x |  |  |
| Altered diorite wt Cu staining | RL108 | RG67 | $\mathrm{Cu}, \mathrm{Cc}$ |  | Qz, Plag/Ab, Kfs, Bt, Mu, Ep, Aln, clay min. | Stb, Thor | x |  |  |
|  |  |  | WBD 10630001; RL232.59; 9986.19 | E/ 10628.24 BGM N |  |  |  |  |  |
| Fine-gr. diorite, patch/dissem. Cp | 252.5 | RG21 | Cp, Po, Cub, Sph |  | Qz, Bt, Plag/Ab, Ep , Chl | Tit |  |  |  |
| Fine-grained diorite wt sulph. dissem. | 254.9 | RG23 | $\mathrm{Cp}, \mathrm{Po}$ |  | Qz, Bt, Plag, Ep |  |  |  |  |
| Fine-grained diorite wt sulph. dissem. | 255.9 | RG24 | Cub, Cp, Po, Pn/Ag-Pn, Sph | $\mathrm{Bi}, \mathrm{Cls}$ | Bt, Ab/Plag, Qz, Ep | Tit, Zr | x | X |  |
| Felsic patch wt. sulph/Moly in fine-grained diorite | 256.4 | RG25 | Moly, Cp, Po/Lam, Cub, Bn, Cov | Bi-tell,Pb-Bi-tell, Alt, Bi, lk, El, Te?, Naum | Ab/Plag, Qz, Bt, Ep/Clz | Tit | x | Xx | x |
| Veinlet in diorite | 258.2 | RG26 | Cp, Po |  | Ab/Plag, Qz |  |  |  |  |
|  |  |  | BD 105900002; RL 241.46; 10022.95 | M E/ 10591.69 BGM N |  |  |  |  |  |
| Diorite | 339.8 | RG27 | Moly, Cp, Po, Cub, Sph, Stn, Thio? | Bi | Bt, Qz, Plag, Mu, Chl | FI, Tit, Stb, Carb | x | Xx | x |
| Felsic band in diorite | 340 | RG28 | Po, Cub, Moly, Sph | Bi-tell? | Qz, Ab, Ep/Clz, Bt |  |  |  |  |
| Veinlet/diorite | 342.1 | RG29 | Moly |  | $\mathrm{Qz}, \mathrm{Plg} / \mathrm{Ab}, \mathrm{Bt}, \mathrm{Ep}, \mathrm{Mu}$ | Tit | x | x | x |
| Veinlet/diorite | 342.5 | RG30 | Moly, Cp | Bi-tell | Qz, Ab/Plag, Ep, Bt, Mu | Tit, FI, Bi | x |  | x |
| X -cutting veinlets in mylonite | 344.9 | RG31 | Cp, Moly, Po, Sph, Mrs? |  | $\mathrm{Qz}, \mathrm{Bt}, \mathrm{Mu}, \mathrm{Ab}, \mathrm{Ep} / \mathrm{Clz}$ | Tit |  |  |  |
| Veinlet in sheared diorite | 346.6 | RG32 | Moly |  | $\mathrm{Plg} / \mathrm{Ab}, \mathrm{Qz}, \mathrm{Bt}, \mathrm{Chl}$ | Tit |  | x |  |
| Felsic zone/diorite | 353 | RG33 | Po, Cp, Cub, Sph, Py |  | Qz, Ep, Ab/Plag, Bt |  |  |  |  |
| X -cutting veinlets in diorite | 353.5 | RG34 | Cp, Cub, Py |  | Qz, Bt, Ep |  |  |  |  |
| Diorite wt sulphide dissem. | 354.2 | RG35 | Cp, Po |  | Bt, Ab/Plag, Ep, Qz |  |  |  |  |
| Diorite | 359.5 | RG36 | Cp, Po, Thio? |  | $\mathrm{Qz}, \mathrm{Ab}, \mathrm{Bt}, \mathrm{Mu}, \mathrm{Ep} / \mathrm{Clz}, \mathrm{Chl}$ | Tit, Sch | x | Xx |  |

$x=E P M A$ silicate; $\mathrm{X}=\mathrm{EPMA}$ ore min.; *FIB-TEM
Coordinates given in Boddington Gold Mine Coordinates
$R L=$ Relative level

| Table 1 continued Sample description | meter | Sample ID | Ore minerals |  | Gangue/Alteration |  | SEM | EPMA | LAICPMS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Major\&Minor sulphides/ore min. | $\mathrm{Au}-\mathrm{Ag}$-Bi-Pb-Te-Se association | Major | Others |  |  |  |
|  |  | WRD 09775002; RL 278.483; 10560.23 BGM E/ 9763.313 BGM N |  |  |  |  |  |  |  |
| Diorite/veinlet | 168.2 | RG38 | $\mathrm{Cp}, \mathrm{Po}, \mathrm{Sph}$, Thio? |  | Bt, Qz, Plag, Ep/Clz | Tit, Cal |  |  |  |
| Moly veinlet in granite | 402.9 | RG40 | Moly, Cp |  | Qz, KFs/Pert, Ab, Clz, Bt, Mu |  |  |  | x |
| Vein sulphides in granite | 410.2 | RG42 | Po, Cp, Moly, Sph | Hed, Pls?, Bi, Hs, Gn | Qz, Pert, Ab, Mu, Clz, Chl | Sch, Nbts | x | Xx |  |
| Moly/Granite | 411.3 | RG43 | Moly, Po, Cp | Ts, Gn | $\mathrm{Qz}, \mathrm{Ab}, \mathrm{Kfs} / \mathrm{Pert}, \mathrm{Bt}, \mathrm{Mu}$, $\mathrm{Clz} / \mathrm{Ep}, \mathrm{Aln}, \mathrm{Chl}$ | REE-min | x | Xx | x |
| Granite wt patch Moly | 412.6 | RG44 | Moly, Cp |  | Qz, Pert, Plag, Bt, Mu, Ep |  |  |  | x |
| Granite wt large Moly | 416.5 | RG45 | Moly, Cp | Ts, Pb-Bi-tell, Gn, Hs | Ab, Qz | FI | x | $x^{*} x$ | x |
| Large sulph. patch/dissem. in granite | 419.7 | RG46 | Po, Cp | Gn | Qz, Ep, Pert, Plag, Bt | Tit |  | X |  |
| Fluorite vein granite | 448.1 | RG47 | Cp, Po |  | Qz,Pert, Ab, Bt, Mu, Clz, Chl | FI |  | x |  |
| Granite/ Veinlet | 448.4 | RG48 | $\mathrm{Cp}, \mathrm{Po}$, Sph, Thio? |  | $\mathrm{Qz}, \mathrm{Ep} / \mathrm{Clz}, \mathrm{Mu}, \mathrm{Chl}$ |  |  | X |  |
| Fine-grained granite wt patch Moly | 477.3 | RG49 | Moly |  | Qz, Pert, Ab, Bt, Mu, Ep, Chl | Tit |  |  | x |
| Intrusive dyke | 491.9 | RG50 | Moly, Po, Cub, Sph |  | $\mathrm{Qz}, \mathrm{Kfs}, \mathrm{Plag} / \mathrm{Ab}, \mathrm{Bt}, \mathrm{Clz}, \mathrm{Chl}$ | Tit, Carb |  | x |  |
| Fine-grained granite wt patch Moly | 500.3 | RG51 | Moly |  | $\mathrm{Qz}, \mathrm{Plg} / \mathrm{Ab}, \mathrm{Ep}, \mathrm{Bt}, \mathrm{Mu}$ |  | x |  | x |
|  |  |  | RD 10600004; RL 278.483; 9996.546 | E/ 10602.093 BGM N |  |  |  |  |  |
| Veinlet in diorite | 217.5 | RG52 | $\mathrm{Cp}, \mathrm{Po}, \mathrm{Py}$ ?, Sph | MId, Au, Bi, Hed, Bi8Te3, Hs | Ab, Plag, Ad, Qz, Act, Ep, $\mathrm{Bt}, \mathrm{Mu}$ | Tit, Ilm, Ru, Ap, Sch, Nbts | x | X |  |
| Veinlet in diorite | 218.2 | RG53 | $\mathrm{Cp}, \mathrm{Po}$ |  | Ab, Qz, Act |  |  |  |  |
| Veinlet in diorite | 247.4 | RG54 | Cp, Po, Moly, Sph, Pn |  | Qz, Plag, Ep/Clz | Tit |  |  |  |
| Veinlet (in diorite) | 263.6 | RG55 | Moly, Cp |  | Qz, Clz/Ep, Ab |  |  |  | x |
| Felsic zone in diorite | 267.1 | RG56 | Moly, Cp, Po |  | Qz, Plag, Bt, Ep/Clz |  |  |  | x |
| Sulphide veinlet\&dissem. In diorite | 276.6 | RG57 | $\mathrm{Cp}, \mathrm{Po}$ |  | Qz, Bt, Plag, Ep/Clz, Chl |  |  |  |  |
| Sulphide patch in felsic zone (diorite) | 277.6 | RG58 | Cp, Po, Sph, Py |  | Qz, Plag, Ep/Clz, Bt | Tit |  |  |  |
| Coarse diorite | 280 | RG59 | Cub, Po, Py, Sph, Stn, Cp, Thio, Pn | $\mathrm{Bi}, \mathrm{Hs}$, Ltk? | Qz, Plag, Bt, Ep/Clz, Chl | Ap, Sch, Bast | x | X |  |
| Veinlet/diorite | 285.15 | RG60 | Moly, Cp, Py, Sph | $\mathrm{El}, \mathrm{Au}, \mathrm{Bi}$ | $\mathrm{Qz}, \mathrm{Ab}, \mathrm{Clz} / \mathrm{Ep}, \mathrm{Bt}, \mathrm{Mu}$ | Cal, Stb | x |  | x |
| Sulph. veinlet in porphyrobl. diorite | 318.6 | RG61 | $\mathrm{Cp}, \mathrm{Po}$, Sph, Thio? |  | Qz,Ab, Ep/Clz |  |  |  |  |
| Coarse Plag.-felsic band /diorite | 321.5 | RG62 | Moly, Cp, Po | Bi-tell (Bi4Te3, Bi2Te), Au, El | Qz, Plag/Ab, Bt, Ep, Mu, Chl | Ilm, Tit, REE-min | x | x* | x |
| Veinlet in diorite | 323.5 | RG63 | Cp, Po |  |  |  |  |  |  |
| Fine-grained diorite | 371.3 | RG64 | Moly, Cp, Sph |  | Qz, Bt, Plag, Ep/Clz | Tit |  |  | x |
| Veinlet/diorite | 379.3 | RG65 | Moly, Cp, Po, Cub |  | Qz, Bt, Plag/Ab, Clz |  |  |  | x |
| Bleached diorite | 924.25 | RG66 | Cp, Po/Lam, Sph | $\mathrm{Bi}, \mathrm{Hs}, \mathrm{Gn}$, Vol? | Bt, Mu, Ab/Plag, Qz, Ep, Chl | Tit | $x$ | X |  |

[^0]Coordinates given in Boddington Gold Mine Coordinates

Table 2. Mineral Abbreviations

| Mineral | Abbreviation |
| :---: | :---: |
| Actinolite | Act |
| Adularia | Ad |
| Albite | Ab |
| Allanite | Aln |
| Altaite | Alt |
| Apatite | Ap |
| Bastnasite | Bast |
| Biotite | Bt |
| Bismuth tellurides | Bi-tell |
| Bornite | Bn |
| Calcite | Cal |
| Calcium niobate | $\mathrm{Nb}-\mathrm{Ca}$ |
| Carbonate | Carb |
| Chalcocite | Cc |
| Chalcopyrite | Cp |
| Chamosite | Cham |
| Clausthalite | Cls |
| Clinozoisite | Clz |
| Covellite | Co |
| Cubanite | Cub |
| Electrum | El |
| Epidote | Ep |
| Fergusonite | Ferg-(Y-Niobate) |
| Flourite | Fl |
| Galena | Gn |
| Gold | Au |
| Hedleyite | Hed |
| Hessite | Hs |
| Hornblende | Hbn |
| Ikunolite | Ik |
| Ilmenite | 11 m |
| K-feldspar | KFs |
| Laitakarite | Ltk |
| Maldonite | Mld |
| Marcasite | Mrs |
| Molybdenite | Moly |
| Muscovite | Mu |
| Naummanite | Naum |
| Niobates | Nbts |
| Native bismuth | Bi |
| Native copper | Cu |
| Native silver | Ag |
| Oligoclase | OI |
| Pentlandite | Pn |
| Perthitic feldspar | KFs-Pe |
| Pilsenite | Pls |
| Plagiocalse | Plag |
| Pyrite | Py |
| Pyrrhotite | Po |
| Quartz | Qz |
| Rare Earth Minerals | REE-min |
| Rutile | Ru |
| Scheelite | Sch |
| Sphalerite | Sph |
| Stannite | Stn |
| Stilbite | Stb |
| Thiospinel | Thio |
| Thorite | Tho |
| Titanite | Tit |
| Tsumoite | Ts |
| Volynskite | Vol |
| Zircon | Zr |

Table 3. Microprobe analyses of chlorite group minerals


## Table 3 continued.

| Ca | 0.002 | 0.002 |
| :--- | :---: | :---: |
| Na | - |  |
| K | 0.042 | 0.050 |
| Total | 5.886 |  |
|  |  |  |
| F | 0.120 | 0.010 |
| Cl | 0.004 | 0.005 |
| (OH) | 15.876 | 0.008 |
| total | 16.000 |  |
|  |  |  |
| Fe/(Fe+Mg+Mn) | 0.95 | 0.00 |
|  |  |  |
| \% clinochlore | 3.8 | 0.1 |
| \% chamosite | 94.9 | 0.3 |
| \% pennantite | 1.3 | 0.2 |
|  |  |  |
| Geothermometric calculations | 366 | 20 |
| Cath 1988 | 386 | 20 |
| J 1991 |  |  |


| 0.003 | 0.002 | 0.069 | 0.001 | 0.002 | 0.010 | 0.008 | 0.090 | 0.009 | 0.004 | 0.003 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.027 | 0.053 | 0.043 | - | 0.012 | - | 0.002 | 0.020 | 0.010 | - | 0.006 |
| 0.005 | 0.002 | 0.015 | 0.001 | 0.003 | 0.002 | 0.001 | 0.012 | 0.004 | 0.004 | 0.002 |
| 5.950 |  | 5.786 | 6.100 | 5.963 | 5.983 |  | 5.806 |  | 5.960 |  |
| 0.110 | 0.008 | 0.104 | 0.112 | 0.128 | 0.156 | 0.017 | 0.114 | 0.017 | 0.096 | 0.018 |
| - |  | 0.004 | 0.002 | 0.001 | 0.004 | 0.004 | 0.004 | 0.002 | 0.003 | 0.003 |
| 15.888 | 0.007 | 15.892 | 15.885 | 15.871 | 15.840 | 0.020 | 15.882 | 0.014 | 15.901 | 0.017 |
| 16.000 |  | 16.000 | 16.000 | 16.000 | 16.000 |  | 16.000 |  | 16.000 |  |
| 0.71 | 0.03 | 0.82 | 0.50 | 0.38 | 0.64 | 0.01 | 0.78 | 0.03 | 0.51 | 0.006 |
|  |  |  |  |  |  |  |  |  |  |  |
| 28.0 | 3.4 | 17.8 | 49.6 | 60.9 | 35.1 | 0.6 | 20.0 | 3.2 | 47.7 | 0.6 |
| 70.9 | 3.4 | 81.7 | 49.7 | 38.5 | 63.5 | 0.6 | 78.5 | 3.4 | 51.4 | 0.6 |
| 1.1 | 0.1 | 0.5 | 0.7 | 0.6 | 1.4 | 0.2 | 1.5 | 0.2 | 0.9 | 0.1 |
|  |  |  |  |  |  |  |  |  |  |  |
| 350 | 28 | 194 | 416 | 300 | 367 | 11 | 219 | 21 | 362 | 8 |
| 362 | 29 | 211 | 421 | $\mathbf{3 0 2}$ | 377 | 11 | 235 | 21 | 367 | 8 |

using formulae from Caritat et al. (1993) Cath = Cathelineau (1988)
J = Jowett (1991) - applies when $\mathrm{Fe} /(\mathrm{Mg}+\mathrm{Fe})<0.6$
Bolding on geothermometric calculated values signifies chosen calibrations

Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, F 0.17, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, Na 0.04, P 0.03, Si 0.02, Ti 0.02

Table 4. Electron probe analyses of pyrrhotite

|  | RG52 |  |  |  | RG24 |  |  | RG11 |  |  |  |  |  | RG25 |  | RG27 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | mean | SD | mean | SD | mean | SD |  | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD |
|  |  |  | LIGHT $\mathrm{n}=11$ |  | DARK $\mathrm{n}=3$ |  | LIGHT | DARK $\mathrm{n}=17$ |  | LIGHT $\mathrm{n}=15$ |  | no lam. $\mathrm{N}=3$ |  | no lam n=4 |  | no lam $\mathrm{n}=15$ |  |
| Cu | 0.07 | 0.06 | 0.08 | 0.10 | 0.21 | 0.14 | 0.00 | 0.06 | 0.04 | 0.05 | 0.05 | 0.05 | 0.01 | 0.22 | 0.09 | 0.07 | 0.05 |
| Fe | 59.41 | 0.38 | 61.84 | 0.69 | 59.34 | 0.40 | 61.45 | 60.08 | 0.57 | 61.95 | 0.59 | 61.79 | 0.17 | 59.61 | 0.49 | 58.87 | 1.62 |
| Co | 0.03 | 0.03 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.01 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 |
| Ni | 0.20 | 0.12 | 0.03 | 0.07 | 0.02 | 0.02 | 0.00 | 0.19 | 0.05 | 0.03 | 0.08 | 0.03 | 0.01 | 0.03 | 0.01 | 0.09 | 0.09 |
| S | 38.17 | 0.26 | 36.51 | 0.85 | 38.68 | 0.38 | 36.84 | 38.42 | 0.23 | 36.60 | 0.55 | 36.26 | 0.16 | 38.88 | 0.27 | 38.52 | 1.61 |
| Total | 97.87 |  | 98.46 |  | 98.25 |  | 98.29 | 98.76 |  | 98.64 |  | 98.13 |  | 98.75 |  | 97.55 |  |
| Formulae (2 a.p.f.u.) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Cu | 0.001 | 0.001 | 0.001 | 0.001 | 0.003 | 0.002 | 0.000 | 0.001 | 0.000 | 0.001 | 0.001 | 0.001 | 0.000 | 0.003 | 0.001 | 0.001 | 0.001 |
| Fe | 0.941 | 0.006 | 0.985 | 0.017 | 0.934 | 0.006 | 0.978 | 0.944 | 0.004 | 0.985 | 0.012 | 0.988 | 0.001 | 0.934 | 0.006 | 0.933 | 0.031 |
| Co | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Ni | 0.003 | 0.002 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.003 | 0.001 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.001 |
| Total M | 0.947 |  | 0.987 |  | 0.939 |  | 0.978 | 0.948 |  | 0.987 |  | 0.990 |  | 0.938 |  | 0.937 |  |
| S | 1.053 | 0.006 | 1.013 | 0.017 | 1.061 | 0.005 | 1.021 | 1.051 | 0.005 | 1.013 | 0.011 | 1.010 | 0.001 | 1.061 | 0.008 | 1.063 | 0.030 |
| M/S | 0.90 | 0.01 | 0.97 | 0.03 | 0.88 | 0.01 | 0.96 | 0.90 | 0.01 | 0.97 | 0.02 | 0.98 | 0.00 | 0.88 | 0.01 | 0.88 | 0.05 |


|  | RG36 |  |  |  | RG59 |  | RG66 |  | RG42 |  | RG46 |  | RG48 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD |
|  | DARK $\mathrm{n}=24$ |  | LIGHT $\mathrm{n}=7$ |  | no lam $\mathrm{n}=2$ |  | no lam $\mathrm{n}=14$ |  | no lam $\mathrm{n}=20$ |  | no lam $\mathrm{n}=18$ |  | no lam n=37 |  |
| Cu | 0.10 | 0.08 | 0.25 | 0.20 | 0.06 | 0.03 | 0.14 | 0.10 | 0.07 | 0.05 | 0.01 | 0.02 | 0.03 | 0.03 |
| Fe | 59.89 | 0.45 | 61.91 | 0.56 | 59.18 | 0.76 | 59.66 | 0.32 | 59.24 | 0.73 | 59.21 | 0.48 | 59.33 | 0.60 |
| Co | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.00 | 0.04 | 0.02 | 0.00 | 0.00 | 0.02 | 0.01 | 0.05 | 0.01 |
| Ni | 0.05 | 0.05 | 0.01 | 0.01 | 0.01 | 0.01 | 0.03 | 0.02 | 0.01 | 0.01 | 0.02 | 0.01 | 0.01 | 0.01 |
| S | 38.60 | 0.28 | 36.67 | 0.49 | 39.05 | 0.75 | 38.28 | 0.20 | 38.93 | 0.33 | 38.57 | 0.16 | 38.60 | 0.80 |
| Total | 98.64 |  | 98.84 |  | 98.31 |  | 98.15 |  | 98.25 |  | 97.82 |  | 98.02 |  |
| Formulae (2 a.p.f.u.) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Cu | 0.001 | 0.001 | 0.004 | 0.003 | 0.001 | 0.000 | 0.002 | 0.001 | 0.001 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 |
| Fe | 0.941 | 0.005 | 0.982 | 0.011 | 0.929 | 0.016 | 0.943 | 0.004 | 0.931 | 0.008 | 0.936 | 0.005 | 0.936 | 0.012 |
| Co | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.001 | 0.000 |
| Ni | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 | 0.000 |
| Total M | 0.943 |  | 0.986 |  | 0.931 |  | 0.946 |  | 0.933 |  | 0.938 |  | 0.939 |  |
| S | 1.057 | 0.006 | 1.013 | 0.011 | 1.068 | 0.016 | 1.053 | 0.004 | 1.066 | 0.008 | 1.062 | 0.005 | 1.061 | 0.012 |
| M/S | 0.89 | 0.01 | 0.97 | 0.02 | 0.87 | 0.03 | 0.90 | 0.01 | 0.88 | 0.01 | 0.88 | 0.01 | 0.88 | 0.02 |

Detection limits (wt\%) Cu 0.06, Co 0.04, Ni 0.03, Fe 0.05, S 0.06
$\mathrm{Ag}, \mathrm{As}, \mathrm{Sb}$ and Se were also measured; all were $<\mathrm{mdl}$ ( $0.14,0.15,0.09$ and $0.15 \mathrm{wt} \%$, respectively)

Table 5. Electron probe microanalyses of sphalerite and stannite


Coexisting sphalerite-stannite pairs

| $\log X_{\text {Fes }} / X_{\text {Zns sphalerite }}$ |  | $\log X_{\text {Fes }} / X_{\text {Zns stannite }}$ |  |
| :--- | :--- | :--- | :--- |
| $R G 59 / 116$ | -0.52 | $R G 59 / 115$ | -0.30 |
| RG59/121 | -0.67 | $R G 59 / 122$ | 0.03 |
| RG59/123 | -0.58 | $R G 59 / 124$ | 0.19 |
| RG59/126 | -0.60 | $R G 59 / 125$ | 0.31 |
| RG59/133 | -0.59 | $R G 59 / 134$ | 0.34 |
| RG59/135 | -0.75 | $R G 59 / 136$ | 1.08 |
| RG27/153 | -0.64 | $R G 27 / 152$ | 0.66 |

[^1]Ag was also measured but was <mdl ( $0.14 \mathrm{wt} \%$ )

Table 6. Electron probe microanalyses of gold/electrum, maldonite and hessite

|  | NATIVE G OLD / ELECTRUM |  |  |  |  |  |  |  |  | MALDONITE |  |  | HESSITE |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 25RG1 | 25RG33 | 62RG35 | 62RG37 | 62RG38 | 62RG44 | 52RG54 | 52RG56 | 52RG67 | 52RG59 | 52RG60 | 52RG61 | 52RG49 | 52RG63 |
| Au | 57.32 | 85.68 | 86.71 | 95.55 | 95.70 | 86.54 | 92.92 | 70.21 | 93.35 | 62.63 | 62.46 | 60.53 | <mdl | <mdl |
| Ag | 35.48 | 7.36 | 7.38 | 5.56 | 5.87 | 10.01 | 4.62 | 3.81 | 1.92 | 0.52 | 0.54 | 0.30 | 64.62 | 62.93 |
| Bi | <mdl | <mdl | <mdl | 0.19 | 0.24 | <mdl | 0.42 | 18.34 | 0.43 | 35.69 | 35.65 | 37.28 | <mdl | 0.39 |
| Te | 0.10 | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | 0.64 | 37.29 | 37.53 |
| S | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | <mdl | 0.17 | <mdl |
| Total | 92.90 | 93.05 | 94.09 | 101.31 | 101.81 | 96.55 | 97.95 | 92.37 | 95.70 | 98.85 | 98.64 | 98.74 | 102.08 | 100.84 |
| Formulae calculated to 1 a.p.f.u. |  |  |  |  |  |  |  |  |  | 3 a.p.f.u. |  |  | 3 a.p.f.u. |  |
| Au | 0.469 | 0.864 | 0.865 | 0.902 | 0.897 | 0.826 | 0.913 | 0.743 | 0.960 | 1.933 | 1.930 | 1.868 | - | - |
| Ag | 0.530 | 0.136 | 0.134 | 0.096 | 0.100 | 0.174 | 0.083 | 0.074 | 0.036 | 0.030 | 0.030 | 0.017 | 2.000 | 1.983 |
| Bi | 0.000 | 0.000 | 0.000 | 0.002 | 0.002 | - | 0.004 | 0.183 | 0.004 | 1.038 | 1.038 | 1.084 | - | 0.006 |
| Te | 0.001 | - | - | - | - | - | - | - | - | - | - | 0.031 | 0.976 | 0.999 |
| S | - | - | - | - | - | - | - | - | - | - | - | - | 0.018 | - |

Detection limits (wt\%) Au 0.37, Ag 0.16, Bi 0.15, Te 0.10, S 0.06
Se was also measured but was <mdl ( $0.15 \mathrm{wt} \%$ )
Analyses with low totals (small grain size) are retained to represent compositional variation

Table 7. Electron probe microanalyses of bismuth telluride species

|  |  | RG45 (in MoS2) | RG43 (in | MoS2) | $\begin{aligned} & \text { TSUMOITE (+S) } \\ & \text { RG43 (in MoS2) } \end{aligned}$ |  | Bi4(Te,S)5 RG43 (in MoS2) |  | $\mathrm{Bi} 4(\mathrm{Te}, \mathrm{~S}) 5$ <br> RG62 | HEDLEYITE <br> RG52 |  | Bi2Te <br> RG52 |  | Bi8Te3 <br> RG52 | HED Bi2Te <br> RG42 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | mean | SD | mean | SD | mean | SD | mean | SD |  | mean | SD | mean | SD |  |  |  |
| (wt. \%) | $\mathrm{n}=25$ |  | $\mathrm{n}=8$ |  | $\mathrm{n}=2$ |  | $\mathrm{n}=5$ |  |  | $\mathrm{n}=4$ |  | $\mathrm{n}=4$ |  |  |  |  |
| Ag | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl | <mdl |  | <mdl |  | <mdl | 0.15 | 0.66 |
| Pb | 0.89 | 0.24 | 1.27 | 0.09 | 1.21 | 0.03 | 0.91 | 0.20 | 0.60 | <mdl |  | <mdl |  | 0.08 | <mdl |  |
| Bi | 61.02 | 1.10 | 62.92 | 0.50 | 60.84 | 0.47 | 59.35 | 0.68 | 61.06 | 79.09 | 0.92 | 75.45 | 0.70 | 78.50 | 84.48 | 75.34 |
| Sb | 0.28 | 0.11 | <mdl |  | <mdl |  | <mdl |  | 0.26 | 0.25 | 0.06 | 0.23 | 0.07 | 0.17 | <mdl |  |
| Te | 36.64 | 0.72 | 35.78 | 0.50 | 34.53 | 1.02 | 35.82 | 0.30 | 29.89 | 19.82 | 0.70 | 22.62 | 0.70 | 17.64 | 18.79 | 27.49 |
| Se | 0.15 | 0.05 | 0.20 | 0.04 | 0.20 | 0.02 | 0.17 | 0.05 | 0.20 | <mdl |  | <mdl |  | <mdl | <mdl |  |
| S | 0.52 | 0.50 | 0.31 | 0.24 | 1.59 | 0.31 | 2.18 | 0.39 | 3.98 | <mdl |  | <mdl |  | 0.08 | <mdl |  |
| Total | 99.50 |  | 100.48 |  | 98.37 |  | 98.42 |  | 95.99 | 99.16 |  | 98.31 |  | 96.47 | 103.42 | 103.48 |
| Formulae |  | (2 a.p.f.u.) |  |  |  |  | (9 a.p.f.u.) |  | (9 a.p.f.u.) | (10 a.p.f.u.) |  | (3 a.p.f.u.) |  | $\begin{aligned} & (11 \\ & \text { a.p.f.u.) } \end{aligned}$ | (10 a.p. |  |
| Ag | - |  | - |  | - |  | - |  | - | - |  | - |  | - | 0.025 | 0.031 |
| Pb | 0.014 | 0.004 | 0.020 | 0.002 | 0.019 | 0.000 | 0.062 | 0.014 | 0.039 | - |  | - |  | 0.008 | - |  |
| Bi | 0.969 | 0.025 | 1.004 | 0.011 | 0.940 | 0.007 | 3.998 | 0.095 | 3.996 | 7.023 | 0.063 | 1.994 | 0.019 | 7.945 | 7.287 | 1.847 |
| Sb | 0.005 | 0.004 | - |  | - |  | - |  | 0.029 | 0.038 | 0.010 | 0.011 | 0.003 | 0.030 | - |  |
| Total M | 0.988 | 0.024 | 1.025 | 0.013 | 0.959 | 0.007 | 4.065 | 0.102 | 4.064 | $7.066 \quad 0.059$ |  | 2.011 | 0.018 | 7.983 | 7.312 | 1.879 |
| Te | 0.953 | 0.027 | 0.935 | 0.016 | 0.874 | 0.025 | 3.951 | 0.059 | 3.203 | 2.882 | 0.098 | 0.979 | 0.016 | 2.924 | 2.655 | 1.104 |
| Se | 0.006 | 0.002 | 0.008 | 0.001 | 0.008 | 0.001 | 0.030 | 0.009 | 0.035 | - |  | - |  | - | - |  |
| S | 0.053 | 0.050 | 0.032 | 0.025 | 0.160 | 0.031 | 0.954 | 0.154 | 1.697 | - |  | 0.003 | 0.002 | 0.056 | - |  |
| Total S,Se,Te | 1.012 | 0.024 | 0.975 | 0.013 | 1.041 | 0.007 | 4.935 | 0.102 | 4.936 | 2.934 | 0.059 | 0.989 | 0.018 | 3.017 | 2.688 | 1.121 |


|  | PHASE 'C' | ALEKSITE <br> RG45 |  | SADDLEBACKITE |
| :--- | :---: | :---: | :---: | :---: |
| (wt. \%) |  |  |  |  |
| Pb | 12.28 | 21.94 | 17.77 | 37.42 |
| Bi | 53.30 | 43.75 | 45.11 | 31.58 |
| Te | 32.91 | 25.38 | 26.99 | 18.22 |
| Se | 0.31 | 0.75 | 0.83 | 0.72 |
| S | 1.86 | 6.50 | 6.46 | 9.74 |
| Total | $\mathbf{1 0 0 . 7 3}$ | $\mathbf{9 8 . 3 6}$ | $\mathbf{9 7 . 2 0}$ | $\mathbf{9 7 . 7 0}$ |
|  |  |  |  |  |
| Formulae | $\mathbf{( 1 0}$ a.p.f.u.) | $\mathbf{( 7}$ a.p.f.u.) | (9 a.p.f.u.) |  |
| Pb | 0.934 | 1.020 | 0.828 | 2.064 |
| Bi | 4.018 | 2.017 | 2.082 | 1.726 |
| Total M | $\mathbf{4 . 9 6 1}$ | $\mathbf{3 . 0 4 1}$ | $\mathbf{2 . 9 1 4}$ | $\mathbf{3 . 7 9 3}$ |
| Te | 4.063 | 1.916 | 2.041 | 1.631 |


| Se | 0.061 | 0.091 | 0.101 | 0.104 |
| :--- | :--- | :--- | :--- | :--- |
| S | 0.915 | 1.952 | 1.944 | 3.472 |
| Total S,Se,Te | $\mathbf{5 . 0 3 9}$ | $\mathbf{3 . 9 5 9}$ | $\mathbf{4 . 0 8 6}$ | $\mathbf{5 . 2 0 7}$ |

Table 8. Summary of LA-ICP-MS spot analyses of molybdenite

| SampleName |  | Fe | Co | Ni | Cu | Zn | As | Se | Ag | Sn | Sb | Te | W | Re | Au | TI | Pb | Bi | $\mathbf{U}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| NK25 | mean | 10360 | 8.4 | 25 | 128 | 17 | 19 | 698 | 3.0 | 1.5 | 13 | 108 | 162 | 374 | 15 | 3.6 | 183 | 343 | 0.37 |
| $\mathrm{n}=10$ | SD | 10940 | 8.0 | 21 | 97 | 16 | 16 | 111 | 2.0 | 1.4 | 9.0 | 39 | 73 | 252 | 6.8 | 3.9 | 121 | 210 | 0.91 |
|  | max | 30682 | 24 | 66 | 391 | 47 | 48 | 930 | 7.2 | 4.9 | 30 | 151 | 262 | 905 | 27 | 14 | 440 | 685 | 2.9 |
|  | min | 813 | 1.7 | 5.8 | 68 | 2.7 | 4.9 | 516 | 1.0 | 0.11 | 2.9 | 45 | 69 | 94 | 7.5 | 0.31 | 50 | 82 | <mdl |
| RG8 | mean | 5499 | 7.6 | 12 | 996 | 44 | 4.7 | 553 | 51 | 2.8 | 5.3 | 188 | 153 | 122 | 78 | 0.85 | 194 | 876 | 0.04 |
| $\mathrm{n}=20$ | SD | 8112 | 9.4 | 13 | 1431 | 48 | 3.2 | 95 | 53 | 3.7 | 4.6 | 117 | 120 | 99 | 81 | 0.87 | 129 | 803 | 0.03 |
|  | max | 35218 | 42 | 49 | 6269 | 182 | 15 | 711 | 219 | 17 | 15 | 432 | 549 | 481 | 347 | 3.0 | 458 | 2855 | 0.11 |
|  | min | 253 | 0.47 | 1.1 | 54 | 3.3 | 1.3 | 414 | 1.5 | 0.24 | 0.17 | 16 | 58 | 36 | 2.1 | 0.04 | 25 | 42 | <mdl |
| RG10 | mean | 3336 | 2.0 | 3.6 | 15 | 14 | 13 | 1149 | 19 | 0.95 | 6.7 | 295 | 122 | 350 | 24 | 1.6 | 217 | 530 | 0.03 |
| $\mathrm{n}=15$ | SD | 3947 | 1.7 | 3.5 | 12 | 13 | 8.9 | 208 | 28 | 1.0 | 6.6 | 163 | 146 | 146 | 30 | 1.6 | 128 | 354 | 0.04 |
|  | max | 12669 | 5.6 | 12 | 51 | 43 | 32 | 1543 | 115 | 3.6 | 22 | 670 | 536 | 713 | 126 | 5.3 | 402 | 1338 | 0.12 |
|  | min | 58 | 0.15 | 0.25 | 2.7 | 1.5 | 3.6 | 779 | 0.16 | 0.12 | 0.33 | 133 | 8.1 | 150 | 0.63 | 0.03 | 16 | 34 | <mdl |
| RG25 | mean | 26853 | 17 | 53 | 158 | 79 | 11 | 603 | 57 | 3.2 | 8.0 | 865 | 306 | 4.7 | 70 | 3.5 | 1532 | 2929 | 0.41 |
| $\mathrm{n}=17$ | SD | 40893 | 14 | 54 | 280 | 106 | 5.8 | 94 | 62 | 2.4 | 5.5 | 922 | 172 | 4.2 | 87 | 2.3 | 1051 | 3076 | 0.47 |
|  | max | 141646 | 54 | 187 | 1201 | 343 | 24 | 740 | 229 | 8.3 | 19 | 3912 | 823 | 16 | 314 | 8.5 | 3440 | 12108 | 1.4 |
|  | min | 610 | 0.87 | 3.1 | 6.6 | 4.6 | 3.3 | 462 | 2.4 | 0.24 | 0.79 | 40 | 113 | 0.34 | 1.6 | 0.29 | 150 | 71 | <mdl |
| RG27 | mean | 4797 | 4.9 | 17 | 1261 | 43 | 5.1 | 485 | 9.8 | 13 | 0.75 | 46 | 104 | 126 | 9.4 | 0.50 | 105 | 113 | 0.08 |
| $\mathrm{n}=16$ | SD | 3429 | 7.3 | 25 | 1326 | 46 | 3.8 | 132 | 32 | 24 | 1.0 | 22 | 59 | 101 | 11 | 0.40 | 116 | 118 | 0.13 |
|  | max | 12591 | 29 | 104 | 4267 | 165 | 15 | 786 | 129 | 89 | 4.2 | 85 | 280 | 347 | 43 | 1.2 | 399 | 363 | 0.43 |
|  | min | 3.9 | 0.14 | 0.17 | 2.8 | 0.4 | 1.4 | 292 | 0.0 | 0.1 | 0.02 | 20 | 42 | 8.4 | 0.17 | 0.01 | 4.4 | 6.0 | <mdl |
| RG29 | mean | 1575 | 0.85 | 2.2 | 228 | 8.0 | 8.8 | 413 | 3.1 | 11 | 0.72 | 107 | 81 | 0.16 | 25 | 0.12 | 120 | 287 | 1.2 |
| $\mathrm{n}=10$ | SD | 1632 | 0.92 | 1.7 | 406 | 6.4 | 7.6 | 35 | 5.4 | 21 | 0.74 | 107 | 29 | 0.12 | 32 | 0.20 | 124 | 280 | 3.2 |
|  | max | 4606 | 3.1 | 4.7 | 1137 | 22 | 21 | 493 | 17 | 68 | 2.6 | 399 | 152 | 0.44 | 81 | 0.67 | 342 | 694 | 10 |
|  | min | 2.6 | 0.11 | 0.31 | 0.65 | 1.4 | 1.5 | 368 | <mdl | 0.23 | 0.04 | 42 | 50 | 0.03 | 0.03 | 0.01 | 0.28 | 1.3 | <mdl |
| RG30 | mean | 4045 | 1.2 | 4.8 | 96 | 11 | 5.2 | 328 | 42 | 6.5 | 0.9 | 105 | 79 | 0.43 | 116 | 0.38 | 96 | 186 | 0.17 |
| $\mathrm{n}=10$ | SD | 6790 | 1.4 | 7.3 | 123 | 15 | 7.7 | 30 | 133 | 7.7 | 1.3 | 126 | 57 | 0.55 | 349 | 0.64 | 91 | 359 | 0.45 |
|  | max | 22554 | 5.1 | 25 | 362 | 52 | 26 | 387 | 421 | 22 | 4.1 | 456 | 230 | 1.8 | 1110 | 2.1 | 257 | 1166 | 1.5 |
|  | min | 27 | 0.44 | 0.38 | 1.3 | 1.2 | 0.5 | 275 | <mdl | 0.16 | 0.08 | 39 | 40 | 0.03 | 0.24 | 0.01 | 1.4 | 4.4 | 0.003 |

Table 8
continued.

| Sample |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Name |  | Fe | Co | Ni | Cu | Zn | As | Se | Ag | Sn | Sb | Te | W | Re | Au | TI | Pb | Bi | U |
| RG40 | mean | 5549 | 3.0 | 0.27 | 202 | 12 | 6.0 | 208 | 6.5 | 7.4 | 1.7 | 723 | 206 | 5.2 | 11 | 1.9 | 428 | 1014 | 6.9 |
| $\mathrm{n}=12$ | SD | 6920 | 2.5 | 0.52 | 360 | 16 | 4.7 | 98 | 6.5 | 7.3 | 1.4 | 1144 | 114 | 3.9 | 15 | 1.8 | 330 | 1680 | 9.8 |
|  | max | 24128 | 7.8 | 1.9 | 1242 | 50 | 17 | 357 | 22 | 22 | 4.6 | 4256 | 378 | 17 | 53 | 5.2 | 963 | 6195 | 33 |
|  | $\min$ | <mdl | <mdl | <mdl | 0.77 | 0.30 | 0.45 | 88 | 0.02 | 0.08 | 0.02 | 40 | 35 | 0.66 | <mdl | <mdl | 0.19 | 0.07 | <mdl |
| RG43 | mean | 2544 | 2.4 | 0.19 | 400 | 165 | 4.9 | 444 | 17 | 2.5 | 2.4 | 4042 | 92 | 4.0 | 0.86 | 7.9 | 1863 | 6628 | 1.3 |
| $\mathrm{n}=12$ | SD | 2522 | 2.4 | 0.28 | 513 | 522 | 5.0 | 106 | 20 | 3.3 | 2.0 | 6624 | 94 | 4.8 | 0.90 | 8.3 | 1211 | 10808 | 2.8 |
|  | max | 6257 | 8.4 | 1.0 | 1433 | 1820 | 17 | 553 | 76 | 11 | 6.8 | 24705 | 335 | 18 | 3.3 | 25 | 4022 | 40308 | 9.4 |
|  | min | <mdl | 0.01 | <mdl | 1.8 | 0.20 | <mdl | 184 | 0.08 | 0.09 | 0.03 | 55 | 15 | 0.40 | 0.01 | 0.05 | 16 | 45 | <mdl |
| RG44 | mean | 2088 | 1.1 | 0.21 | 9.1 | 9.0 | 4.3 | 247 | 1.3 | 8.3 | 1.4 | 1159 | 69 | 2.4 | 0.16 | 2.5 | 449 | 1968 | 5.4 |
| $\mathrm{n}=10$ | SD | 2401 | 0.73 | 0.14 | 7.5 | 8.7 | 3.1 | 38 | 1.1 | 8.1 | 0.87 | 924 | 40 | 1.4 | 0.11 | 1.7 | 277 | 1698 | 8.2 |
|  | max | 7084 | 2.7 | 0.47 | 24 | 30 | 11 | 287 | 3.7 | 29 | 3.3 | 3459 | 165 | 4.4 | 0.36 | 6.6 | 910 | 6255 | 27 |
|  | $\min$ | 113 | 0.07 | 0.03 | 2.4 | 1.7 | 0.4 | 158 | 0.16 | 0.21 | 0.21 | 356 | 19 | 0.71 | 0.02 | 0.35 | 85 | 775 | 0.27 |
| RG45 | mean | 420 | 0.56 | 0.06 | 2.7 | 1.9 | 2.0 | 347 | 32 | 1.7 | 4.6 | 5093 | 77 | 2.9 | 0.61 | 2.4 | 4774 | 7885 | 0.50 |
| $\mathrm{n}=16$ | SD | 765 | 1.0 | 0.07 | 4.6 | 2.7 | 2.5 | 90 | 43 | 3.7 | 5.5 | 9435 | 109 | 1.3 | 1.0 | 3.6 | 4866 | 14166 | 1.4 |
|  | max | 2896 | 3.6 | 0.23 | 17 | 9.6 | 9.0 | 580 | 132 | 11 | 17 | 35599 | 364 | 5.6 | 3.3 | 11 | 15602 | 52682 | 5.1 |
|  | min | 1.6 | 0.01 | <mdl | 0.38 | 0.33 | 0.10 | 213 | 0.78 | 0.06 | 0.13 | 38 | 25 | 0.66 | <mdl | 0.02 | 46 | 23 | <mdl |
| RG49 | mean | 854 | 0.44 | 0.08 | 2.1 | 2.5 | 4.3 | 205 | 3.5 | 6.1 | 0.18 | 36 | 124 | 3.2 | 0.01 | 1.1 | 281 | 51 | 3.9 |
| $\mathrm{n}=10$ | SD | 1419 | 0.48 | 0.07 | 1.9 | 3.0 | 3.0 | 87 | 2.8 | 7.3 | 0.20 | 20 | 68 | 1.9 | 0.01 | 1.3 | 388 | 52 | 4.1 |
|  | max | 4044 | 1.4 | 0.17 | 6.1 | 9.8 | 10 | 317 | 7.9 | 25 | 0.53 | 68 | 235 | 7.6 | 0.04 | 3.9 | 983 | 152 | 11 |
|  | min | 2.3 | 0.02 | <mdl | 0.48 | 0.31 | 0.35 | 80 | 0.10 | 0.35 | <mdl | 10 | 50 | 1.3 | <mdl | <mdl | 2.8 | 0.84 | 0.002 |
| RG51 | mean | 1870 | 1.0 | 0.10 | 4.2 | 5.5 | 6.5 | 249 | 4.3 | 4.8 | 0.86 | 45 | 124 | 5.9 | 0.02 | 1.1 | 146 | 32 | 1.9 |
| $\mathrm{n}=12$ | SD | 2481 | 1.4 | 0.08 | 3.2 | 5.4 | 4.4 | 103 | 7.0 | 5.9 | 1.4 | 33 | 106 | 3.4 | 0.04 | 1.9 | 216 | 46 | 2.8 |
|  | max | 6743 | 4.3 | 0.32 | 9.2 | 16 | 12 | 340 | 23 | 18 | 4.4 | 129 | 354 | 14 | 0.14 | 5.4 | 612 | 139 | 7.5 |
|  | $\min$ | 1.3 | 0.01 | <mdl | 0.57 | 0.87 | 0.6 | 59 | 0.03 | 0.29 | <mdl | 8.0 | 38 | 0.92 | <mdl | <mdl | 0.19 | 0.02 | <mdl |
| RG55 | mean | 2501 | 2.7 | 8.4 | 837 | 24 | 5.7 | 501 | 5.2 | 4.7 | 1.2 | 123 | 244 | 1.3 | 41 | 0.44 | 150 | 389 | 0.54 |
| $\mathrm{n}=10$ | SD | 2237 | 1.8 | 6.5 | 709 | 22 | 2.5 | 53 | 4.3 | 3.3 | 1.2 | 79 | 128 | 1.2 | 48 | 0.31 | 99 | 303 | 0.64 |
|  | max | 8445 | 6.6 | 22 | 2473 | 86 | 8.8 | 593 | 14 | 12 | 4.4 | 212 | 476 | 4.0 | 148 | 1.1 | 292 | 803 | 2.1 |
|  | $\min$ | 914 | 0.57 | 2.1 | 237 | 9.3 | 1.1 | 419 | 0.54 | 1.1 | 0.22 | 32 | 64 | 0.36 | 2.1 | 0.09 | 26 | 27 | 0.01 |

Table 8
continued.

| Sample <br> Name |  | Fe | Co | Ni | Cu | Zn | As | Se | Ag | Sn | Sb | Te | W | Re | Au | TI | Pb | Bi | U |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| RG56 | mean | 7709 | 7.4 | 16 | 227 | 23 | 7.3 | 311 | 22 | 13 | 0.53 | 118 | 181 | 0.90 | 56 | 1.2 | 301 | 363 | 1.6 |
| $\mathrm{n}=14$ | SD | 7553 | 5.3 | 11 | 160 | 17 | 6.3 | 45 | 40 | 18 | 0.31 | 46 | 109 | 1.0 | 112 | 0.98 | 165 | 252 | 2.8 |
|  | max | 22375 | 20 | 37 | 524 | 62 | 24 | 409 | 125 | 69 | 1.2 | 184 | 410 | 3.3 | 435 | 2.5 | 577 | 831 | 8.8 |
|  | $\min$ | 147 | 0.15 | 0.50 | 15 | 1.2 | 0.36 | 268 | 0.17 | 0.16 | 0.01 | 48 | 41 | 0.04 | 0.62 | 0.02 | 7.5 | 12 | 0.01 |
| RG60 | mean | 1390 | 14 | 20 | 267 | 28 | 5.2 | 400 | 96 | 2.8 | 0.44 | 107 | 123 | 1.8 | 37 | 0.23 | 230 | 322 | 0.25 |
| $\mathrm{n}=12$ | SD | 2863 | 23 | 34 | 209 | 35 | 2.9 | 54 | 88 | 3.4 | 0.34 | 44 | 49 | 2.1 | 37 | 0.26 | 176 | 180 | 0.31 |
|  | max | 10401 | 83 | 125 | 760 | 125 | 12 | 469 | 326 | 13 | 1.4 | 182 | 207 | 5.9 | 122 | 1.0 | 610 | 655 | 0.93 |
|  | $\min$ | 57 | 1.4 | 2.0 | 34 | 2.3 | 2.6 | 317 | 16 | 0.14 | 0.16 | 50 | 65 | 0.15 | 1.6 | 0.03 | 54 | 99 | 0.00 |
| RG62 | mean | 3167 | 2.8 | 5.6 | 123 | 59 | 4.7 | 429 | 2.2 | 0.75 | 0.20 | 169 | 133 | 0.79 | 9.4 | 0.61 | 57 | 265 | 0.13 |
| $\mathrm{n}=15$ | SD | 4376 | 2.9 | 6.7 | 195 | 74 | 4.2 | 60 | 2.1 | 0.53 | 0.13 | 129 | 70 | 0.78 | 8.3 | 0.54 | 60 | 244 | 0.18 |
|  | max | 13456 | 11 | 25 | 706 | 270 | 18 | 540 | 5.5 | 1.9 | 0.39 | 543 | 269 | 2.0 | 26 | 1.6 | 234 | 999 | 0.52 |
|  | min | 40 | 0.51 | 0.43 | 15 | 0.18 | 0.28 | 342 | 0.14 | 0.07 | 0.02 | 43 | 50 | 0.05 | 0.55 | 0.02 | 1.3 | 12 | <mdl |
| RG64 | mean | 9345 | 14 | 12 | 365 | 29 | 12 | 497 | 64 | 3.4 | 2.8 | 193 | 255 | 36 | 85 | 4.9 | 413 | 572 | 2.7 |
| $\mathrm{n}=19$ | SD | 9718 | 14 | 11 | 682 | 29 | 8.5 | 70 | 101 | 5.2 | 3.2 | 95 | 175 | 71 | 138 | 5.4 | 392 | 402 | 8.7 |
|  | max | 33273 | 50 | 46 | 3010 | 85 | 31 | 687 | 386 | 22 | 11.4 | 434 | 602 | 218 | 517 | 18 | 1491 | 1356 | 39 |
|  | min | 1201 | 1.8 | 2.3 | 16 | 4.3 | 2.8 | 385 | 0.71 | 0.23 | 0.10 | 48 | 67 | 0.25 | 0.58 | 0.43 | 9.0 | 15 | 0.01 |
| RG65 | mean | 2864 | 3.6 | 3.9 | 157 | 35 | 5.1 | 344 | 63 | 14 | 1.5 | 141 | 96 | 1.7 | 44 | 0.57 | 84 | 569 | 2.0 |
| $\mathrm{n}=10$ | SD | 2971 | 5.6 | 2.9 | 229 | 39 | 3.4 | 29 | 144 | 19 | 3.0 | 137 | 24 | 1.4 | 84 | 0.56 | 82 | 674 | 2.1 |
|  | max | 9457 | 19 | 10 | 767 | 132 | 13 | 387 | 464 | 56 | 9.9 | 457 | 144 | 4.1 | 273 | 1.5 | 244 | 2105 | 5.1 |
|  | $\min$ | 28 | 0.04 | 0.47 | 3.0 | 2.4 | 1.8 | 310 | 0.05 | 0.34 | <mdl | 27 | 60 | 0.26 | 0.01 | 0.01 | 2.2 | 2.8 | 0.01 |

## APPENDIX

## APPENDIX FIGURE CAPTIONS

FIGURE A1. (a) Map view of the structural controls on the Central Diorite orebody showing the main shear zones and faults confining the ore body. (b) Map view of the major shear zones within the deposit and the location of the interlayered dykes that intruded the orebody.

FIGURE A2. (a) Map view showing the Wourhaming granite situated with respect to the shear zones and intrusive dykes. The location of drillcore WRD 09775002 is shown through the granite pluton. Inset is a table of granite samples collected and relative depths. (b) Oblique view showing the Wourhaming granite intercepting the intrusive dykes and major shear zones. The direction of drillcore WRD 09775002 is shown in relation to the granite.

FIGURE A3. (a) NW-SE cross-section showing lithologies in the Central Diorite orebody with the outline of the high-grade Southern Diorite Deep core. Sampled drillcores and also shown. (b) NW-SE cross-section showing the same area as in (a) but with Au concentrations plotted along the drillcores. The sampling intervals are shown in the three diorite drillcores.

FIGURE A4. (a) Plan view of the southern section of the South Pit showing the blocks with different grades and samples collected from the high grade zones. Inset the location of the South Pit with respect to the surrounding volcanics and Central Diorite. (b) Photograph of diorite sample with zeolite facies alteration and native copper dissemination and staining. (c) Photograph of South Pit wall to the east showing the trajectory of the South-Eastern shear. (d) Photograph of the pit floor after blast with ore grades marked out and chip-rock sample collection.

FIGURE A5. Schematic lithological columns and assay data for $\mathrm{Cu}, \mathrm{Mo}, \mathrm{Au}$ and Bi in the four drillcores sampled in this project. The sampling intervals are also marked.
FIGURE A6. (a) Lithological and mineralisation details for the sampled interval in the drillcore that intersects the granite (WRD 09775002). The individual sample locations are also marked. (b) Photograph illustrating a mineralised contact between sheared of diorite and granite. (c) Photograph showing fluorite veinlet in coarse grained granite. (d) Typical coarse molybdenite patches in coarse-grained granite (e) and (f) Polished blocks showing molybdenite in both coarse and fine-grained granite.
FIGURE A7. (a) Lithological and mineralisation details for the sampled interval in diorite from drillcore WBD 1063001. The individual sample locations are also marked. (b) Photograph of diorite showing the appearance of cross-cutting veinlets with mineralisation. (c) Photograph of diorite showing quartz veining with molybdenite occurring on the margin.
(d) Photograph showing the typical appearance of sheared diorite. (e) Polished block showing sulphides in a biotite rich zone.

FIGURE A8. (a) Lithological and mineralisation details for the sampled interval in diorite from drillcore WBD 10590002. The individual sample locations are also marked. (b-e) Polished blocks showing occurrence of molybdenite and other sulphides

FIGURE A9. (a) Lithological and mineralisation details for the sampled interval in diorite from drillcore WRD 10600004. The individual sample locations are also marked. (b) Photograph of quartz $\pm$ clinozoisite veinlets in diorite that have undergone deformation. (c) Photograph of diorite showing fine-grained shear zones with biotite. (d) Photograph of diorite showing porphyroblastic albite. (e) Polished block showing large, pegmatitic albite and associated molybdenite. (f) Photograph of large vein in diorite comprising milky quartz $\pm$ clinozoisite.

FIGURE A10. BSE images illustrating lithologies and alteration assemblages. (a) Typical Qz-Plag-Bt association in coarser diorite (b) Mica-rich bands characterising foliation in the fine-grained diorite. (c) Coarse-grained granite showing typical perthitic feldspar and the presence of Clz along the edges of molybdenite (d, e) Typical break-down of coarse pyroxene resulting in Hbn-Ti-minerals and epidote replacement in dolerite. (f) Enrichment in apatite and scheelite grains in diorite. Inclusions of REE-minerals in granite (g) and Ilmenite surrounding a rutile core (h). (i) Detail on the dusty inclusions in ilmenite from (h). (j) Sulphides along a veinlet of actinolite. The complexity of alteration assemblage is shown and also the transition to fine-grained diorite with little mineralisation. The veinlet selvage comprises epidote (replacing actinolite?) and biotite. (k-l) Clinozoisite embedding sulphides in the granite. (m) Oscillatory zonation in clinozoisite. (n) Biotite-Clz alteration in granite; note the dusty REE-mineral inclusions accompanying break-down of $\mathrm{Clz} /$ allanite. (o) Fe-rich chlorite (chamosite) interlayered with muscovite in granite.

FIGURE A11. BSE images showing the occurrence of Ca-rich stilbite (a-c) and aspects of accompanying minerals in the zeolite facies alteration. Stilbite intergrown with molybdenite (a), and replacing feldspar in diorite (b). Detail from (b) showing typical acicular aggregates of stilbite (c). Halo of quartz+clay minerals replacing plagioclase porphyroblast in diorite. (e) Late hydrothermal biotite interlayered with chlorite (f) and overgrown by Clz (f). (g, h) Breakdown of allanite resulting in fields of dusty inclusions with REE-minerals and enrichment in Th. (i) Large grain of thorite nucleating in areas as in (h).

FIGURE A12. BSE images of sulphide assemblages in diorite host rock. (a) Typical twophase pyrrhotite with lamellar exsolutions. Note the pointed edges of the lamellae. (b)

Skeletal exsolutions of Po in Cub. (c) Argentopentlandite as inclusions in Pn. (d) Fields of small exsolutions of Sph-Stn in cubanite. (e) Coarser-Sph-Stn inclusions in Cub (f) Stn segregated at the boundary between Sph and Cp . (g) Co-bearing Mck in Cp. (h-j) Unusual thiospinel occurrence (embedded within bastnasite) occurring at the boundaries between pyrite and chalcopyrite. (k) Minor bornite and covellite replacing chalcopyrite with cubanite exsolutions. Chalcocite (l) and native $\mathrm{Cu}(\mathrm{m})$ in the diorite altered to zeolite facies.

FIGURE A13. LA-ICPMS element map of a coarse molybdenite lamella in diorite from the Open Pit (sample RG10) showing a tight microfold on one side and silicate replacement on the other side. This grain shows high- Re concentration with slight depletion along replacement boundaries. Tungsten shows strong concentration relative to replacement boundaries. The chalcophile elements $(\mathrm{Bi}, \mathrm{Pb}, \mathrm{Te})$ show a strong correlation with one another and correlate with spots of high-Au concentration.

FIGURE A14. LA-ICPMS element map of coarse-grained, highly-deformed molybdenite from diorite in ABreccia (North Pit, sample NK25). Rhenium is highly concentrated within molybdenite lamellae and depleted along the boundaries. Tungsten shows an inverse trend. Small, spot-concentration of Au and Ag are present and show correlation with Pb and Bi distribution.

FIGURE A15. Secondary electron and ion beam images showing the experimental procedure during FIB-TEM sample preparation (as marked). The duplication in ( $\mathbf{a}, \mathbf{b}$ ) and (d, e) is meant to show details of the tellurides that have been sliced, e.g. pilsenite in 62RG in (a), tsumoite in 45 RG in (b, e) and the Bi-Pb-tellurides that have been analysed by EPMA on sample 45 RG in (e); the latter slice was lost during welding on the Cu grid.

FIGURE A16. Secondary electron images showing location and details of the five FIB-TEM prepared foils from molybdenite in granite and diorite. These targeted specific-site extraction of material for TEM study necessary for crystal-structure and nanoscale characterisation of Bi-Pb-tellurides and host molybdenite.

## APPENDIX TABLES LIST

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Table A7. LA-ICPMS data of molybdenite

Figure A1


## WRD 09775002-Granite



## Figure A3

## Central Diorite Drill Cores



Figure A4


Figure A5

WRD 09775002


WBD 105900002


Shear


Dolerite Dyke
Granite

WRD 09775002
Figure A6


Figure A7

## WBD 1063001



Figure A8

## a WBD 10590002



# WRD10600004 

Figure A9


Figure A10


Figure A11


Figure A12



NK 25


Mo95_cps


Co59_cps


Re185_cps


Te125_cps


S34_cps


Ni60_cps
 W182_cps


Bi209_cps


Figure A14


Al27_cps


Cu63_cps


Au197_cps


Pb208_cps

Sb121_cps


Fe57 cps


Ag107_cps


Figure A15


Figure A16


Table A1. Microprobe analyses of feldspar group minerals

| (Wt. \%) | RG42 |  |  |  | RG43 |  |  | RG45 |  |  |  | RG47 |  |  |  | RG49 |  |  |  | RG50 |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | albite |  | K-feldspar |  | albite | K-feldspar |  | plagioclase |  | K-feldspar |  | albite |  | K-feldspar |  | albite |  | K-feldspar |  | albite |  | K-feldspar |  |
|  | mean | SD | mean | SD |  | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD |
|  | $\mathrm{n}=2$ |  | $\mathrm{n}=3$ |  |  | $\mathrm{n}=4$ |  | $\mathrm{n}=5$ |  | $\mathrm{n}=6$ |  | $\mathrm{n}=8$ |  | $\mathrm{n}=15$ |  | $\mathrm{n}=7$ |  | $\mathrm{n}=7$ |  | $\mathrm{n}=10$ |  | $\mathrm{n}=2$ |  |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| CaO | 0.04 | 0.01 | <mdl |  | 0.66 | <mdl |  | 0.46 | 0.29 | <mdl |  | 0.69 | 0.44 | <mdl |  | 0.54 | 0.28 | <mdl |  | 3.45 | 1.91 | 0.16 | 0.04 |
| Na 2 O | 10.68 | 1.54 | 0.34 | 0.07 | 11.22 | 0.28 | 0.03 | 9.66 | 3.04 | 0.31 | 0.08 | 11.35 | 0.22 | 0.35 | 0.07 | 11.48 | 0.20 | 0.27 | 0.66 | 9.64 | 1.25 | 1.35 | 0.36 |
| K2O | 0.10 | 0.01 | 16.39 | 0.27 | 0.05 | 16.41 | 0.08 | 0.95 | 1.91 | 16.19 | 0.14 | 0.09 | 0.03 | 16.25 | 0.21 | 0.07 | 0.02 | 16.35 | 0.85 | 0.39 | 0.58 | 13.81 | 1.86 |
| FeO | 0.04 | 0.01 | 0.06 | 0.02 | 0.15 | 0.05 | 0.05 | 0.29 | 0.49 | <mdl |  | <mdl |  | 0.04 | 0.04 | <mdl |  | <mdl |  | 0.27 | 0.17 | 0.33 | 0.45 |
| TiO2 | <mdl |  | <mdl |  | <mdl | 0.04 | 0.03 | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | 0.03 | 0.04 |
| MgO | <mdl |  | <mdl |  | <mdl | <mdl |  | 0.05 | 0.09 | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  |
| SiO2 | 71.26 | 3.67 | 64.06 | 0.52 | 68.38 | 63.95 | 0.30 | 67.12 | 3.52 | 64.07 | 0.33 | 67.77 | 0.34 | 64.07 | 0.60 | 67.35 | 0.78 | 64.49 | 0.54 | 64.11 | 2.46 | 60.91 | 4.52 |
| MnO | <mdl |  | <mdl |  | 0.03 | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  | <mdl |  |
| Al2O3 | 18.08 | 2.01 | 18.23 | 0.24 | 20.56 | 18.31 | 0.17 | 20.08 | 0.54 | 18.23 | 0.29 | 20.32 | 0.70 | 17.94 | 0.35 | 20.28 | 0.50 | 18.36 | 0.28 | 22.32 | 1.49 | 21.95 | 4.44 |
| Cl | <mdl |  | 0.02 | 0.03 | <mdl | <mdl |  | 0.04 | 0.06 | 0.02 | 0.01 | <mdl |  | 0.03 | 0.10 | <mdl |  | <mdl |  | 0.02 | 0.05 | <mdl |  |
| Total | 100.2 |  | 99.1 |  | 101.0 | 99.0 |  | 98.6 |  | 98.8 |  | 100.2 |  | 98.6 |  | 99.7 |  | 99.5 |  | 100.2 |  | 98.5 |  |
| Formula calculated to $8 \mathbf{O}$ a.p.f.u. |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Ca | 0.002 | - | - | 0.001 | 0.031 | - | 0.000 | 0.022 | 0.014 | - | 0.000 | 0.032 | 0.020 | - | 0.002 | 0.025 | 0.013 | - | 0.006 | 0.163 | 0.091 | 0.008 | 0.0021 |
| Na | 0.895 | 0.1405 | 0.031 | 0.006 | 0.941 | 0.026 | 0.002 | 0.830 | 0.260 | 0.028 | 0.007 | 0.960 | 0.0233 | 0.032 | 0.007 | 0.976 | 0.016 | 0.024 | 0.060 | 0.825 | 0.099 | 0.122 | 0.0333 |
| K | 0.006 | - | 0.977 | 0.013 | 0.003 | 0.979 | 0.004 | 0.054 | 0.105 | 0.966 | 0.010 | 0.005 | 0.0017 | 0.972 | 0.014 | 0.004 | 0.001 | 0.969 | 0.048 | 0.022 | 0.033 | 0.823 | 0.1034 |
| Total | 0.903 | 0.140 | 1.008 | 0.008 | 0.975 | 1.004 | 0.004 | 0.905 | 0.167 | 0.995 | 0.009 | 0.998 | 0.0121 | 1.004 | 0.015 | 1.006 | 0.012 | 0.993 | 0.020 | 1.010 | 0.014 | 0.953 | 0.068 |
| Al | 0.921 | 0.1143 | 1.004 | 0.008 | 1.049 | 1.009 | 0.010 | 1.049 | 0.022 | 1.005 | 0.014 | 1.045 | 0.029 | 0.991 | 0.015 | 1.049 | 0.028 | 1.006 | 0.014 | 1.161 | 0.084 | 1.208 | 0.2555 |
| Si | 3.081 | 0.120 | 2.993 | 0.004 | 2.959 | 2.990 | 0.007 | 2.974 | 0.050 | 2.996 | 0.010 | 2.957 | 0.0269 | 3.004 | 0.009 | 2.955 | 0.025 | 2.997 | 0.009 | 2.829 | 0.086 | 2.845 | 0.1856 |
| Mg | - |  | - |  | - | - |  | 0.003 | 0.007 | - |  | - |  | - |  | - |  | - |  | - |  | - |  |
| Fe | 0.002 | - | 0.002 | 0.001 | 0.006 | 0.002 | 0.002 | 0.011 | 0.019 | - |  | - |  | 0.001 | 0.001 | - |  | - |  | 0.010 | 0.006 | 0.013 | 0.0178 |
| Mn | - |  | - |  | 0.001 | - |  | - |  | - |  | - |  | - |  | - |  | - |  | - |  | - |  |
| Ti | - |  | - |  | - | 0.001 | 0.001 | - |  | - |  | - |  | - |  | - |  | - |  | - |  | 0.001 | 0.0014 |
| Total | 0.003 | - | 0.005 | 0.001 | 0.007 | 0.004 | 0.003 | 0.015 | 0.026 | 0.002 | 0.001 | 0.003 | 0.0018 | 0.003 | 0.002 | 0.002 | 0.001 | 0.001 | 0.002 | 0.012 | 0.006 | 0.016 | 0.0192 |
| ab \% | 99.1 | 0.2 | 3.1 | 0.6 | 96.5 | 2.5 | 0.2 | 91.7 | 16.1 | 2.9 | 0.8 | 96.3 | 2.1 | 3.2 | 0.7 | 97.1 | 1.3 | 2.4 | 5.7 | 81.7 | 10.4 | 12.8 | 4.4 |
| an \% | 0.2 | 0.1 | 0.0 | 0.1 | 3.2 | 0.0 | 0.0 | 2.4 | 1.5 | 0.0 | 0.0 | 3.2 | 2.0 | 0.0 | 0.2 | 2.5 | 1.3 | 0.0 | 0.5 | 16.1 | 8.9 | 0.8 | 0.3 |
| Kfs \% | 0.6 | 0.1 | 96.9 | 0.6 | 0.3 | 97.5 | 0.2 | 6.0 | 17.0 | 97.1 | 0.7 | 0.5 | 0.2 | 96.8 | 0.6 | 0.4 | 0.1 | 97.6 | 6.2 | 2.2 | 3.3 | 86.4 | 4.7 |

Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, Na 0.04, P 0.03, Si 0.02 , Ti 0.02
F and Cr contents were also measured but all values were $<\mathrm{mdl}(0.17 \mathrm{wt} \%, 0.01 \mathrm{wt} \%$, respectively $)$

Table A1. Microprobe analyses of feldspar group minerals (continued)


Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, Na $0.04, \mathrm{P} 0.03, \mathrm{Si} 0.02$, Ti 0.02

F and Cr contents were also measured but all values were $<\mathrm{mdl}(0.17 \mathrm{wt} \%, 0.01 \mathrm{wt} \%$, respectively)

Table A2. Microprobe analyses of biotite group minerals

|  | RG45 | RG43 |  | RG50 |  | RG27 |  | RG36 |  | RG25 |  | RG32 |  | RG29 |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD |
| (Wt.\%) |  | $\mathrm{n}=3$ |  | $\mathrm{n}=9$ |  | $n=30$ |  | $\mathrm{n}=18$ |  | $\mathrm{n}=18$ |  | $\mathrm{n}=5$ |  | $1 \mathrm{n}=6$ |  | II $\mathrm{n}=3$ |  |
| CaO | 0.00 | 0.13 | 0.03 | 0.48 | 1.36 | 0.03 | 0.06 | 0.03 | 0.06 | 0.18 | 0.56 | <mdl |  | <mdl |  | <mdl |  |
| Na 2 O | 0.06 | 0.06 | 0.04 | 0.05 | 0.03 | 0.05 | 0.03 | 0.22 | 0.48 | 0.11 | 0.23 | 0.06 | 0.03 | 0.07 | 0.03 | 0.08 | 0.05 |
| K2O | 9.41 | 8.05 | 0.65 | 9.37 | 0.43 | 9.38 | 0.18 | 9.48 | 0.58 | 9.28 | 0.59 | 9.55 | 0.10 | 9.28 | 0.44 | 8.92 | 0.50 |
| FeO | 18.80 | 26.49 | 0.85 | 22.68 | 1.17 | 19.54 | 0.53 | 17.82 | 1.54 | 17.19 | 0.71 | 19.92 | 0.39 | 18.29 | 2.17 | 15.57 | 0.75 |
| TiO2 | 0.57 | 1.16 | 0.04 | 0.96 | 1.71 | 0.65 | 0.05 | 0.90 | 0.24 | 0.97 | 0.52 | 0.74 | 0.06 | 0.68 | 0.07 | 0.74 | 0.07 |
| MgO | 12.36 | 5.87 | 0.04 | 8.43 | 0.73 | 11.47 | 0.35 | 10.38 | 0.81 | 11.97 | 0.66 | 10.54 | 0.29 | 11.74 | 1.93 | 14.21 | 0.09 |
| SiO 2 | 38.31 | 34.33 | 0.26 | 35.42 | 0.55 | 36.59 | 0.62 | 38.01 | 4.17 | 36.31 | 1.67 | 36.48 | 0.67 | 36.90 | 1.27 | 37.12 | 0.85 |
| MnO | 0.36 | 0.25 | 0.04 | 0.30 | 0.05 | 0.32 | 0.05 | 0.22 | 0.05 | 0.19 | 0.03 | 0.27 | 0.05 | 0.28 | 0.06 | 0.22 | 0.01 |
| Cr2O3 | 0.02 | 0.01 | 0.01 | <mdl |  | 0.01 | 0.01 | 0.02 | 0.01 | 0.02 | 0.01 | 0.02 | 0.01 | 0.01 | 0.01 | <mdl |  |
| Al2O3 | 15.19 | 17.41 | 0.33 | 16.26 | 0.64 | 15.27 | 0.32 | 17.19 | 1.25 | 16.63 | 0.68 | 16.73 | 0.28 | 16.18 | 0.73 | 16.29 | 0.32 |
| F | 2.71 | 0.50 | 0.04 | 1.44 | 0.11 | 2.45 | 0.13 | 0.89 | 0.06 | 1.23 | 0.10 | 1.71 | 0.11 | 1.74 | 0.39 | 1.28 | 0.12 |
| Cl | 0.05 | 0.08 | 0.01 | 0.07 | 0.03 | 0.06 | 0.02 | 0.04 | 0.02 | 0.06 | 0.02 | 0.06 | 0.01 | 0.06 | 0.02 | 0.06 | 0.03 |
| Total | 97.77 | 94.16 |  | 95.45 |  | 95.74 |  | 94.94 |  | 94.13 |  | 96.07 |  | 95.23 |  | 94.49 | 1.02 |

Formula based on Biotite $\mathrm{K} 2(\mathrm{Fe}, \mathrm{Mg}, \mathrm{Mn}) 6 \mathrm{Al} 2 \mathrm{Si6O20}(\mathrm{OH}, \mathrm{F}) 4$

| Na | 0.019 | 0.019 | 0.012 | 0.014 | 0.009 | 0.016 | 0.010 | 0.019 | 0.006 | 0.031 | 0.067 | 0.019 | 0.009 | 0.020 | 0.005 | 0.022 | 0.014 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K | 1.823 | 1.641 | 0.131 | 1.887 | 0.093 | 1.852 | 0.037 | 1.894 | 0.024 | 1.831 | 0.081 | 1.882 | 0.020 | 1.879 | 0.127 | 1.734 | 0.084 |
| Total | 1.841 | 1.660 |  | 1.901 |  | 1.868 |  | 1.913 |  | 1.862 |  | 1.901 |  | 1.899 |  | 1.756 |  |
| Mg | 2.796 | 1.399 | 0.005 | 1.983 | 0.178 | 2.668 | 0.072 | 2.470 | 0.043 | 2.763 | 0.120 | 2.425 | 0.054 | 2.431 | 0.055 | 3.228 | 0.030 |
| Fe | 2.387 | 3.541 | 0.115 | 2.994 | 0.155 | 2.574 | 0.068 | 2.395 | 0.049 | 2.227 | 0.090 | 2.573 | 0.063 | 2.554 | 0.070 | 1.985 | 0.108 |
| Mn | 0.047 | 0.034 | 0.006 | 0.040 | 0.007 | 0.040 | 0.006 | 0.027 | 0.006 | 0.024 | 0.005 | 0.036 | 0.006 | 0.042 | 0.006 | 0.028 | 0.002 |
| Ca | 0.000 | 0.022 | 0.005 | 0.080 | 0.228 | 0.003 | 0.003 | 0.001 | 0.002 | 0.029 | 0.091 | - |  | - |  | - |  |
| Ti | 0.066 | 0.139 | 0.005 | 0.113 | 0.201 | 0.077 | 0.005 | 0.108 | 0.019 | 0.113 | 0.059 | 0.087 | 0.008 | 0.075 | 0.006 | 0.085 | 0.008 |
| Total | 5.296 | 5.134 |  | 5.211 |  | 5.362 |  | 5.000 |  | 5.156 |  | 5.122 |  | 5.105 |  | 5.327 |  |
| Al | 2.719 | 3.280 | 0.046 | 3.026 | 0.131 | 2.813 | 0.069 | 3.155 | 0.062 | 3.036 | 0.072 | 3.046 | 0.064 | 2.953 | 0.040 | 2.927 | 0.051 |
| Cr | 0.003 | 0.001 | 0.002 | - |  | 0.001 | 0.001 | 0.002 | 0.002 | 0.002 | 0.002 | 0.002 | 0.002 | 0.001 | 0.002 | - |  |
| Total | 2.721 | 3.282 |  | 3.027 |  | 2.814 |  | 3.157 |  | 3.038 |  | 3.048 |  | 2.955 |  | 2.928 |  |
| Si | 5.818 | 5.487 | 0.049 | 5.593 | 0.097 | 5.703 | 0.061 | 5.600 | 0.029 | 5.622 | 0.076 | 5.635 | 0.070 | 5.719 | 0.027 | 5.659 | 0.073 |
| F | 1.301 | 0.254 | 0.021 | 0.720 | 0.057 | 1.200 | 0.077 | 0.432 | 0.031 | 0.606 | 0.054 | 0.833 | 0.050 | 0.972 | 0.143 | 0.617 | 0.063 |
| Cl | 0.012 | 0.021 | 0.003 | 0.018 | 0.008 | 0.017 | 0.006 | 0.011 | 0.007 | 0.015 | 0.006 | 0.016 | 0.003 | 0.016 | 0.005 | 0.016 | 0.007 |
| OH | 2.687 | 3.725 | 0.022 | 3.262 | 0.054 | 2.783 | 0.075 | 3.557 | 0.029 | 3.379 | 0.054 | 3.151 | 0.051 | 3.011 | 0.146 | 3.367 | 0.058 |
| Total | 4.000 | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  |
| TOTAL | 19.68 | 19.56 |  | 19.73 |  | 19.75 |  | 19.67 |  | 19.68 |  | 19.70 |  | 19.68 |  | 19.67 |  |
| Mol.\% end-mem |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| Annite | 45.1 | 69.0 | 0.6 | 57.6 | 4.3 | 48.0 | 1.2 | 47.9 | 0.7 | 43.2 | 1.4 | 50.2 | 0.9 | 50.0 | 0.9 | 37.3 | 1.1 |
| Phlogopite | 52.8 | 27.3 | 0.6 | 38.2 | 4.2 | 49.8 | 1.1 | 49.4 | 1.0 | 53.6 | 1.8 | 47.3 | 1.1 | 47.6 | 1.0 | 60.6 | 1.0 |
| Fluorphlogopite | 32.5 | 6.3 | 0.5 | 18.0 | 1.4 | 30.0 | 1.9 | 10.8 | 0.8 | 15.1 | 1.4 | 20.8 | 1.2 | 24.3 | 3.6 | 15.4 | 1.6 |

Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, F 0.17, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, $\mathrm{Na} 0.04, \mathrm{P} 0.03, \mathrm{Si} 0.02, \mathrm{Ti} 0.02$

|  | RG29 |  | RG32 |  | RG42 |  | RG43 |  | RG45 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | SD | mean | SD | mean | SD | mean | SD | mean | SD |
| (Wt.\%) | $\mathrm{n}=3$ |  | $n=7$ |  | $\mathrm{n}=6$ |  | $\mathrm{n}=2$ |  | $\mathrm{n}=3$ |  |
| CaO | <mdl |  | 0.05 | 0.11 | <mdl |  | <mdl |  | 0.02 | 0.01 |
| Na 2 O | 0.23 | 0.04 | 0.36 | 0.43 | 0.15 | 0.04 | 0.16 | 0.02 | 0.17 | 0.03 |
| K2O | 10.01 | 0.25 | 9.30 | 2.10 | 10.34 | 1.41 | 11.01 | 0.01 | 10.73 | 0.18 |
| FeO | 2.76 | 0.47 | 2.34 | 0.58 | 4.40 | 1.09 | 3.81 | 0.37 | 1.89 | 0.09 |
| TiO2 | 0.10 | 0.03 | 0.12 | 0.04 | 0.08 | 0.03 | 0.22 | 0.11 | <mdl |  |
| MgO | 2.03 | 0.09 | 1.86 | 0.43 | 1.22 | 0.28 | 1.28 | 0.40 | 0.62 | 0.04 |
| SiO 2 | 47.77 | 1.03 | 51.33 | 9.80 | 47.23 | 1.15 | 47.22 | 0.98 | 46.23 | 0.42 |
| MnO | 0.06 | 0.01 | 0.03 | 0.02 | 0.03 | 0.03 | 0.09 | 0.02 | 0.03 | 0.02 |
| Cr 2 O 3 | <mdl |  | 0.02 | 0.01 | 0.01 | 0.01 | <mdl |  | 0.01 | 0.02 |
| Al2O3 | 31.30 | 0.40 | 29.69 | 5.80 | 31.23 | 1.80 | 31.59 | 1.80 | 34.37 | 0.33 |
| F | 0.69 | 0.08 | 0.49 | 0.14 | 0.39 | 0.05 | 0.47 | 0.02 | <mdl |  |
| Cl | <mdl |  | <mdl |  | <mdl |  | <mdl |  | 0.02 | 0.01 |
| Total | 94.94 |  | 95.57 |  | 95.08 |  | 95.85 |  | 94.09 |  |

Formula on basis of KAl2(Si3AI)O10(OH,F)2

| Na | 0.029 | 0.005 | 0.044 | 0.049 | 0.020 | 0.005 | 0.021 | 0.003 | 0.022 | 0.004 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| K | 0.862 | 0.023 | 0.791 | 0.194 | 0.898 | 0.137 | 0.948 | 0.000 | 0.925 | 0.013 |
| Total | 0.891 |  | 0.835 |  | 0.918 |  | 0.968 |  | 0.947 |  |
| Total AI | 2.490 | 0.025 | 2.332 | 0.513 | 2.497 | 0.095 | 2.513 | 0.140 | 2.737 | 0.033 |
| Si | 3.223 | 0.040 | 3.375 | 0.458 | 3.206 | 0.038 | 3.187 | 0.070 | 3.123 | 0.020 |
| Al (iv) | 0.777 | 0.040 | 0.625 | 0.458 | 0.794 | 0.038 | 0.813 | 0.070 | 0.877 | 0.020 |
| Total | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  | 4.000 |  |
| Al (iv) | 1.712 | 0.021 | 1.708 | 0.055 | 1.702 | 0.079 | 1.700 | 0.070 | 1.860 | 0.013 |
| Cr | - |  | 0.001 | 0.001 | 0.001 | 0.000 | - |  | 0.001 | 0.001 |
| Mg | 0.204 | 0.012 | 0.185 | 0.046 | 0.123 | 0.030 | 0.129 | 0.041 | 0.063 | 0.004 |
| Fe | 0.156 | 0.028 | 0.131 | 0.036 | 0.250 | 0.063 | 0.215 | 0.021 | 0.107 | 0.005 |
| Mn | 0.003 | 0.001 | 0.002 | 0.001 | 0.002 | 0.002 | 0.005 | 0.001 | 0.002 | 0.001 |
| Ca | - |  | 0.003 | 0.007 | - |  | - |  | 0.001 | 0.001 |
| Ti | 0.005 | 0.002 | 0.006 | 0.002 | 0.004 | 0.001 | 0.011 | 0.005 | - |  |
| Total | 2.082 |  | 2.035 |  | 2.083 |  | 2.061 |  | 2.034 |  |
| TOTAL | 6.973 |  | 6.870 |  | 7.001 |  | 7.029 |  | 6.981 |  |
| F | 0.148 | 0.016 | 0.103 | 0.032 | 0.085 | 0.012 | 0.101 | 0.005 | 0.025 | 0.019 |
| Cl | 0.000 |  | 0.000 |  | 0.000 |  | 0.000 |  | 0.002 | 0.001 |
| OH | 1.852 | 0.016 | 1.896 | 0.033 | 1.915 | 0.012 | 1.899 | 0.005 | 1.973 | 0.019 |
| Total | 2.000 |  | 2.000 |  | 2.000 |  | 2.000 |  | 2.000 |  |
| Phengite component \% | 17.7 | 1.8 | 15.9 | 3.0 | 18.2 | 2.9 | 17.5 | 3.2 | 8.5 | 0.4 |
| $\mathrm{Fe} /(\mathrm{Fe}+\mathrm{Mg})$ | 0.43 | 0.03 | 0.41 | 0.02 | 0.66 | 0.08 | 0.63 | 0.05 | 0.63 | 0.01 |
| \% F-end-member | 7.4 | 0.8 | 5.2 | 1.6 | 4.2 | 0.6 | 5.0 | 0.2 | 1.2 | 1.0 |
| \% Cl-end-member | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.0 | 0.1 | 0.0 |

Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, F 0.17, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, Na 0.04, P 0.03, Si 0.02, Ti 0.02

Table A4. Microprobe analyses of epidote group minerals

|  | RG42 |  | RG43 |  | RG45 |  | RG49 |  | RG50 |  | RG21 | RG25 | RG29 |  | RG36 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | mean | SD | mean | SD | mean | SD | mean | SD | mean | SD |  |  | mean | SD | mean | SD |
| (Wt.\%) | $\mathrm{n}=25$ |  | $\mathrm{n}=10$ |  | $\mathrm{n}=20$ |  | $\mathrm{n}=-10$ |  | $\mathrm{n}=-19$ |  |  |  | $\mathrm{n}=-24$ |  | $\mathrm{n}=-34$ |  |
| CaO | 24.77 | 0.33 | 23.75 | 0.81 | 24.31 | 0.36 | 24.25 | 0.35 | 24.28 | 0.94 | 22.03 | 24.69 | 24.69 | 0.19 | 24.70 | 0.46 |
| Na 2 O | <mdl |  | 0.05 | 0.04 | <mdl |  | <mdl |  | 0.08 | 0.17 | <mdl | <mdl | <mdl |  | <mdl |  |
| K2O | 0.01 | 0.01 | 0.01 | 0.01 | 0.02 | 0.01 | 0.01 | 0.01 | 0.01 | 0.02 | 0.12 | <mdl | 0.01 | 0.01 | 0.01 | 0.01 |
| FeO (meas.) | 7.71 | 1.20 | 8.78 | 0.51 | 8.04 | 0.99 | 10.53 | 2.19 | 8.16 | 1.92 | 11.40 | 6.11 | 6.60 | 0.57 | 7.99 | 1.16 |
| Fe2O3 (calc.) | 8.57 | 1.33 | 9.76 | 0.57 | 8.94 | 1.11 | 11.70 | 2.43 | 9.07 | 2.13 | 12.67 | 6.79 | 7.33 | 0.63 | 8.89 | 1.28 |
| TiO2 | <mdl |  | 0.04 | 0.03 | 0.03 | 0.03 | 0.06 | 0.04 | 0.34 | 0.94 | 0.05 | <mdl | 0.04 | 0.04 | 0.04 | 0.04 |
| MgO | <mdl |  | <mdl |  | <mdl |  | <mdl |  | 0.03 | 0.04 | 0.65 | <mdl | <mdl |  | <mdl |  |
| SiO2 | 37.80 | 0.53 | 37.52 | 0.22 | 37.65 | 0.28 | 36.80 | 1.12 | 37.72 | 0.72 | 35.05 | 38.29 | 37.41 | 0.70 | 37.92 | 0.41 |
| MnO | 0.07 | 0.07 | 0.13 | 0.05 | 0.12 | 0.09 | 0.07 | 0.07 | 0.12 | 0.07 | 0.17 | 0.14 | 0.17 | 0.06 | 0.07 | 0.05 |
| Cr203 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.01 | 0.03 | 0.03 | 0.01 | 0.01 | 0.01 | 0.02 |
| Al203 | 26.95 | 1.16 | 25.34 | 0.54 | 26.09 | 0.82 | 24.51 | 2.02 | 26.08 | 2.10 | 23.64 | 28.22 | 27.00 | 0.67 | 26.49 | 1.07 |
| F | <mdl |  | 0.21 | 0.08 | 0.18 | 0.06 | 0.19 | 0.04 | 0.17 | 0.08 | 0.16 | <mdl | 0.17 | 0.05 | <mdl |  |
| Cl | <mdl |  | 0.04 | 0.05 | <mdl |  | <mdl |  | <mdl |  | 0.02 | <mdl | <mdl |  | <mdl |  |
| Total | 98.17 |  | 96.86 |  | 97.33 |  | 97.59 |  | 97.92 |  | 94.60 | 98.16 | 96.84 |  | 98.13 |  |

Formula, based on $\mathrm{Ca} 2(\mathrm{Al}, \mathrm{Fe} 3+$ ) $3 \mathrm{Si} 3 \mathrm{O} 2(\mathrm{OH})$

| Ca | 2.079 | 0.014 | 2.031 | 0.044 | 2.064 | 0.023 | 2.078 | 0.024 | 2.050 | 0.065 | 1.95184 | 2.05625 | 2.0991 | 0.03053 | 2.07601 | 0.02753 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Na | - |  | 0.008 | 0.006 | - |  | - |  | 0.012 | 0.026 | - | - | - |  | - |  |
| K | 0.001 | 0.001 | 0.001 | 0.001 | 0.002 | 0.001 | 0.001 | 0.001 | 0.002 | 0.003 | 0.01301 | - | 0.001 | 0.001 | 0.001 | 0.00114 |
| Total | 2.082 |  | 2.040 |  | 2.068 |  | 2.081 |  | 2.064 |  | 1.9657 | 2.0563 | 2.102 |  | 2.0788 |  |
| Cr | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.001 | 0.002 | 0.002 | 0.001 | 0.001 | 0.001 | 0.001 |
| Mg | - |  | - |  | - |  | - |  | 0.003 | 0.005 | 0.080 | - | - |  | - |  |
| Fe | 0.506 | 0.083 | 0.586 | 0.031 | 0.533 | 0.066 | 0.707 | 0.159 | 0.539 | 0.135 | 0.789 | 0.397 | 0.438 | 0.038 | 0.525 | 0.079 |
| Mn | 0.004 | 0.005 | 0.009 | 0.004 | 0.008 | 0.006 | 0.004 | 0.005 | 0.008 | 0.005 | 0.012 | 0.009 | 0.012 | 0.004 | 0.004 | 0.003 |
| Ti | - |  | 0.002 | 0.002 | 0.002 | 0.002 | 0.003 | 0.003 | 0.020 | 0.056 | 0.003 | - | 0.002 | 0.002 | 0.003 | 0.002 |
| Al | 2.488 | 0.085 | 2.385 | 0.029 | 2.437 | 0.072 | 2.309 | 0.147 | 2.422 | 0.167 | 2.305 | 2.586 | 2.525 | 0.038 | 2.449 | 0.078 |
| Total | 2.999 |  | 2.984 |  | 2.980 |  | 3.024 |  | 2.993 |  | 3.191 | 2.995 | 2.978 |  | 2.982 |  |
| Si | 2.961 | 0.018 | 2.996 | 0.035 | 2.984 | 0.015 | 2.943 | 0.034 | 2.974 | 0.044 | 2.900 | 2.977 | 2.969 | 0.011 | 2.975 | 0.023 |
| (OH) | 0.961 | 0.015 | 0.942 | 0.027 | 0.952 | 0.015 | 0.952 | 0.010 | 0.956 | 0.020 | 0.955 | 0.971 | 0.957 | 0.012 | 0.961 | 0.015 |
| F | - |  | 0.053 | 0.021 | 0.046 | 0.016 | 0.047 | 0.011 | 0.043 | 0.020 | 0.042 | - | 0.043 | 0.012 | - |  |
| Cl | - |  | 0.005 | 0.007 | - |  | - |  | - |  | 0.003 | - | - |  | - |  |
| XFe | 16.9 | 2.7 | 19.6 | 1.0 | 17.9 | 2.2 | 23.4 | 5.1 | 18.0 | 4.5 | 24.7 | 13.3 | 14.7 | 1.3 | 17.6 | 2.6 |
| XMn | 0.1 | 0.2 | 0.3 | 0.1 | 0.3 | 0.2 | 0.1 | 0.2 | 0.3 | 0.2 | 0.4 | 0.3 | 0.4 | 0.1 | 0.1 | 0.1 |
| XAI | 82.9 | 2.8 | 79.9 | 1.0 | 81.8 | 2.3 | 76.4 | 5.1 | 80.9 | 5.2 | 72.2 | 86.3 | 84.8 | 1.2 | 82.1 | 2.6 |

Detection Limits (wt\%) Al 0.01, Ca 0.02, Cr 0.01, Cl 0.02, F 0.17, Fe 0.03, K 0.01, Mn 0.03, Mg 0.02, Na 0.04, P 0.03, Si $0.02, \mathrm{Ti} 0.02$

Table A5. Electron probe microanalyses of stilbite


Table A6. Electron probe microanalyses of pyrite, pentlandite, argentopentlandite, cubanite, mackinawite and Fe-Ni-thiospinel

|  | $\begin{gathered} \text { PYRITE } \\ \text { RG18 } \\ \hline \end{gathered}$ |  | PENTLANDITERG24 |  | ARGENTOPENTLANDITE RG25 |  | CUBANITE |  |  | CO-NI MACKINAWITE |  |  | FE-NI THIOSPINEL RG36 |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | RG24 |  |  |  |  |  |  | RG36 |  |  |
|  | mean | SD |  |  | mean | SD | mean | SD |  | mean | SD | mean | SD |  | mean | SD |
| Wt.\% | $\mathrm{n}=31$ |  | $\mathrm{n}=5$ |  | $\mathrm{n}=3$ |  |  | $\mathrm{n}=18$ |  | $\mathrm{n}=8$ |  |  | $\mathrm{n}=3$ |  |
| Ag |  |  | <mdl |  | 11.47 | 1.13 | <mdl |  |  | <mdl |  |  | <mdl |  |
| Cu | <mdl |  | 1.19 | 1.04 | 2.17 | 1.31 | 23.36 | 22.95 | 0.24 | 0.59 | 0.30 | 0.40 | 0.44 | 0.20 |
| Fe | 45.94 | 0.46 | 32.11 | 0.50 | 37.04 | 1.14 | 39.67 | 40.52 | 0.21 | 53.06 | 1.06 | 56.23 | 35.57 | 3.89 |
| Co | 0.08 | 0.06 | 4.68 | 0.88 | 0.19 | 0.17 | <mdl | <mdl |  | 4.99 | 1.57 | 0.32 | 0.38 | 0.19 |
| Ni | 0.04 | 0.03 | 27.54 | 2.15 | 15.98 | 0.40 | <mdl | <mdl |  | 3.69 | 0.48 | 5.20 | 20.18 | 1.52 |
| S | 52.91 | 1.75 | 33.25 | 0.25 | 32.05 | 0.66 | 35.40 | 35.58 | 0.20 | 35.63 | 0.56 | 35.54 | 40.35 | 1.22 |
| Total | 98.97 |  | 98.77 |  | 98.90 |  | 98.42 | 99.04 |  | 97.96 |  | 97.69 | 96.92 |  |
| Formula (3 a.p.f.u.) |  |  | Formula (17 a.p.f.u.) |  | Formula (17 a.p.f.u.) |  | Formula (6 a.p.f.u.) |  |  | Formula (2 a.p.f.u.) |  |  | Formula (7 a.p.f.u.) |  |
| Ag |  |  | - |  | 0.916 | 0.102 | - | - |  | - |  |  | - |  |
| Cu | 0.000 | 0.000 | 0.146 | 0.127 | 0.293 | 0.176 | 1.010 | 0.986 | 0.010 | 0.008 | 0.004 | 0.006 | 0.022 | 0.010 |
| Fe | 0.997 | 0.026 | 4.483 | 0.058 | 5.714 | 0.246 | 1.952 | 1.982 | 0.009 | 0.856 | 0.019 | 0.909 | 1.978 | 0.203 |
| Co | 0.002 | 0.001 | 0.620 | 0.118 | 0.027 | 0.025 | - | - |  | 0.076 | 0.024 | 0.005 | 0.020 | 0.010 |
| Ni | 0.001 | 0.001 | 3.659 | 0.286 | 2.345 | 0.028 | - | - |  | 0.057 | 0.008 | 0.080 | 1.068 | 0.084 |
| Total M | 1.000 | 0.027 | 8.911 | 0.040 | 8.387 | 0.060 | 2.966 | 2.968 | 0.009 | 0.999 | 0.004 | 0.999 | 3.088 | 0.108 |
| s | 2.000 | 0.027 | 8.089 | 0.040 | 8.613 | 0.060 | 3.034 | 3.032 | 0.009 | 1.001 | 0.004 | 1.001 | 3.912 | 0.108 |
| M/S | 0.50 | 0.02 | 1.10 | 0.01 | 0.97 | 0.01 | 0.98 | 0.98 | 0.01 | 0.997 | 0.009 | 0.999 | 0.790 | 0.049 |

Detection limits (wt\%) Ag 0.15, Cu 0.06, Co 0.04, Ni 0.03, Fe 0.05, S 0.06
$\mathrm{Mn}, \mathrm{As}, \mathrm{Sb}$ and Se were also measured; all were <mdl ( $0.03,0.15,0.09$ and $0.15 \mathrm{wt} \%$, respectively)

Table A7. LA-ICPMS data of molybdenite.



| AU04A129 | RG27.1 | 1662 | 1.1 | 11 | 91 | 16 | 3.8 | 428 | 1.1 | 3.1 | 0.27 | 23 | 133 | 26 | 3.9 | 0.34 | 58 | 40 | 0.13 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| AU04A130 | RG27.2 | 4477 | 6.7 | 26 | 4267 | 30 | 2.8 | 786 | 2.8 | 9.9 | 1.01 | 85 | 54 | 131 | 14 | 0.82 | 155 | 94 | 0.00 |
| AU04A131 | RG27.3 | 3068 | 3.9 | 8.8 | 1368 | 17 | 1.8 | 605 | 0.98 | 2.8 | 1.12 | 72 | 51 | 234 | 5.2 | 0.38 | 304 | 132 | 0.01 |
| AU04A132 | RG27.4 | 6338 | 5.2 | 14 | 1357 | 66 | 2.2 | 492 | 3.1 | 89 | 1.20 | 64 | 112 | 46 | 22 | 0.52 | 181 | 247 | 0.28 |
| AU04A133 | RG27.5 | 4921 | 0.8 | 2.8 | 110 | 16 | 3.0 | 750 | 0.14 | 3.5 | 0.16 | 69 | 60 | 117 | 0.49 | 0.09 | 9.1 | 12 | 0.03 |
| AU04A134 | RG27.6 | 283 | 0.7 | 2.1 | 191 | 13 | 2.7 | 465 | 0.50 | 0.65 | 0.07 | 25 | 98 | 79 | 3.6 | 0.01 | 6.5 | 18 | 0.00 |
| AU04A135 | RG27.7 | 5137 | 3.9 | 20 | 1767 | 80 | 5.0 | 502 | 2.1 | 8.4 | 0.95 | 57 | 78 | 22 | 13 | 1.1 | 198 | 175 | 0.02 |
| AU04A138 | RG27.8 | 4962 | 1.5 | 6.4 | 13 | 11 | 7.6 | 392 | 0.6 | 0.47 | 0.14 | 54 | 76 | 162 | 0.17 | 0.52 | 16 | 70 | 0.01 |
| AU04A139 | RG27.9 | 7051 | 29 | 104 | 199 | 64 | 11 | 439 | 129 | 5.4 | 0.39 | 40 | 162 | 98 | 14 | 0.81 | 57 | 363 | 0.03 |
| AU04A140 | RG27.10 | 3069 | 1.5 | 3.6 | 3430 | 16 | 10 | 376 | 2.8 | 4.5 | 0.38 | 41 | 56 | 8.4 | 12 | 0.07 | 42 | 65 | 0.00 |
| AU04A141 | RG27.11 | 12591 | 13 | 30 | 3211 | 165 | 4.9 | 467 | 10 | 55 | 4.24 | 69 | 280 | 122 | 43 | 1.15 | 399 | 361 | 0.43 |
| AU04A142 | RG27.12 | 5526 | 2.4 | 8.5 | 1112 | 36 | 2.2 | 436 | 1.2 | 7.4 | 0.71 | 20 | 99 | 347 | 4.4 | 0.45 | 71 | 41 | 0.08 |
| AU04A143 | RG27.13 | 9601 | 4.3 | 15 | 1200 | 123 | 4.0 | 397 | 0.97 | 5.7 | 0.76 | 41 | 107 | 48 | 6.2 | 0.91 | 112 | 132 | 0.30 |
| AU04A144 | RG27.14 | 7493 | 3.6 | 24 | 1242 | 38 | 4.5 | 292 | 1.1 | 13 | 0.41 | 22 | 119 | 291 | 7.5 | 0.80 | 53 | 39 | 0.02 |
| AU04A145 | RG27.15 | 3.9 | 0.14 | 0.17 | 2.8 | 0.41 | 15 | 390 | 0.05 | 0.12 | 0.02 | 22 | 135 | 59 | 0.20 | 0.01 | 4.4 | 6.8 | 0.00 |
| AU04A146 | RG27.16 | 573 | 0.27 | 2.4 | 618 | 2.2 | 1.4 | 539 | 0.16 | 0.69 | 0.22 | 25 | 42 | 221 | 0.21 | 0.01 | 12 | 6.0 | 0.00 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A163 | RG29.1 | 3155 | 0.62 | 3.1 | 10 | 7.5 | 16 | 493 | 0.24 | 1.02 | 0.74 | 44 | 79 | 0.44 | 6.7 | 0.14 | 45 | 153 | 0.06 |
| AU04A164 | RG29.2 | 1633 | 1.3 | 4.3 | 106 | 11 | 2.0 | 411 | 17 | 11 | 1.0 | 118 | 95 | 0.13 | 66 | 0.10 | 257 | 586 | 10 |
| AU04A165 | RG29.3 | 73 | 0.23 | 0.31 | 4.7 | 2.6 | 1.5 | 427 | 0.16 | 0.23 | 0.12 | 60 | 61 | 0.24 | 3.4 | 0.01 | 13 | 25 | 0.06 |
| AU04A166 | RG29.4 | 1012 | 1.6 | 3.2 | 99 | 5.5 | 9.2 | 410 | 1.2 | 3.9 | 0.81 | 70 | 86 | 0.09 | 10 | 0.17 | 342 | 128 | 0.41 |
| AU04A167 | RG29.5 | 563 | 0.42 | 0.74 | 43 | 5.2 | 19 | 428 | 0.26 | 1.7 | 0.44 | 80 | 152 | 0.15 | 5.0 | 0.05 | 119 | 141 | 0.55 |
| AU04A168 | RG29.6 | 2.6 | 0.19 | 0.43 | 0.7 | 1.4 | 21 | 377 | 0.00 | 15.19 | 0.04 | 48 | 66 | 0.03 | 0.03 | 0.01 | 0.28 | 1.3 | 0.00 |
| AU04A169 | RG29.7 | 4606 | 0.64 | 1.6 | 43 | 12 | 6.2 | 417 | 5.7 | 4.26 | 0.90 | 79 | 57 | 0.22 | 67 | 0.04 | 71 | 694 | 0.39 |
| AU04A170 | RG29.8 | 1017 | 0.27 | 3.4 | 1137 | 22 | 2.0 | 368 | 0.90 | 1.0 | 0.45 | 399 | 73 | 0.11 | 12 | 0.02 | 85 | 626 | 0.00 |
| AU04A171 | RG29.9 | 3538 | 3.1 | 4.7 | 829 | 11 | 2.8 | 409 | 6.0 | 68 | 2.6 | 131 | 91 | 0.04 | 81 | 0.67 | 265 | 500 | 0.43 |
| AU04A172 | RG29.10 | 148 | 0.11 | 0.49 | 4.4 | 1.4 | 7.3 | 388 | 0.08 | 0.46 | 0.13 | 42 | 50 | 0.17 | 0.20 | 0.02 | 4.3 | 11 | 0.02 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A211 | RG30.1 | 5680 | 0.70 | 3.7 | 362 | 10.8 | 3.1 | 320 | 0.06 | 2.9 | 0.24 | 51 | 67 | 0.09 | 0.24 | 0.21 | 19 | 7.4 | 0.03 |
| AU04A212 | RG30.2 | 2015 | 0.52 | 2.1 | 4.1 | 7.4 | 2.3 | 338 | 0.10 | 0.86 | 0.08 | 57 | 66 | 0.22 | 2.9 | 0.19 | 7.2 | 16 | 0.06 |
| AU04A213 | RG30.3 | 2397 | 1.0 | 2.5 | 87 | 7.7 | 0.5 | 341 | 0.37 | 3.4 | 0.18 | 66 | 112 | 0.03 | 6.7 | 0.22 | 23 | 48 | 0.05 |
| AU04A214 | RG30.4 | 5056 | 1.6 | 8.5 | 278 | 18.3 | 0.6 | 298 | 0.44 | 16 | 0.58 | 62 | 40 | 0.29 | 4.3 | 0.68 | 144 | 77 | 0.03 |
| AU04A215 | RG30.5 | 22554 | 5.1 | 25 | 41 | 52.0 | 4.4 | 275 | 421 | 22 | 4.1 | 456 | 230 | 1.8 | 1110 | 2.1 | 257 | 1166 | 1.5 |
| AU04A216 | RG30.6 | 27 | 0.44 | 0.38 | 21 | 1.2 | 2.0 | 340 | 0.03 | 0.37 | 0.17 | 49 | 43 | 0.57 | 0.34 | 0.01 | 150 | 5.7 | 0.01 |


|  | AU04A217 | RG30.7 | 666 | 0.99 | 1.8 | 79 | 2.5 | 3.6 | 346 | 1.4 | 4.9 | 2.5 | 120 | 58 | 0.21 | 24 | 0.08 | 181 | 299 | 0.04 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | AU04A218 | RG30.8 | 439 | 0.63 | 1.5 | 63 | 3.9 | 0.5 | 317 | 0.73 | 14 | 1.2 | 105 | 42 | 0.16 | 8.7 | 0.12 | 155 | 211 | 0.03 |
|  | AU04A219 | RG30.9 | 461 | 0.66 | 1.0 | 1.3 | 1.7 | 25.8 | 318 | 0.07 | 0.99 | 0.13 | 42 | 65 | 0.03 | 0.35 | 0.04 | 1.4 | 4.4 | 0.04 |
|  | AU04A220 | RG30.10 | 1154 | 0.67 | 2.1 | 23 | 6.0 | 9.4 | 387 | 0.10 | 0.16 | 0.11 | 39 | 67 | 0.89 | 4.5 | 0.10 | 27 | 23 | 0.00 |
| WRD09775002 |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| GRANITE | FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
|  | AU04A183 | RG40.1 | 4605 | 4.0 | 0.10 | 1242 | 50 | 4.5 | 323 | 6.7 | 6.0 | 2.8 | 4256 | 68 | 17 | 7.6 | 2.5 | 888 | 6195 | 0.93 |
|  | AU04A184 | RG40.2 | 668 | 2.9 | 0.03 | 86 | 6.4 | 3.2 | 89 | 6.0 | 1.4 | 1.8 | 643 | 289 | 3.4 | 13 | 0.67 | 437 | 952 | 0.34 |
|  | AU04A185 | RG40.3 | 2858 | 1.7 | 0.22 | 65 | 2.6 | 3.8 | 128 | 15 | 7.3 | 1.4 | 769 | 248 | 5.6 | 53 | 1.6 | 528 | 1073 | 14 |
|  | AU04A186 | RG40.4 | 88 | 0.73 | 0.02 | 7.2 | 0.64 | 1.5 | 92 | 0.79 | 0.16 | 0.48 | 40 | 364 | 4.6 | 0.48 | 0.05 | 85 | 52 | 0.0 |
|  | AU04A187 | RG40.5 | 10926 | 3.8 | 0.31 | 155 | 13 | 5.3 | 194 | 10 | 22 | 2.3 | 545 | 177 | 5.1 | 6.5 | 5.2 | 442 | 686 | 7.9 |
|  | AU04A188 | RG40.6 | 9650 | 6.5 | 0.20 | 106 | 23 | 10 | 188 | 6.3 | 16 | 3.0 | 652 | 265 | 5.3 | 17 | 4.3 | 963 | 930 | 12 |
|  | AU04A189 | RG40.7 | 0.0 | 0.00 | 0.00 | 0.77 | 0.30 | 0.45 | 357 | 0.0 | 0.1 | 0.02 | 67 | 35 | 4.2 | 0.00 | 0.00 | 0 | 0 | 0.0 |
|  | AU04A190 | RG40.8 | 2165 | 7.8 | 0.06 | 69 | 2.4 | 8.1 | 223 | 6.1 | 4.1 | 4.6 | 784 | 182 | 7.0 | 20 | 3.7 | 812 | 1132 | 1.1 |
|  | AU04A191 | RG40.9 | 2133 | 1.7 | 0.08 | 16 | 2.3 | 9.2 | 323 | 1.8 | 1.3 | 0.89 | 169 | 44 | 0.66 | 3.2 | 0.44 | 214 | 214 | 1.3 |
|  | AU04A192 | RG40.10 | 7743 | 5.1 | 1.9 | 96 | 4.0 | 17 | 199 | 22 | 12 | 2.8 | 225 | 205 | 4.0 | 6.6 | 3.1 | 522 | 319 | 33 |
|  | AU04A193 | RG40.11 | 1628 | 0.64 | 0.01 | 23 | 1.0 | 2.2 | 88 | 1.3 | 2.3 | 0.43 | 211 | 378 | 1.9 | 2.6 | 0.47 | 121 | 261 | 0.37 |
|  | AU04A194 | RG40.12 | 24128 | 1.1 | 0.38 | 560 | 33 | 6.4 | 294 | 2.5 | 16 | 0.35 | 315 | 214 | 4.4 | 1.3 | 0.58 | 125 | 351 | 13 |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
|  | FileName AU04A093 | SampleName RG43.1 | Fe57 0.0 | Co59 0.01 | Ni60 0.01 | Cu65 1.8 | Zn66 0.20 | As75 0.0 | Se77 495 | Ag107 0.08 | Sn118 0.10 | Sb121 0.03 | Te125 55 | W182 15 | Re185 18 | Au197 0.01 | T1205 0.0 | Pb208 16 | Bi209 45 | U238 0.05 |
|  | AU04A094 | RG43.2 | 5030 | 4.8 | 1.0 | 60 | 32 | 8.2 | 184 | 20 | 6.2 | 4.1 | 3646 | 335 | 6.2 | 1.2 | 12 | 2562 | 6218 | 9.4 |
|  | AU04A095 | RG43.3 | 5188 | 8.4 | 0.31 | 1157 | 1820 | 17 | 470 | 76 | 11 | 6.8 | 2048 | 94 | 2.2 | 3.3 | 25 | 4022 | 3470 | 4.1 |
|  | AU04A096 | RG43.4 | 6257 | 2.6 | 0.26 | 73 | 10 | 5.1 | 512 | 10 | 2.8 | 2.2 | 3109 | 49 | 5.9 | 1.5 | 9.3 | 2842 | 5076 | 0.80 |
|  | AU04A097 | RG43.5 | 5858 | 2.6 | 0.22 | 531 | 33 | 7.7 | 445 | 10 | 4.3 | 1.9 | 4380 | 49 | 0.54 | 0.66 | 9.1 | 1692 | 7258 | 0.25 |
|  | AU04A098 | RG43.6 | 600 | 0.30 | 0.09 | 770 | 4.2 | 0.2 | 493 | 5.0 | 0.23 | 0.97 | 1710 | 22 | 4.3 | 0.31 | 1.4 | 1279 | 2776 | 0.00 |
|  | AU04A099 | RG43.7 | 57 | 1.5 | 0.08 | 8.7 | 0.7 | 1.1 | 486 | 19 | 0.12 | 1.9 | 1859 | 26 | 2.1 | 0.53 | 1.4 | 1678 | 3040 | 0.00 |
|  | AU04A100 | RG43.8 | 1666 | 0.37 | 0.00 | 734 | 13 | 1.5 | 481 | 6.8 | 0.09 | 0.86 | 347 | 28 | 1.9 | 0.06 | 0.24 | 806 | 558 | 0.00 |
|  | AU04A101 | RG43.9 | 4109 | 3.4 | 0.10 | 1433 | 59 | 7.1 | 450 | 17 | 3.0 | 2.9 | 2202 | 146 | 1.4 | 1.2 | 8.4 | 2279 | 3820 | 0.52 |
|  | AU04A102 | RG43.10 | 1301 | 3.5 | 0.09 | 21 | 3.8 | 1.9 | 553 | 26 | 0.91 | 5.0 | 24705 | 114 | 1.4 | 0.54 | 21 | 3459 | 40308 | 0.00 |
|  | AU04A103 | RG43.11 | 16 | 0.63 | 0.00 | 2.8 | 0.3 | 9.0 | 481 | 5.3 | 0.16 | 0.65 | 1624 | 40 | 0.40 | 0.15 | 1.2 | 434 | 2469 | 0.00 |
|  | AU04A104 | RG43.12 | 445 | 0.93 | 0.08 | 6.4 | 1.7 | 1.3 | 275 | 9.1 | 0.66 | 1.2 | 2814 | 191 | 4.2 | 0.80 | 5.4 | 1282 | 4501 | 0.53 |
|  | FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
|  | AU04A221 | RG44.1 | 2262 | 1.8 | 0.47 | 24 | 15 | 11 | 158 | 2.7 | 29 | 2.4 | 356 | 165 | 2.0 | 0.27 | 3.1 | 904 | 775 | 27 |
|  | AU04A222 | RG44.2 | 396 | 1.4 | 0.08 | 12 | 2.2 | 4.6 | 250 | 3.7 | 5.9 | 1.5 | 1113 | 80 | 1.4 | 0.15 | 3.2 | 621 | 1706 | 4.1 |


| AU04A223 | RG44.3 | 5683 | 0.69 | 0.28 | 5.1 | 14 | 2.5 | 251 | 0.38 | 11 | 1.6 | 496 | 48 | 4.4 | 0.08 | 2.6 | 257 | 779 | 0.67 |
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| AU04A224 | RG44.4 | 1502 | 0.53 | 0.24 | 4.3 | 5.7 | 2.4 | 233 | 0.60 | 4.2 | 0.88 | 765 | 58 | 1.4 | 0.11 | 1.1 | 258 | 1134 | 1.0 |
| AU04A225 | RG44.5 | 113 | 0.07 | 0.11 | 2.8 | 1.9 | 0.4 | 286 | 0.16 | 0.21 | 0.21 | 734 | 19 | 4.4 | 0.02 | 0.35 | 85 | 1176 | 0.37 |
| AU04A226 | RG44.6 | 229 | 1.1 | 0.03 | 2.4 | 1.7 | 1.9 | 287 | 0.46 | 0.73 | 1.0 | 1483 | 35 | 4.4 | 0.08 | 2.7 | 351 | 2565 | 0.27 |
| AU04A227 | RG44.7 | 510 | 2.7 | 0.26 | 7.6 | 4.7 | 7.9 | 278 | 1.4 | 9.3 | 3.3 | 3459 | 64 | 0.71 | 0.27 | 6.6 | 910 | 6255 | 11 |
| AU04A228 | RG44.8 | 7084 | 0.92 | 0.34 | 21 | 30 | 3.8 | 259 | 1.3 | 9.0 | 0.92 | 845 | 92 | 2.0 | 0.11 | 2.8 | 321 | 1324 | 4.9 |
| AU04A229 | RG44.9 | 1127 | 0.91 | 0.18 | 5.4 | 6.6 | 3.0 | 227 | 0.90 | 4.3 | 1.4 | 1786 | 68 | 2.6 | 0.36 | 1.4 | 355 | 3111 | 0.81 |
| AU04A230 | RG44.10 | 1975 | 0.83 | 0.09 | 7.2 | 8.2 | 5.6 | 236 | 1.3 | 10 | 1.1 | 550 | 65 | 1.2 | 0.13 | 1.5 | 432 | 852 | 3.6 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A018 | RG45.1 | 86 | 0.45 | 0.05 | 0.99 | 0.47 | 1.9 | 444 | 132 | 0.20 | 8.6 | 14902 | 25 | 4.9 | 0.40 | 8.3 | 7202 | 23342 | 0.03 |
| AU04A019 | RG45.2 | 299 | 0.49 | 0.03 | 1.42 | 1.6 | 1.1 | 354 | 18 | 0.81 | 1.0 | 101 | 43 | 2.2 | 0.23 | 0.63 | 4167 | 179 | 0.05 |
| AU04A020 | RG45.3 | 290 | 0.09 | 0.02 | 1.04 | 1.1 | 0.3 | 325 | 12 | 0.16 | 0.79 | 962 | 30 | 0.89 | 0.04 | 1.2 | 1908 | 1540 | 0.01 |
| AU04A021 | RG45.4 | 4.4 | 0.07 | 0.03 | 0.83 | 0.39 | 0.7 | 360 | 33 | 0.12 | 1.7 | 1188 | 27 | 2.0 | 0.05 | 0.18 | 6522 | 1865 | 0.00 |
| AU04A022 | RG45.5 | 5.1 | 0.06 | 0.02 | 0.83 | 0.69 | 0.2 | 314 | 5.3 | 0.10 | 1.0 | 785 | 32 | 1.8 | 0.07 | 0.06 | 1608 | 1210 | 0.01 |
| AU04A025 | RG45.6 | 2.3 | 0.01 | 0.03 | 0.49 | 0.33 | 0.64 | 316 | 3.1 | 0.07 | 17 | 3037 | 32 | 3.1 | 0.01 | 0.02 | 1762 | 4911 | 0.04 |
| AU04A026 | RG45.7 | 1.6 | 0.06 | 0.00 | 0.58 | 0.45 | 3.3 | 355 | 0.8 | 0.08 | 0.75 | 711 | 45 | 2.0 | 0.0 | 0.04 | 243 | 1075 | 0.00 |
| AU04A027 | RG45.8 | 334 | 0.79 | 0.00 | 0.64 | 1.1 | 1.8 | 580 | 130 | 0.14 | 16 | 35599 | 52 | 3.7 | 0.50 | 8.7 | 15602 | 52682 | 0.00 |
| AU04A028 | RG45.9 | 303 | 0.09 | 0.06 | 1.8 | 1.6 | 5.9 | 283 | 0.8 | 1.5 | 0.13 | 38 | 62 | 0.7 | 0.26 | 0.16 | 46 | 23 | 0.13 |
| AU04A029 | RG45.10 | 35 | 0.2 | 0.02 | 0.38 | 0.35 | 0.66 | 477 | 64 | 0.07 | 9.1 | 13832 | 37 | 2.5 | 0.18 | 1.1 | 10940 | 22156 | 0.00 |
| AU04A030 | RG45.11 | 2896 | 3.6 | 0.23 | 17 | 9.6 | 9.0 | 236 | 44 | 11 | 8.7 | 1938 | 364 | 3.1 | 2.0 | 11 | 11390 | 3271 | 5.1 |
| AU04A031 | RG45.12 | 1326 | 2.3 | 0.15 | 11 | 6.4 | 4.6 | 213 | 32 | 11 | 5.1 | 514 | 346 | 5.6 | 3.3 | 4.6 | 9619 | 954 | 2.5 |
| AU04A032 | RG45.13 | 24 | 0.04 | 0.00 | 1.2 | 0.45 | 0.10 | 311 | 1.1 | 0.09 | 0.22 | 70 | 33 | 3.7 | 0.08 | 0.04 | 216 | 64 | 0.01 |
| AU04A033 | RG45.14 | 4.9 | 0.12 | 0.03 | 2.0 | 0.52 | 1.1 | 322 | 2.6 | 0.06 | 0.61 | 495 | 38 | 2.8 | 2.3 | 0.05 | 611 | 602 | 0.00 |
| AU04A034 | RG45.15 | 68 | 0.23 | 0.07 | 1.2 | 0.53 | 0.31 | 305 | 6.1 | 0.20 | 0.38 | 64 | 39 | 3.1 | 0.09 | 0.08 | 1144 | 73 | 0.03 |
| AU04A035 | RG45.16 | 1035 | 0.30 | 0.20 | 2.1 | 4.5 | 0.86 | 357 | 28 | 1.1 | 3.7 | 7250 | 34 | 3.8 | 0.24 | 2.9 | 3410 | 12212 | 0.12 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A273 | RG49.1 | 3.4 | 0.02 | 0.06 | 0.48 | 0.80 | 6.9 | 285 | 4.5 | 1.7 | 0.01 | 48 | 72 | 1.9 | 0.01 | 0.27 | 3.8 | 0.84 | 4.1 |
| AU04A274 | RG49.2 | 2.3 | 0.02 | 0.09 | 1.2 | 0.31 | 0.74 | 247 | 0.47 | 1.0 | 0.03 | 10 | 67 | 2.6 | 0.01 | 0.06 | 10 | 1.6 | 0.29 |
| AU04A275 | RG49.3 | 1555 | 0.48 | 0.09 | 2.3 | 3.3 | 3.8 | 80 | 2.9 | 6.8 | 0.23 | 27 | 147 | 7.6 | 0.01 | 1.4 | 155 | 62 | 5.0 |
| AU04A276 | RG49.4 | 2.7 | 0.03 | 0.00 | 0.78 | 0.67 | 0.35 | 211 | 0.36 | 0.93 | 0.04 | 11 | 177 | 2.5 | 0.00 | 0.12 | 29 | 6.9 | 0.09 |
| AU04A277 | RG49.5 | 4044 | 0.61 | 0.14 | 5.1 | 9.8 | 4.0 | 242 | 2.1 | 4.6 | 0.25 | 31 | 73 | 1.3 | 0.01 | 0.84 | 68 | 30 | 6.7 |
| AU04A278 | RG49.6 | 3.3 | 0.06 | 0.09 | 0.53 | 1.4 | 2.5 | 173 | 0.10 | 6.1 | 0.00 | 19 | 163 | 3.7 | 0.00 | 0.00 | 2.8 | 1.4 | 0.00 |
| AU04A279 | RG49.7 | 56 | 0.77 | 0.00 | 1.5 | 0.88 | 6.4 | 91 | 7.9 | 25 | 0.49 | 48 | 196 | 4.2 | 0.04 | 3.9 | 700 | 152 | 9.3 |
| AU04A280 | RG49.8 | 15 | 0.84 | 0.14 | 1.7 | 1.0 | 10 | 317 | 4.6 | 0.35 | 0.17 | 57 | 56 | 1.6 | 0.02 | 0.32 | 983 | 62 | 0.02 |


|  | AU04A281 | RG49.9 | 2557 | 1.4 | 0.17 | 6.1 | 5.3 | 4.6 | 113 | 5.0 | 11 | 0.53 | 38 | 235 | 4.5 | 0.02 | 2.7 | 810 | 102 | 11 |
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|  | AU04A282 | RG49.10 | 301 | 0.09 | 0.00 | 1.5 | 1.3 | 3.6 | 290 | 7.0 | 4.5 | 0.11 | 68 | 50 | 2.1 | 0.01 | 1.4 | 45 | 90 | 1.9 |
|  | FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
|  | AU04A147 | RG51.1 | 3953 | 1.1 | 0.09 | 1.7 | 7.6 | 0.59 | 106 | 23 | 1.1 | 0.07 | 8.6 | 301 | 4.9 | 0.14 | 0.30 | 7.9 | 1.8 | 2.7 |
|  | AU04A148 | RG51.2 | 5856 | 4.3 | 0.11 | 5.8 | 15 | 11 | 296 | 6.2 | 18 | 4.4 | 70 | 85 | 5.4 | 0.00 | 5.4 | 414 | 84 | 7.2 |
|  | AU04A149 | RG51.3 | 3583 | 2.2 | 0.06 | 6.7 | 10 | 5.9 | 153 | 8.6 | 9.6 | 1.7 | 36 | 146 | 14 | 0.01 | 4.2 | 342 | 66 | 7.5 |
|  | AU04A150 | RG51.4 | 789 | 1.7 | 0.09 | 8.7 | 4.0 | 9.0 | 319 | 2.3 | 2.4 | 1.8 | 129 | 60 | 3.8 | 0.00 | 1.1 | 612 | 139 | 0.43 |
|  | AU04A151 | RG51.5 | 547 | 0.18 | 0.00 | 1.3 | 1.0 | 0.65 | 59 | 0.39 | 0.50 | 0.13 | 22 | 354 | 11 | 0.00 | 0.13 | 38 | 20 | 1.2 |
|  | AU04A152 | RG51.6 | 1.3 | 0.01 | 0.32 | 0.79 | 2.2 | 3.0 | 333 | 0.03 | 13 | 0.00 | 46 | 53 | 6.3 | 0.00 | 0.00 | 0.33 | 0.10 | 0.01 |
|  | AU04A153 | RG51.7 | 6.7 | 0.01 | 0.00 | 0.57 | 0.87 | 8.8 | 320 | 0.03 | 0.56 | 0.01 | 37 | 68 | 6.1 | 0.00 | 0.00 | 0.19 | 0.02 | 0.00 |
|  | AU04A154 | RG51.8 | 28 | 0.02 | 0.10 | 1.8 | 1.2 | 12 | 331 | 0.05 | 0.29 | 0.03 | 42 | 73 | 2.9 | 0.01 | 0.01 | 0.43 | 0.05 | 0.01 |
|  | AU04A155 | RG51.9 | 104 | 0.03 | 0.11 | 7.5 | 2.5 | 12 | 340 | 0.04 | 2.9 | 0.01 | 43 | 38 | 3.4 | 0.00 | 0.01 | 0.78 | 0.25 | 0.02 |
|  | AU04A156 | RG51.10 | 690 | 0.28 | 0.12 | 4.0 | 3.4 | 0.58 | 138 | 0.09 | 0.78 | 0.31 | 8.0 | 199 | 6.5 | 0.01 | 0.09 | 9.4 | 5.2 | 0.25 |
|  | AU04A157 | RG51.11 | 137 | 0.07 | 0.06 | 9.2 | 2.2 | 8.4 | 296 | 0.06 | 1.5 | 0.02 | 29 | 41 | 0.92 | 0.02 | 0.01 | 1.0 | 0.35 | 0.41 |
|  | AU04A158 | RG51.12 | 6743 | 2.5 | 0.12 | 2.9 | 16 | 6.2 | 294 | 11 | 6.2 | 1.8 | 66 | 66 | 5.6 | 0.04 | 2.1 | 322 | 72 | 3.4 |
| WRD10600004 | FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
|  | AU04A173 | RG55.1 | 1408 | 1.4 | 2.9 | 291 | 13 | 2.2 | 562 | 0.83 | 1.4 | 0.26 | 41 | 213 | 4.0 | 5.0 | 0.23 | 51 | 81 | 0.36 |
|  | AU04A174 | RG55.2 | 2353 | 1.1 | 4.0 | 2473 | 14 | 7.7 | 498 | 0.66 | 5.6 | 0.22 | 32 | 276 | 0.70 | 6.8 | 0.09 | 55 | 74 | 0.11 |
|  | AU04A175 | RG55.3 | 2213 | 2.4 | 6.8 | 1329 | 26 | 5.6 | 517 | 5.0 | 6.2 | 1.0 | 95 | 276 | 1.2 | 24 | 0.33 | 138 | 306 | 0.44 |
|  | AU04A176 | RG55.4 | 993 | 1.7 | 4.4 | 714 | 9.3 | 4.2 | 483 | 3.1 | 3.3 | 0.40 | 73 | 124 | 2.6 | 7.0 | 0.40 | 85 | 157 | 0.44 |
|  | AU04A177 | RG55.5 | 1113 | 0.6 | 2.1 | 237 | 13 | 1.1 | 433 | 0.54 | 1.6 | 0.23 | 34 | 64 | 1.8 | 2.1 | 0.20 | 26 | 27 | 0.15 |
|  | AU04A178 | RG55.6 | 3499 | 4.6 | 17 | 317 | 30 | 8.1 | 593 | 7.5 | 6.1 | 4.4 | 119 | 418 | 0.52 | 21 | 1.1 | 160 | 403 | 1.2 |
|  | AU04A179 | RG55.7 | 1686 | 3.4 | 9.4 | 1349 | 14 | 6.5 | 518 | 5.1 | 4.0 | 1.8 | 209 | 210 | 0.61 | 75 | 0.57 | 256 | 741 | 0.13 |
|  | AU04A180 | RG55.8 | 8445 | 6.6 | 22 | 363 | 86 | 8.8 | 478 | 5.4 | 12 | 1.1 | 212 | 476 | 0.66 | 29 | 0.72 | 287 | 573 | 2.1 |
|  | AU04A181 | RG55.9 | 914 | 1.7 | 5.5 | 878 | 18 | 7.0 | 512 | 10 | 1.1 | 1.2 | 205 | 247 | 0.79 | 148 | 0.16 | 154 | 803 | 0.01 |
|  | AU04A182 | RG55.10 | 2383 | 3.4 | 9.2 | 416 | 19 | 6.1 | 419 | 14 | 5.5 | 1.6 | 212 | 135 | 0.36 | 92 | 0.67 | 292 | 725 | 0.50 |
|  | FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
|  | AU04A259 | RG56.1 | 1929 | 2.9 | 5.6 | 79 | 8.4 | 24 | 285 | 14 | 26 | 0.35 | 128 | 85 | 0.04 | 36 | 0.11 | 179 | 328 | 0.81 |
|  | AU04A260 | RG56.2 | 3722 | 7.5 | 13 | 308 | 19 | 5.0 | 309 | 15 | 8.0 | 1.2 | 132 | 155 | 0.35 | 25 | 0.48 | 441 | 343 | 3.4 |
|  | AU04A261 | RG56.3 | 8883 | 8.2 | 23 | 322 | 27 | 7.9 | 268 | 104 | 24 | 0.50 | 159 | 131 | 0.24 | 102 | 0.70 | 355 | 405 | 6.9 |
|  | AU04A262 | RG56.4 | 15638 | 11 | 37 | 320 | 62 | 3.9 | 281 | 22 | 16 | 0.70 | 115 | 140 | 0.17 | 27 | 1.1 | 246 | 236 | 8.8 |
|  | AU04A263 | RG56.5 | 190 | 0.15 | 0.50 | 15 | 1.2 | 0.36 | 269 | 0.17 | 0.16 | 0.01 | 48 | 41 | 0.44 | 0.6 | 0.02 | 7.5 | 12 | 0.10 |
|  | AU04A264 | RG56.6 | 147 | 0.56 | 0.69 | 45 | 3.3 | 5.4 | 271 | 3.4 | 0.53 | 0.14 | 53 | 55 | 0.34 | 7.1 | 0.03 | 43 | 46 | 0.01 |


| AU04A265 | RG56.7 | 2205 | 3.3 | 9.4 | 270 | 20 | 6.7 | 306 | 3.9 | 69 | 0.35 | 73 | 190 | 0.25 | 1.2 | 0.12 | 277 | 102 | 0.78 |
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| AU04A266 | RG56.8 | 2556 | 4.8 | 6.9 | 524 | 11 | 9.4 | 280 | 5.3 | 2.8 | 0.45 | 179 | 173 | 0.45 | 34 | 1.2 | 360 | 584 | 0.14 |
| AU04A267 | RG56.9 | 10370 | 7.6 | 18 | 448 | 32 | 16 | 333 | 5.4 | 7.9 | 0.44 | 137 | 347 | 0.18 | 24 | 1.7 | 255 | 415 | 0.45 |
| AU04A268 | RG56.10 | 20770 | 9.3 | 27 | 106 | 39 | 2.6 | 284 | 2.2 | 3.1 | 0.60 | 93 | 145 | 2.2 | 17 | 2.5 | 493 | 347 | 0.35 |
| AU04A269 | RG56.11 | 2556 | 20 | 17 | 358 | 11 | 4.5 | 388 | 4.7 | 5.9 | 0.99 | 175 | 260 | 3.3 | 33 | 2.5 | 577 | 831 | 0.10 |
| AU04A270 | RG56.12 | 12620 | 4.9 | 16 | 184 | 37 | 1.2 | 325 | 1.2 | 1.0 | 0.35 | 77 | 110 | 2.5 | 13 | 2.1 | 177 | 198 | 0.04 |
| AU04A271 | RG56.13 | 3968 | 12 | 13 | 130 | 13 | 6.8 | 409 | 125 | 9.0 | 0.67 | 184 | 410 | 0.88 | 435 | 1.8 | 460 | 825 | 0.44 |
| AU04A272 | RG56.14 | 22375 | 11 | 31 | 70 | 41 | 8.2 | 342 | 4.9 | 2.8 | 0.65 | 101 | 298 | 1.2 | 34 | 2.5 | 338 | 409 | 0.37 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A199 | RG60.1 | 795 | 6.4 | 6.5 | 148 | 13 | 5.6 | 337 | 16 | 4.6 | 0.3 | 58 | 115 | 0.40 | 13 | 0.12 | 107 | 129 | 0.06 |
| AU04A200 | RG60.2 | 340 | 1.4 | 2.0 | 86 | 6.2 | 2.8 | 365 | 33 | 4.7 | 0.6 | 116 | 65 | 0.46 | 55 | 0.06 | 319 | 440 | 0.05 |
| AU04A201 | RG60.3 | 10401 | 14 | 13 | 286 | 40 | 12 | 433 | 146 | 13 | 1.4 | 141 | 159 | 0.57 | 122 | 0.99 | 610 | 655 | 0.93 |
| AU04A202 | RG60.4 | 574 | 5.9 | 9.3 | 329 | 21 | 10 | 386 | 160 | 1.2 | 0.40 | 148 | 99 | 0.42 | 17 | 0.19 | 311 | 356 | 0.05 |
| AU04A203 | RG60.5 | 178 | 2.1 | 3.7 | 161 | 3.8 | 2.9 | 317 | 138 | 0.88 | 0.25 | 168 | 99 | 0.15 | 13 | 0.26 | 92 | 457 | 0.02 |
| AU04A204 | RG60.6 | 748 | 26 | 31 | 375 | 47 | 6.1 | 346 | 33 | 1.1 | 0.18 | 59 | 188 | 0.35 | 1.6 | 0.11 | 111 | 99 | 0.34 |
| AU04A205 | RG60.7 | 678 | 5.4 | 7.4 | 206 | 15 | 5.0 | 438 | 41 | 1.2 | 0.37 | 100 | 98 | 5.2 | 17 | 0.18 | 227 | 268 | 0.20 |
| AU04A206 | RG60.8 | 557 | 2.3 | 7.5 | 152 | 2.3 | 4.5 | 354 | 91 | 0.34 | 0.16 | 95 | 65 | 0.67 | 67 | 0.03 | 54 | 178 | 0.00 |
| AU04A207 | RG60.9 | 57 | 3.6 | 4.0 | 34 | 5.8 | 2.6 | 469 | 40 | 0.14 | 0.16 | 50 | 207 | 1.8 | 32 | 0.04 | 72 | 167 | 0.05 |
| AU04A208 | RG60.10 | 1541 | 14 | 21 | 760 | 49 | 5.3 | 441 | 77 | 2.6 | 0.40 | 89 | 84 | 4.5 | 18 | 0.25 | 189 | 238 | 0.31 |
| AU04A209 | RG60.11 | 572 | 83 | 125 | 537 | 125 | 3.4 | 454 | 326 | 2.6 | 0.76 | 182 | 116 | 5.9 | 83 | 0.36 | 498 | 586 | 0.79 |
| AU04A210 | RG60.12 | 243 | 4.0 | 4.4 | 130 | 6.3 | 2.9 | 458 | 47 | 1.4 | 0.38 | 84 | 180 | 1.6 | 9.3 | 0.24 | 167 | 295 | 0.16 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
| AU04A003 | RG62.1 | 575 | 0.51 | 0.76 | 414 | 5.3 | 1.7 | 535 | 0.33 | 0.38 | 0.11 | 43 | 92 | 1.6 | 3.4 | 0.08 | 30 | 53 | 0.01 |
| AU04A004 | RG62.2 | 635 | 1.5 | 2.1 | 257 | 9.5 | 1.6 | 414 | 4.9 | 0.66 | 0.26 | 92 | 93 | 0.12 | 2.0 | 0.37 | 33 | 257 | 0.08 |
| AU04A005 | RG62.3 | 1029 | 0.55 | 0.92 | 706 | 16 | 1.4 | 410 | 0.68 | 0.45 | 0.09 | 54 | 58 | 1.7 | 2.7 | 0.60 | 18 | 47 | 0.00 |
| AU04A006 | RG62.4 | 1181 | 2.4 | 3.9 | 90 | 15 | 3.0 | 485 | 1.4 | 0.30 | 0.25 | 144 | 197 | 1.9 | 8.9 | 0.32 | 50 | 326 | 0.04 |
| AU04A007 | RG62.5 | 1904 | 11 | 25 | 51 | 85 | 3.9 | 471 | 5.5 | 1.3 | 0.35 | 134 | 221 | 2.0 | 22 | 0.85 | 102 | 371 | 0.23 |
| AU04A008 | RG62.6 | 96 | 0.58 | 0.90 | 23 | 0.18 | 0.28 | 447 | 0.14 | 0.07 | 0.02 | 48 | 50 | 0.15 | 0.55 | 0.04 | 1.3 | 12 | 0.00 |
| AU04A009 | RG62.7 | 9465 | 5.0 | 10 | 29 | 149 | 2.5 | 420 | 5.2 | 1.4 | 0.39 | 179 | 161 | 1.7 | 15 | 1.6 | 106 | 225 | 0.34 |
| AU04A010 | RG62.8 | 3212 | 2.2 | 4.1 | 39 | 67 | 3.9 | 403 | 0.58 | 0.50 | 0.15 | 543 | 82 | 0.21 | 6.6 | 0.58 | 42 | 999 | 0.05 |
| AU04A011 | RG62.9 | 10591 | 4.2 | 9.5 | 15 | 91 | 3.4 | 347 | 0.90 | 1.2 | 0.20 | 75 | 116 | 0.17 | 2.4 | 1.3 | 38 | 81 | 0.08 |
| AU04A012 | RG62.10 | 4205 | 2.9 | 5.7 | 20 | 82 | 4.6 | 540 | 2.3 | 1.1 | 0.21 | 158 | 220 | 0.31 | 11 | 1.1 | 62 | 213 | 0.52 |
| AU04A013 | RG62.11 | 101 | 0.53 | 0.43 | 26 | 5.2 | 18 | 442 | 1.3 | 0.20 | 0.08 | 340 | 96 | 0.06 | 21 | 0.04 | 10 | 521 | 0.03 |
| AU04A014 | RG62.12 | 13456 | 7.4 | 14 | 50 | 270 | 4.8 | 433 | 5.4 | 1.9 | 0.38 | 222 | 269 | 1.3 | 26 | 1.3 | 234 | 274 | 0.48 |


| AU04A015 | RG62.13 | 979 | 3.0 | 4.4 | 68 | 84 | 7.5 | 387 | 3.6 | 1.0 | 0.37 | 158 | 188 | 0.40 | 8.9 | 0.88 | 98 | 207 | 0.16 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| AU04A016 | RG62.14 | 42 | 0.81 | 1.3 | 32 | 1.7 | 7.9 | 362 | 0.29 | 0.49 | 0.08 | 142 | 82 | 0.05 | 1.1 | 0.02 | 21 | 143 | 0.00 |
| AU04A017 | RG62.15 | 40 | 0.75 | 0.84 | 30 | 3.2 | 6.0 | 342 | 1.0 | 0.25 | 0.07 | 203 | 63 | 0.31 | 9.3 | 0.06 | 8.4 | 240 | 0.00 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | TI205 | Pb208 | Bi209 | U238 |
| AU04A057 | RG64.1 | 12365 | 5.7 | 8.7 | 220 | 41 | 3.9 | 475 | 8.6 | 0.75 | 0.93 | 79 | 108 | 16 | 14 | 3.1 | 160 | 159 | 0.11 |
| AU04A058 | RG64.2 | 2609 | 8.6 | 6.3 | 94 | 7.8 | 10 | 452 | 24 | 1.2 | 3.1 | 197 | 105 | 0.46 | 20 | 4.3 | 273 | 475 | 0.09 |
| AU04A059 | RG64.3 | 1495 | 11 | 7.0 | 469 | 5.0 | 5.5 | 463 | 38 | 1.4 | 2.1 | 250 | 124 | 1.1 | 86 | 2.9 | 424 | 629 | 0.01 |
| AU04A060 | RG64.4 | 1557 | 5.6 | 4.8 | 104 | 5.9 | 3.7 | 404 | 46 | 0.2 | 0.9 | 216 | 92 | 0.49 | 111 | 0.92 | 387 | 724 | 0.01 |
| AU04A061 | RG64.5 | 6556 | 44 | 30 | 1018 | 20 | 28 | 455 | 31 | 9.4 | 11.4 | 256 | 247 | 1.3 | 14 | 14 | 1491 | 997 | 2.3 |
| AU04A062 | RG64.6 | 3506 | 7.1 | 4.5 | 75 | 10 | 16 | 491 | 42 | 1.3 | 1.9 | 230 | 84 | 0.25 | 64 | 1.2 | 218 | 448 | 0.11 |
| AU04A063 | RG64.7 | 31885 | 10 | 17 | 23 | 85 | 11 | 502 | 1.2 | 1.7 | 0.39 | 48 | 396 | 110 | 1.4 | 2.4 | 44 | 15 | 0.96 |
| AU04A064 | RG64.8 | 21739 | 7.5 | 13 | 16 | 61 | 8.2 | 385 | 3.0 | 1.2 | 0.30 | 127 | 430 | 218 | 3.9 | 2.0 | 138 | 340 | 2.0 |
| AU04A065 | RG64.9 | 3857 | 2.7 | 4.1 | 23 | 14 | 5.6 | 594 | 386 | 0.54 | 0.34 | 153 | 351 | 135 | 517 | 0.43 | 165 | 379 | 1.5 |
| AU04A066 | RG64.10 | 5307 | 1.8 | 3.2 | 110 | 15 | 10 | 493 | 0.71 | 6.8 | 0.10 | 49 | 359 | 191 | 0.58 | 0.55 | 9.0 | 54 | 0.18 |
| AU04A067 | RG64.11 | 12458 | 5.2 | 7.5 | 158 | 36 | 3.9 | 415 | 32 | 0.88 | 1.1 | 145 | 91 | 1.2 | 18 | 3.6 | 149 | 336 | 0.02 |
| AU04A068 | RG64.12 | 1201 | 2.7 | 2.3 | 118 | 4.3 | 2.8 | 526 | 18 | 0.61 | 1.1 | 130 | 67 | 3.0 | 31 | 0.44 | 207 | 222 | 0.02 |
| AU04A069 | RG64.13 | 5825 | 8.4 | 8.2 | 343 | 14 | 14 | 475 | 81 | 4.6 | 2.2 | 314 | 602 | 1.8 | 122 | 2.9 | 603 | 1146 | 3.3 |
| AU04A070 | RG64.14 | 13372 | 50 | 46 | 344 | 15 | 31 | 566 | 114 | 22 | 6.7 | 278 | 459 | 2.5 | 75 | 18 | 1245 | 1356 | 39 |
| AU04A071 | RG64.15 | 9027 | 38 | 28 | 394 | 26 | 28 | 501 | 285 | 3.6 | 10 | 246 | 567 | 3.3 | 402 | 17 | 835 | 926 | 0.64 |
| AU04A072 | RG64.16 | 4335 | 12 | 8.4 | 88 | 14 | 11 | 518 | 13 | 1.7 | 3.8 | 136 | 159 | 0.66 | 10 | 3.1 | 373 | 300 | 0.24 |
| AU04A073 | RG64.17 | 33273 | 11 | 18 | 103 | 85 | 10 | 490 | 23 | 2.1 | 1.9 | 200 | 329 | 1.4 | 39 | 5.1 | 356 | 536 | 0.24 |
| AU04A074 | RG64.18 | 5110 | 10 | 7.4 | 3010 | 85 | 8.9 | 548 | 32 | 3.3 | 2.1 | 434 | 150 | 1.8 | 46 | 3.2 | 494 | 1287 | 0.56 |
| AU04A075 | RG64.19 | 2073 | 18 | 12 | 227 | 6.1 | 13 | 687 | 37 | 1.0 | 2.5 | 170 | 130 | 0.50 | 47 | 7.4 | 278 | 539 | 0.08 |
| FileName | SampleName | Fe57 | Co59 | Ni60 | Cu65 | Zn66 | As75 | Se77 | Ag107 | Sn118 | Sb121 | Te125 | W182 | Re185 | Au197 | T1205 | Pb208 | Bi209 | U238 |
| AU04A245 | RG65.1 | 5509 | 2.8 | 7.3 | 88 | 39 | 5.3 | 335 | 9.2 | 3.0 | 1.2 | 238 | 96 | 4.1 | 19 | 0.49 | 244 | 718 | 4.5 |
| AU04A246 | RG65.2 | 28 | 1.0 | 0.47 | 29 | 4.0 | 2.2 | 371 | 0.07 | 1.1 | 0.05 | 27 | 70 | 4.0 | 0.01 | 0.01 | 5.4 | 5.1 | 0.01 |
| AU04A247 | RG65.3 | 3029 | 0.9 | 2.9 | 38 | 21 | 5.1 | 314 | 7.4 | 3.6 | 0.31 | 71 | 95 | 1.0 | 2.5 | 0.50 | 72 | 158 | 1.6 |
| AU04A248 | RG65.4 | 9457 | 4.0 | 9.6 | 64 | 43 | 5.6 | 357 | 27 | 7.0 | 0.47 | 206 | 144 | 2.4 | 71 | 0.84 | 160 | 744 | 3.2 |
| AU04A249 | RG65.5 | 153 | 0.0 | 1.7 | 3.0 | 10 | 4.7 | 311 | 0.05 | 34 | 0.00 | 31 | 60 | 1.5 | 0.05 | 0.01 | 2.2 | 2.8 | 0.06 |
| AU04A250 | RG65.6 | 359 | 1.0 | 0.64 | 19 | 2.4 | 1.9 | 325 | 0.29 | 0.34 | 0.04 | 31 | 117 | 0.26 | 0.02 | 0.06 | 3.1 | 5.1 | 0.04 |
| AU04A251 | RG65.7 | 4566 | 2.6 | 3.3 | 240 | 132 | 3.9 | 310 | 26 | 28 | 1.3 | 457 | 104 | 0.66 | 15 | 0.67 | 151 | 2105 | 4.6 |
| AU04A252 | RG65.8 | 1843 | 2.3 | 4.7 | 114 | 31 | 13 | 387 | 464 | 9.2 | 9.9 | 117 | 88 | 1.6 | 273 | 1.4 | 78 | 1087 | 5.1 |
| AU04A253 | RG65.9 | 1057 | 2.1 | 3.5 | 206 | 58 | 6.9 | 373 | 98 | 56 | 1.4 | 189 | 83 | 0.97 | 55 | 1.5 | 114 | 781 | 0.64 |
| AU04A254 | RG65.10 | 2637 | 19 | 5.0 | 767 | 12 | 1.8 | 357 | 1.0 | 0.73 | 0.23 | 40 | 107 | 0.49 | 0.0 | 0.15 | 14 | 84 | 0.02 |


[^0]:    $x=E P M A$ silicate; $X=E P M A$ ore min.; *FIB-TEM

[^1]:    Detection limits (wt\%) Cu 0.06, Zn 0.04, Mn 0.03, Cd 0.09, In 0.12, Sn 0.11, S 0.06

